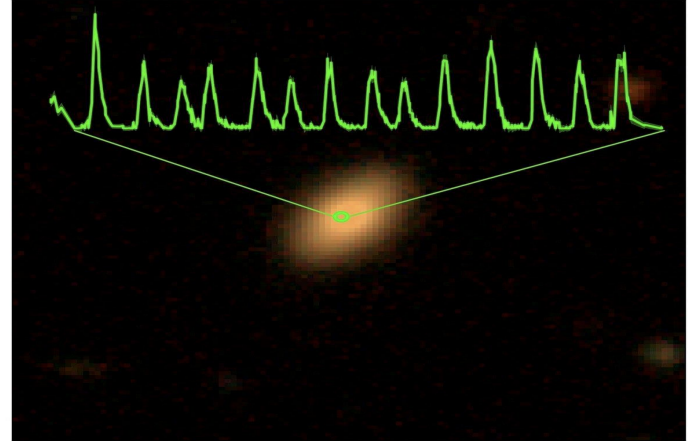
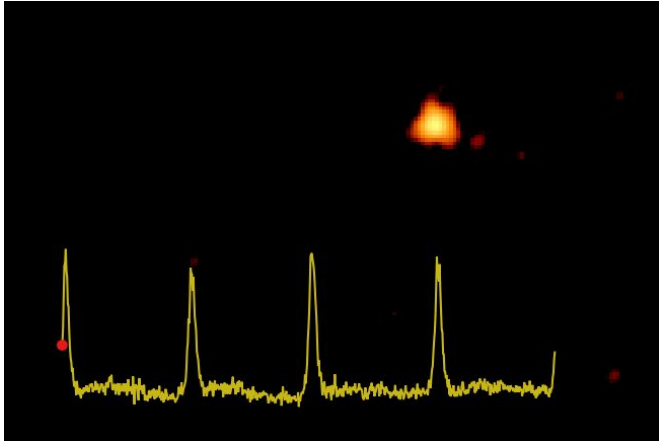


# Quasi-Periodic (X-ray) Eruptions from Galactic Nuclei



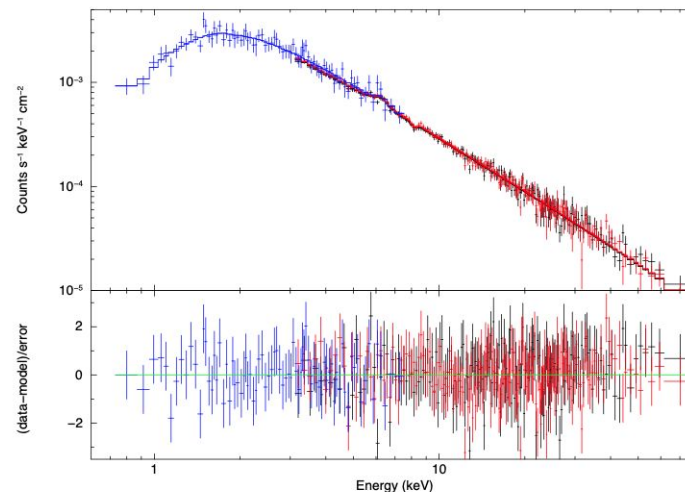
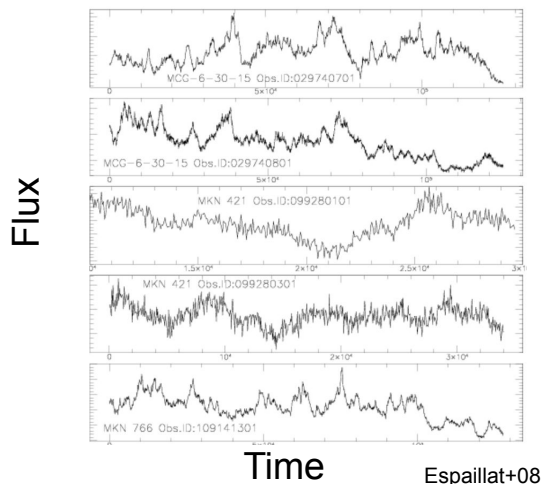
**Muryel Guolo**  
**Johns Hopkins University**

Dheeraj R. Pasham, Michal Zajacek, Eric Coughlin, Suvi Gezari, et al.

# Nuclear Activity: **AGN** or TDEs

## Typical Active Galactic Nucleus (AGN) in the X-ray band:

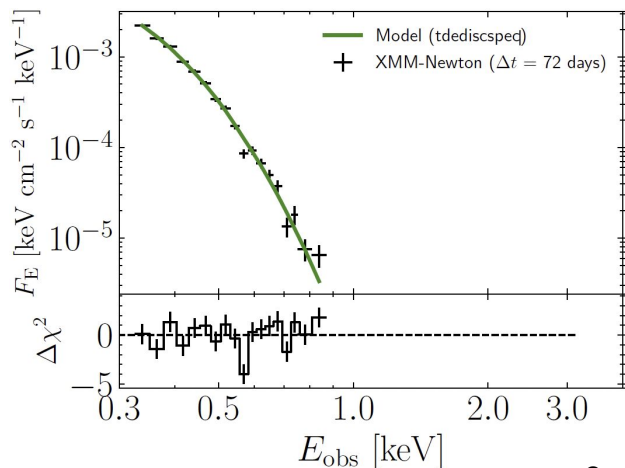
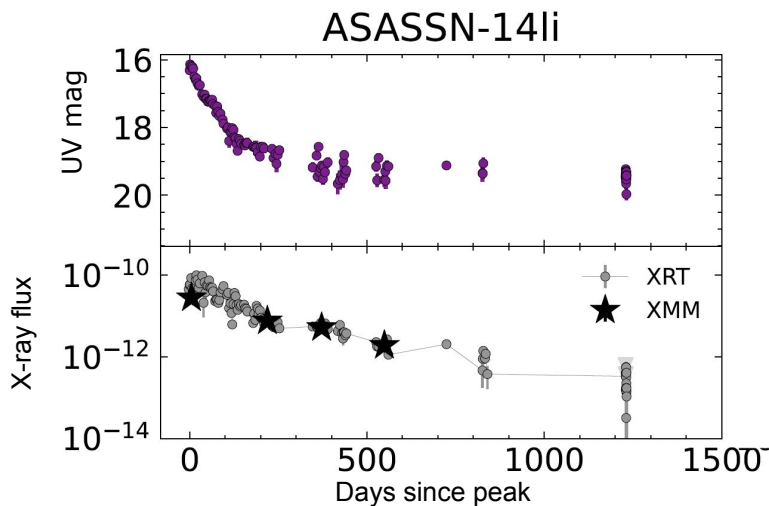
- 30-50% fractional variability; “aperiodic”, “random”, “stochastic”
- Power-law X-ray spectrum extending to hard X-rays



