

Searching for extragalactic transients with SPT-3G



Sam Guns

UC Berkeley

The Transient and
Variable Universe

June 22 2023

What is SPT-3G?

Third generation receiver on the 10 meter South Pole Telescope

Since 2018: running a multiyear survey focused on cosmology (CMB powerspectra & lensing), point sources, and clusters



Why use a CMB telescope for transients?

SPT is optimized for both **CMB** and point sources

- observe where CMB peaks (95 / 150 / 220 GHz)
- low noise (μK)
- obsessively reobserve the same field every day for years

Why use a CMB telescope for transients?

SPT is optimized for both CMB and **point sources**

- observe where CMB peaks (95 / 150 / 220 GHz)
- low noise (μK)
- obsessively reobserve the same field every day for years
- **wide survey area (1500 deg²)**
- **high resolution (1' beam)**
- **good localization ($O(10'')$)**

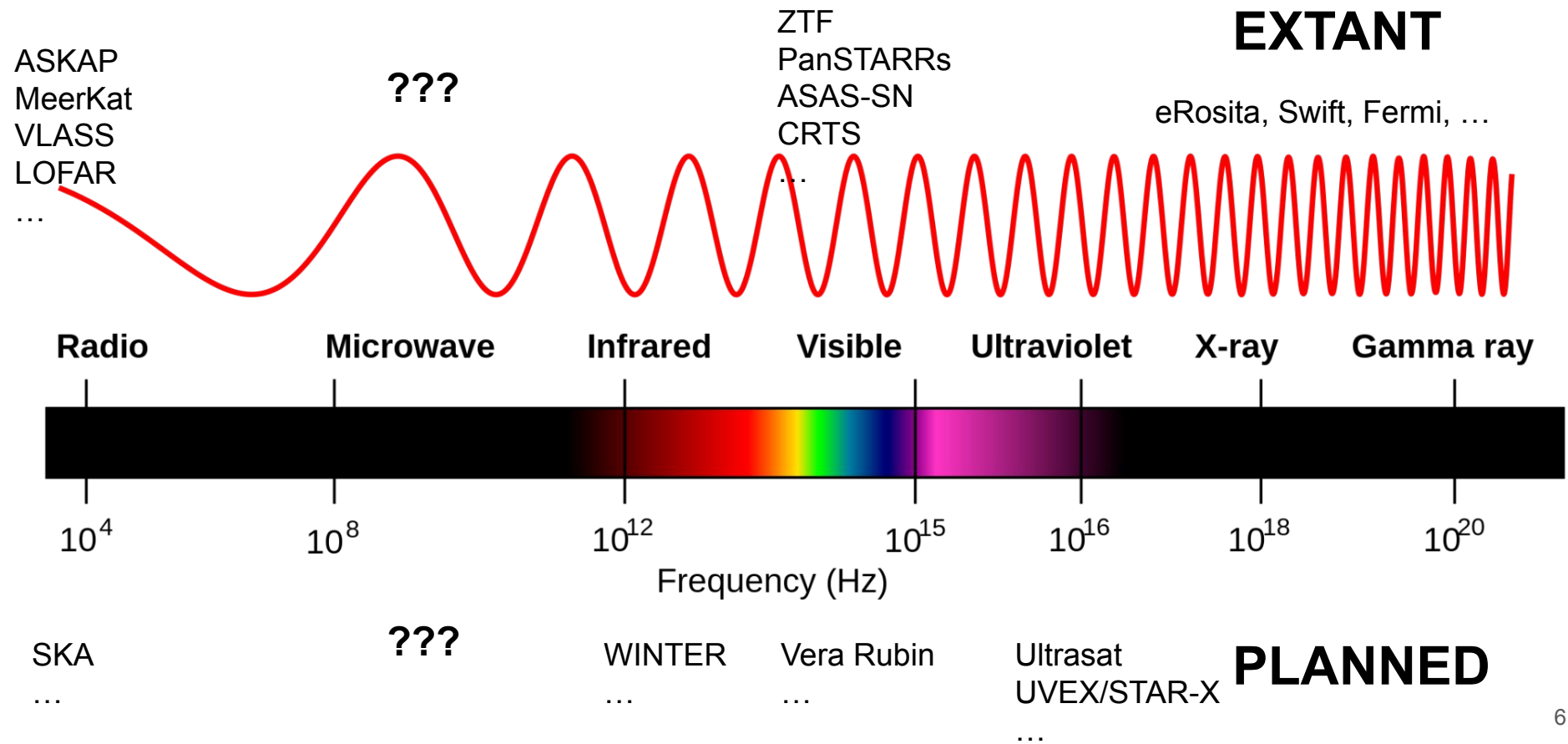
Why use a CMB telescope for transients?

SPT is optimized for both CMB and point sources