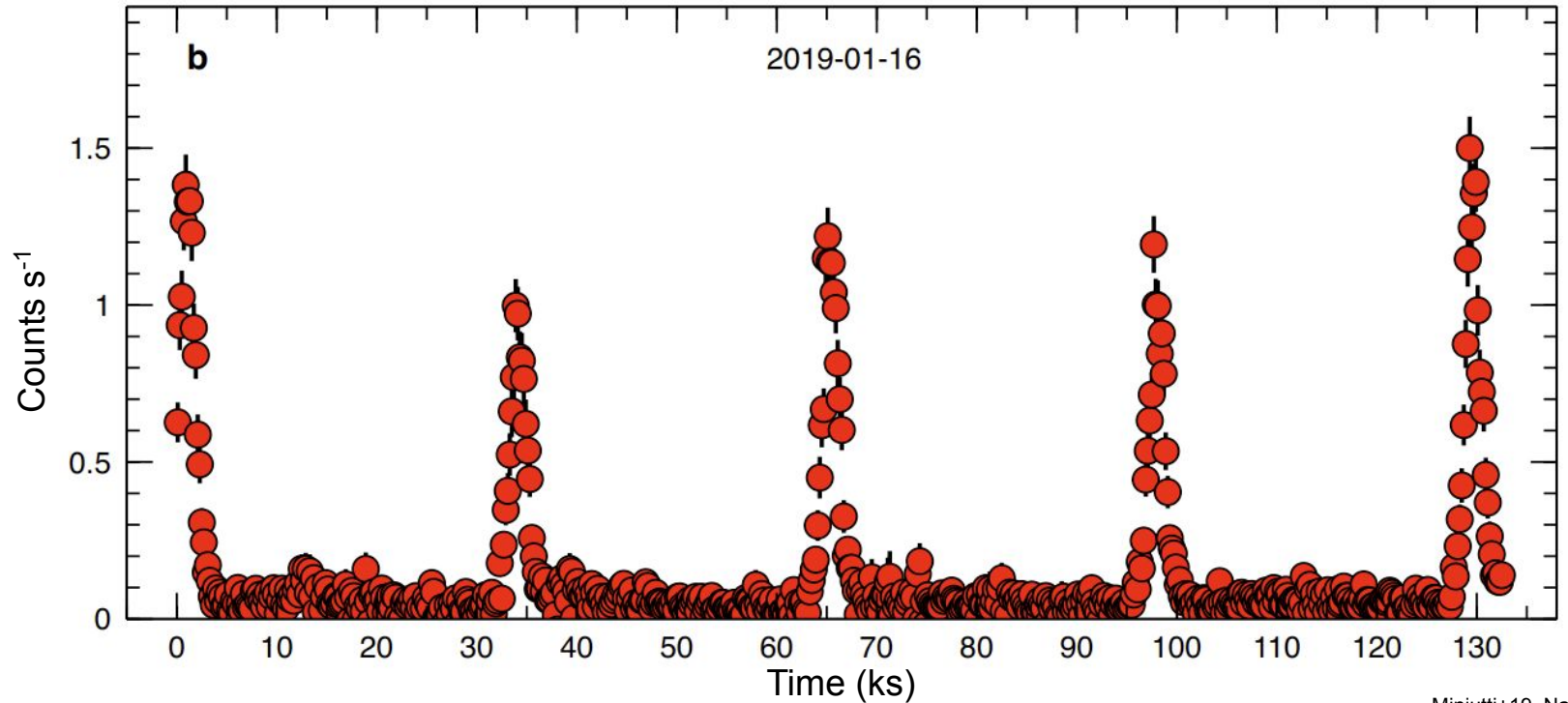
# Nuclear Activity: AGN or **TDEs**

## Typical Tidal Disruption Event (TDE) in the X-ray band:

- Few years-long flare from a previously quiescent galaxy ( $F_x \propto t^{-\alpha}$ ), usually UV bright.
- Soft Thermal X-ray spectrum ( $T \sim 50\text{-}200$  eV)

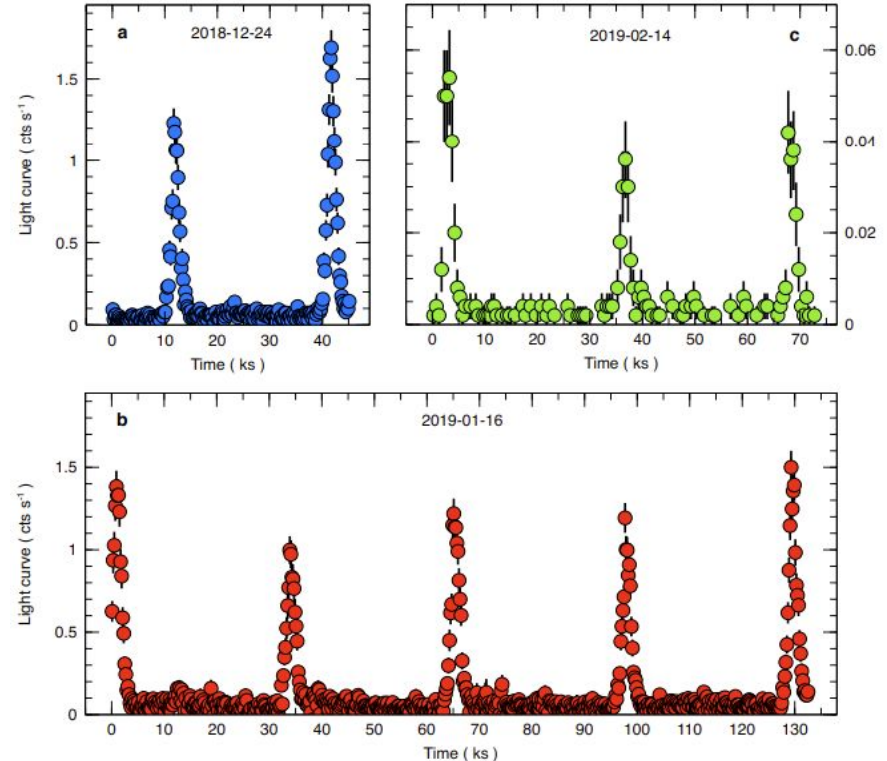


# 2019: X-ray Quasi-Periodic Eruptions (QPEs) discovered in GSN 069



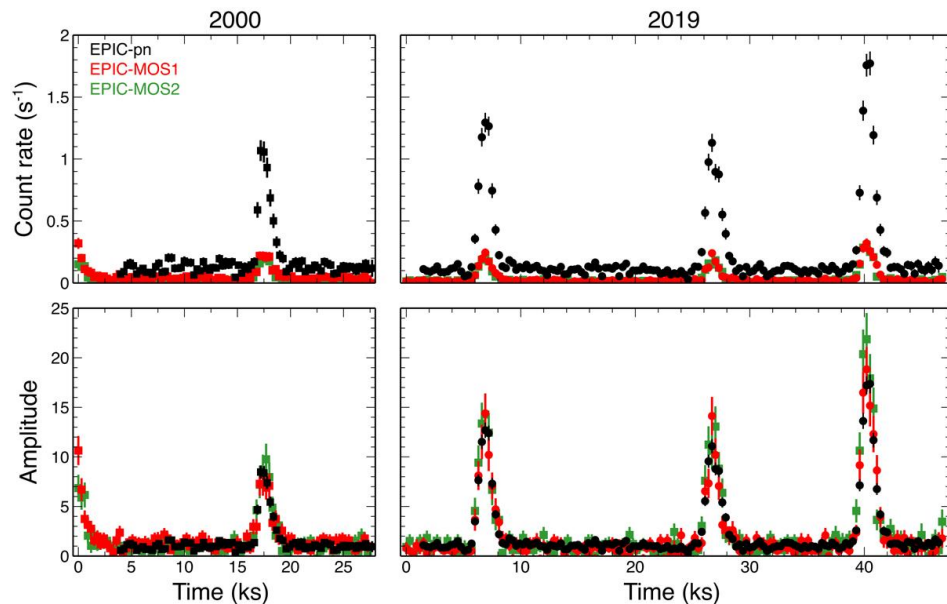
# Nuclear Activity: AGN or TDEs or **QPEs**

- 4 systems:
  - **GSN 069 (Miniutti+19)**
  - RX J1301 (Giustini+20)
  - eRO-QPE1 (Arcodia+21)
  - eRO-QPE2 (Arcodia+21)
- Recurrence times 2-20 hours
- Flare durations 0.5-8 hours
- Peak  $L_X = 10^{42}$ - $10^{43}$  erg/s
- Amplitude 10-200
- No UV/optical counterpart
- $M_{\text{BH}} = 10^5$ - $10^{6.6} M_{\odot}$



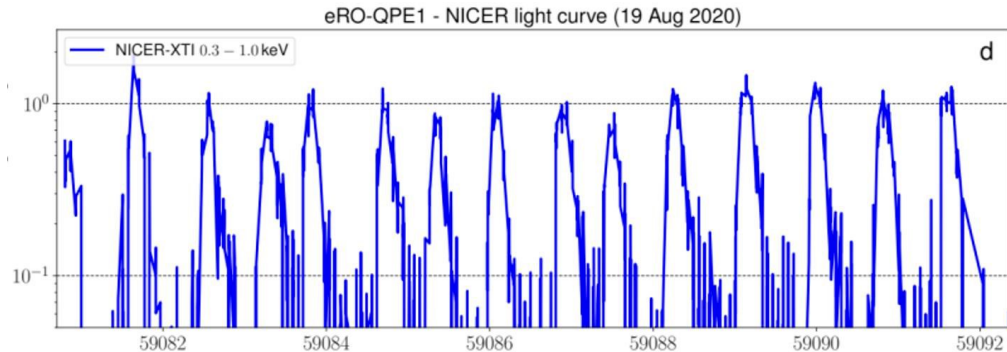
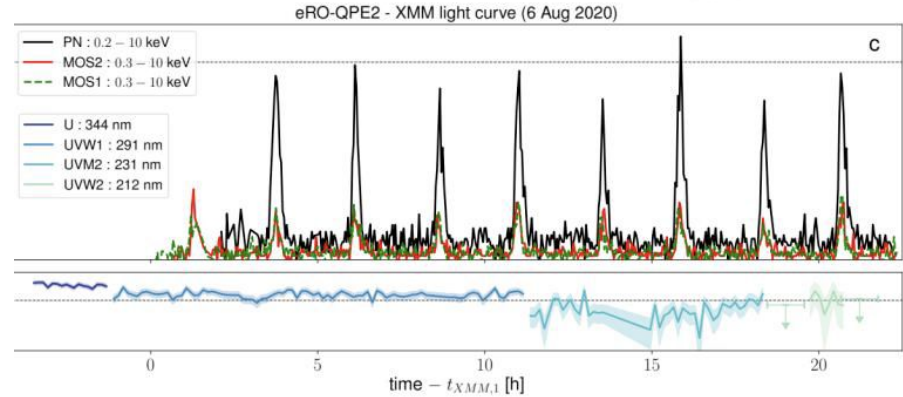
# Nuclear Activity: AGN or TDEs or **QPEs**

- 4 systems:
  - GSN 069 (Miniutti+19)
  - **RX J1301 (Giustini+20)**
  - eRO-QPE1 (Arcodia+21)
  - eRO-QPE2 (Arcodia+21)
- Recurrence times 2-20 hours
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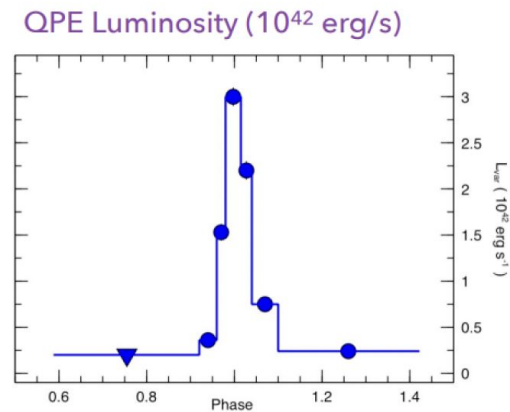
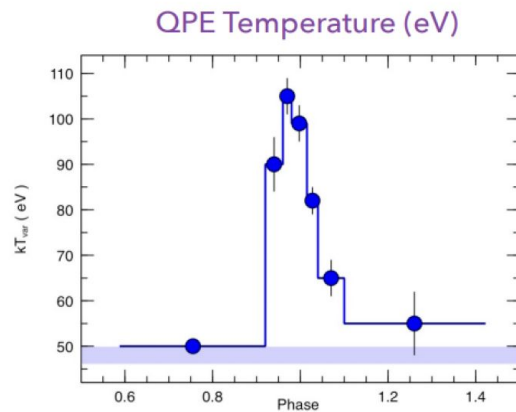
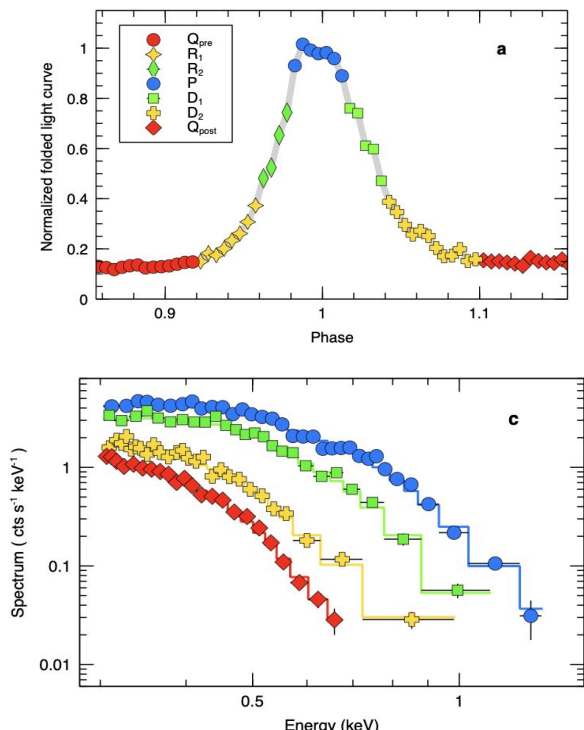
# Nuclear Activity: AGN or TDEs or **QPEs**

- 4 systems:
  - GSN 069 (Miniutti+19)
  - RX J1301 (Giustini+20)
  - **eRO-QPE1 (Arcodia+21)**
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- Amplitude 10-200
- No UV/optical counterpart
- $M_{\text{BH}} = 10^5 - 10^{6.6} M_{\odot}$



# QPEs Spectra

- Blackbody or multi-temperature disk (*diskbb* on Xspec),  $T \sim 50\text{-}200$  eV



Miniutti+19

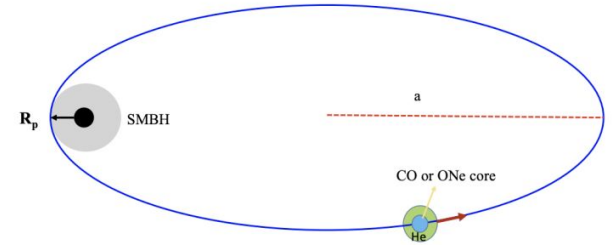


# Models for QPE sources

- ~~Self-lensing black hole binary~~
- ~~Disk instabilities~~
- **Orbiting bodies?**
  - **EM counterparts of extreme mass ratio inspirals (EMRIs)**

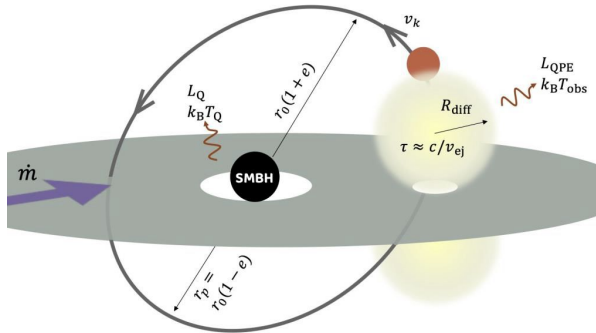
Repeating partial TDE of compact object or main-sequence companion(s)

(King 2020, 2022; Zhao et al. 2021; Metzger & Stone 2022)



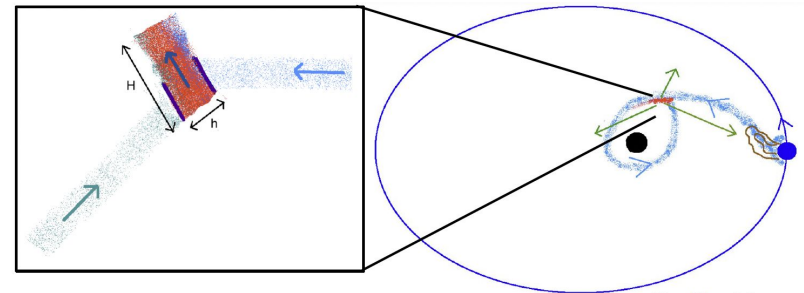
## Star-disk (or ADAF) collisions/interactions

(Xian et al. 2021, Sukova et al. 2021, Linal & Metzger 2023)



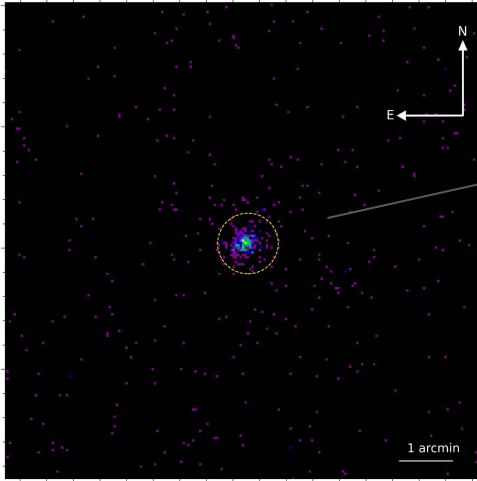
Linal & Metzger 2023

Shock-powered flares from a main-sequence companion (Lu & Quataert 2022; Linal & Sari 2022; Krolik & Linal 2022)

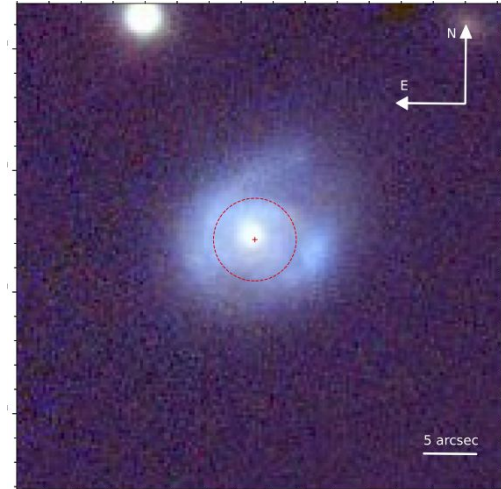


Krolik+22

# Crazy Swift Source



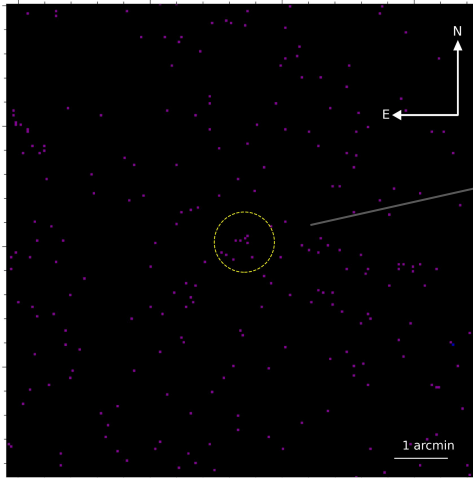
$$F_x \sim 1 \times 10^{-12} \text{ cgs}$$
$$L_x \sim 5 \times 10^{42} \text{ cgs}$$



- Serendipitously discovered by Swift/XRT
- 35x brighter than previous upper limit
- Soft X-ray spectra all counts below 1 keV
- X-ray TDE candidate
- However....

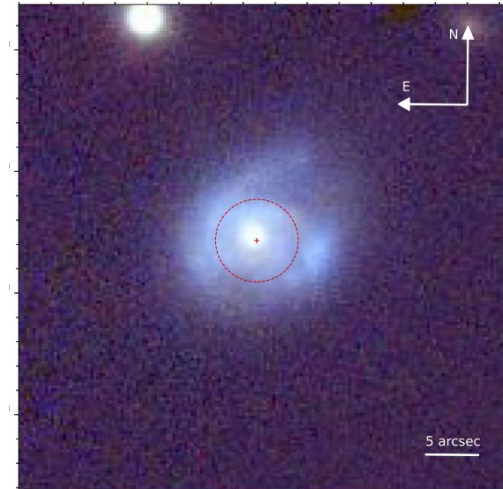
Host: spiral galaxy at 165 Mpc ( $z \sim 0.036$ )

# Crazy Swift Source



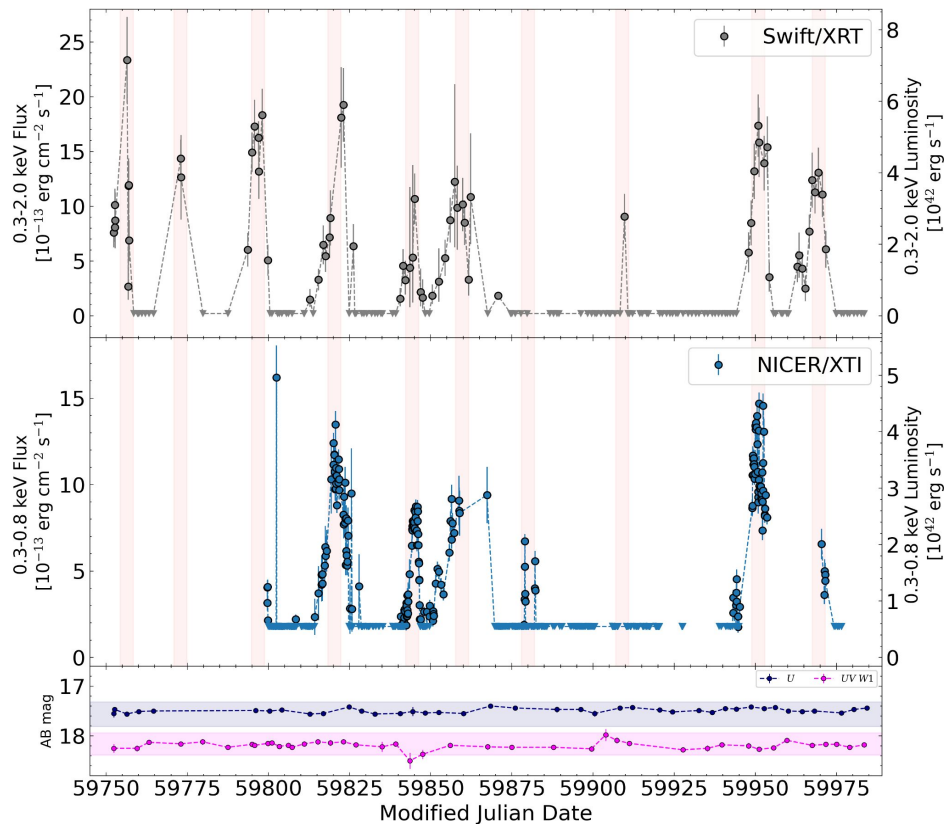
$F_x < 2 \times 10^{-14}$  cgs  
 $L_x < 1 \times 10^{41}$  cgs

- Serendipitously discovered by Swift/XRT
- 35x brighter than previous upper limit
- Soft X-ray spectra all counts below 1 keV
- X-ray TDE candidate
- However.... GONE TWO WEEKS LATER!



Host: spiral galaxy at 165 Mpc ( $z \sim 0.036$ )

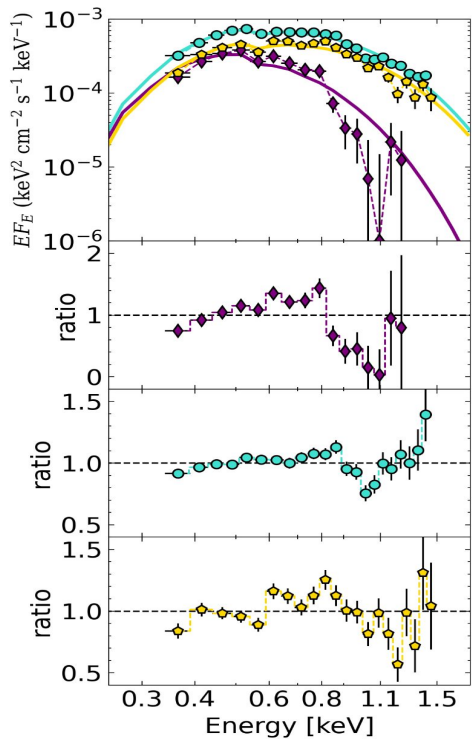
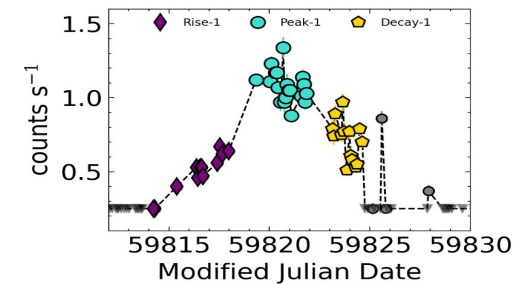
# Crazy Swift Source



- Few days long X-ray eruptions each  $\sim 20$  days, but gaps...
- No detections during quiescent  $F_x < 2 \times 10^{-14}$  cgs
- Amplitude  $> 100$
- No UV optical Counterpart
- No clear AGN in the host galaxy (from X-ray, Radio, optical nor IR), but BPT/WHAN indicates a low-luminosity AGN.
- $\log M_{\text{BH}} = 6.6 \pm 0.4 M_{\odot}$

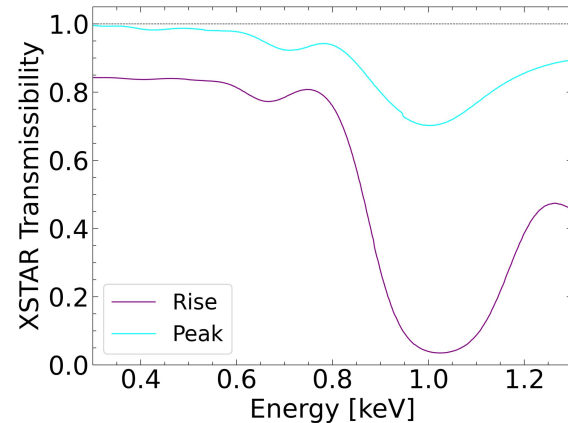
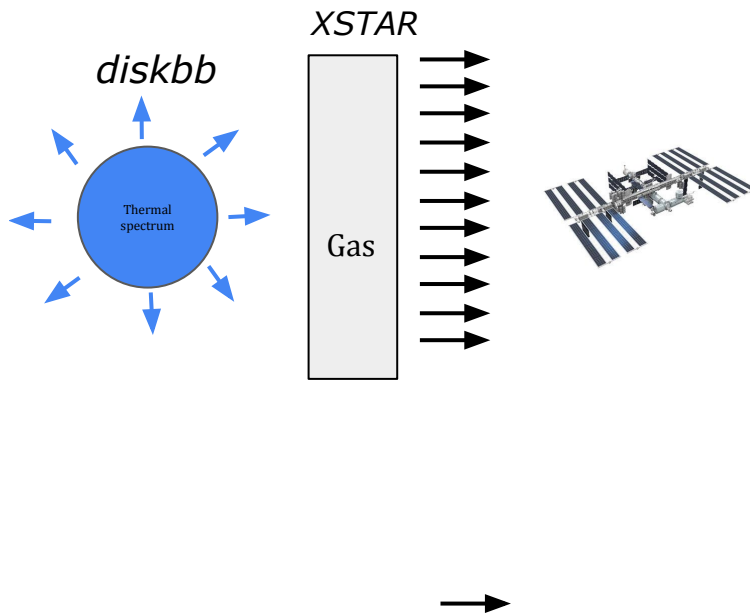
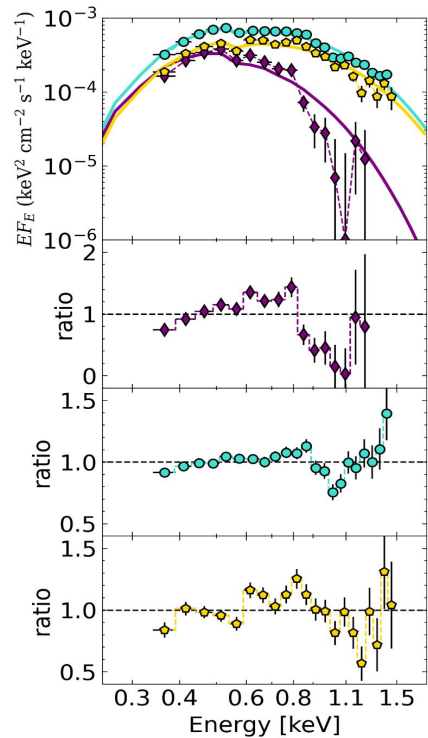
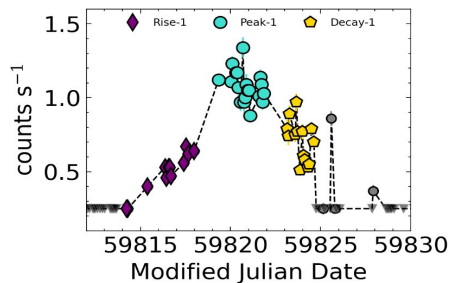
# Crazy Swift Source: Spectra

- Soft/Thermal X-ray spectrum  $T \sim 100\text{-}250\text{ eV}$
- Strong residuals 0.9-1.1 Kev



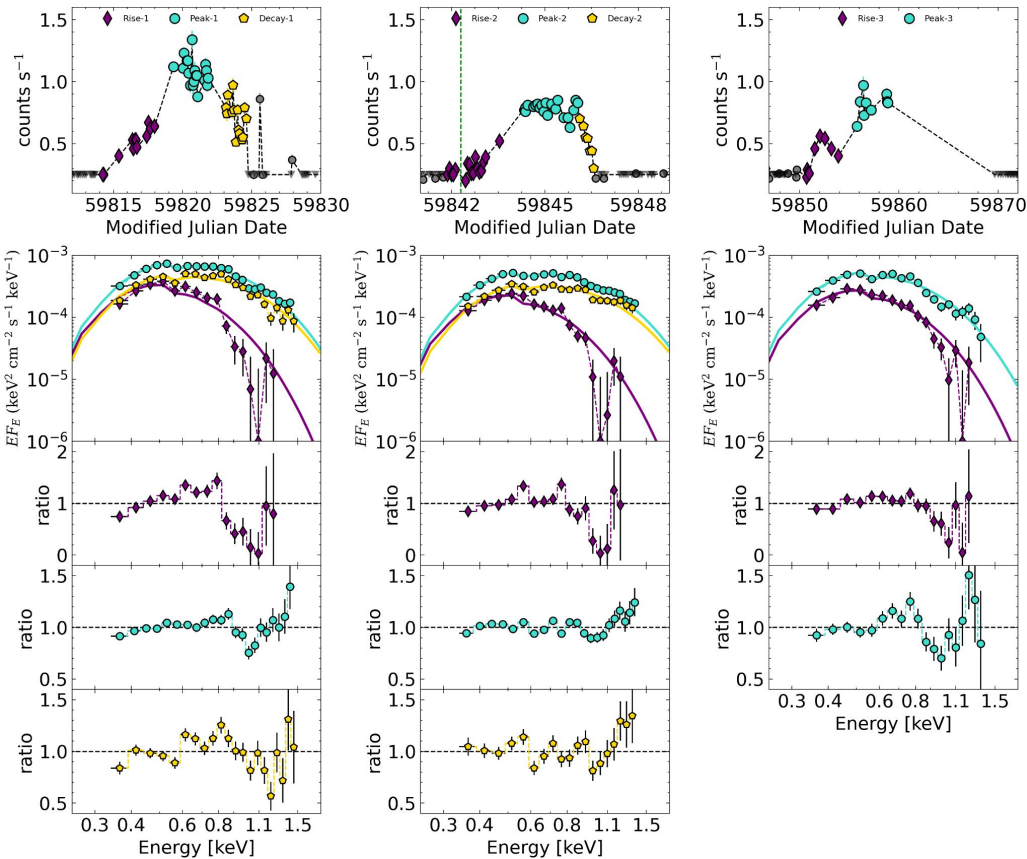
# Crazy Swift Source: Spectra

- Soft/Thermal X-ray spectrum  $T \sim 100\text{-}250\text{ eV}$
- Strong residuals 0.9-1.1 KeV  $\rightarrow$  **Ultra Fast Outflow (UFO),  $v \sim 0.1c$**

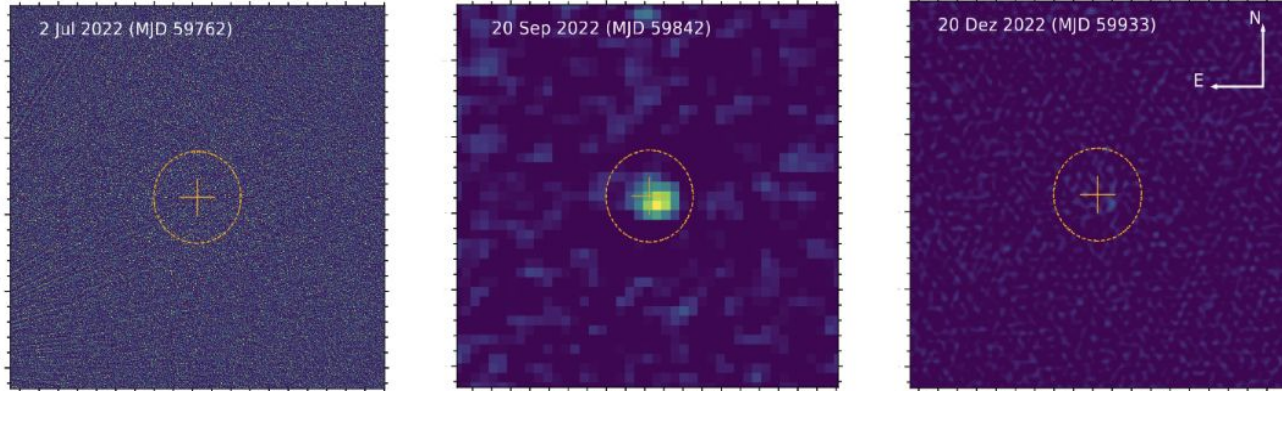


# Outflow X-rays

- **Soft/Thermal X-ray spectrum  $T \sim 100\text{-}250$  eV**
- **Ultra Fast Outflow (UFO)  $v \sim 0.1c$  during the rise of the eruptions**



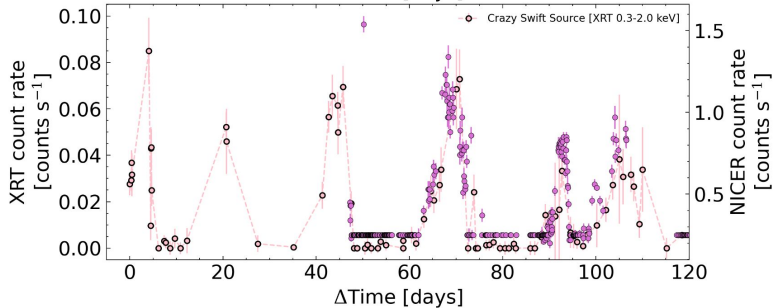
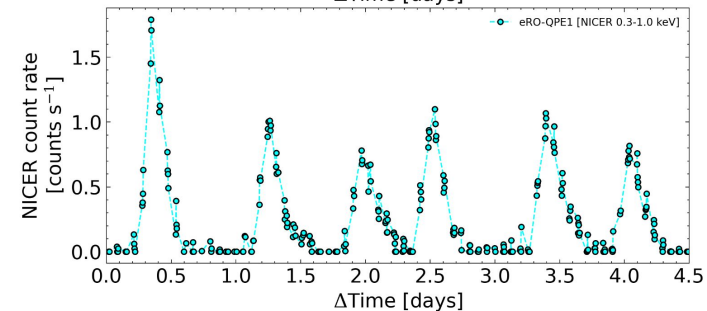
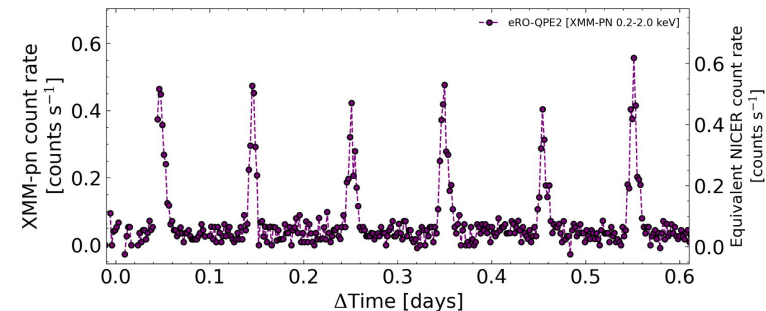
# Outflow: Radio



Transient radio (VLA) detection during one of the X-ray eruptions



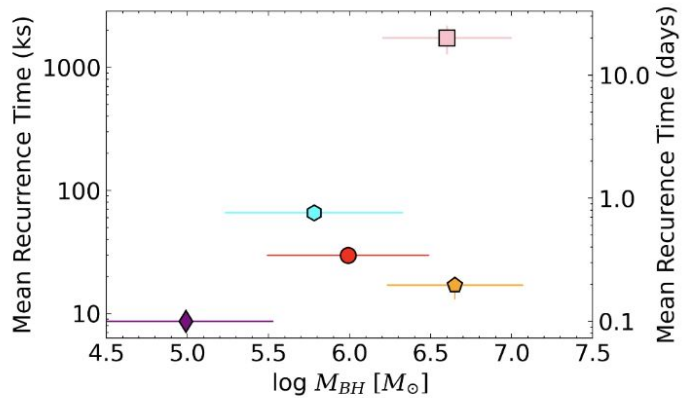
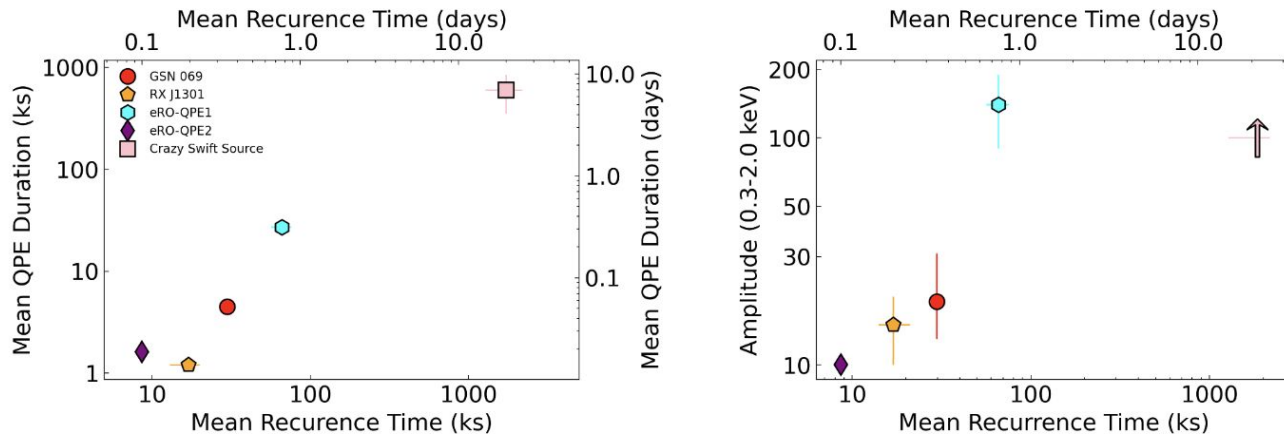
# Crazy Swift Source: as a super-long QPE



All characteristics of a QPE source, however...

- Recurrence time 25x longer than longest known QPE (eRO-QPE1 ~ 20 hours)

# Crazy Swift Source: as a super-long QPE



# Conclusions

- We discovered a new QPE candidate, the fifth and with longest recurrence time.
- It opens an unexplored time-scale for QPEs and nuclear transient
- First QPE with evidence for outflows.
- Theoretical understanding of QPEs is still incomplete, many possible scenarios, but should be scalable from 3 hours to 20 days.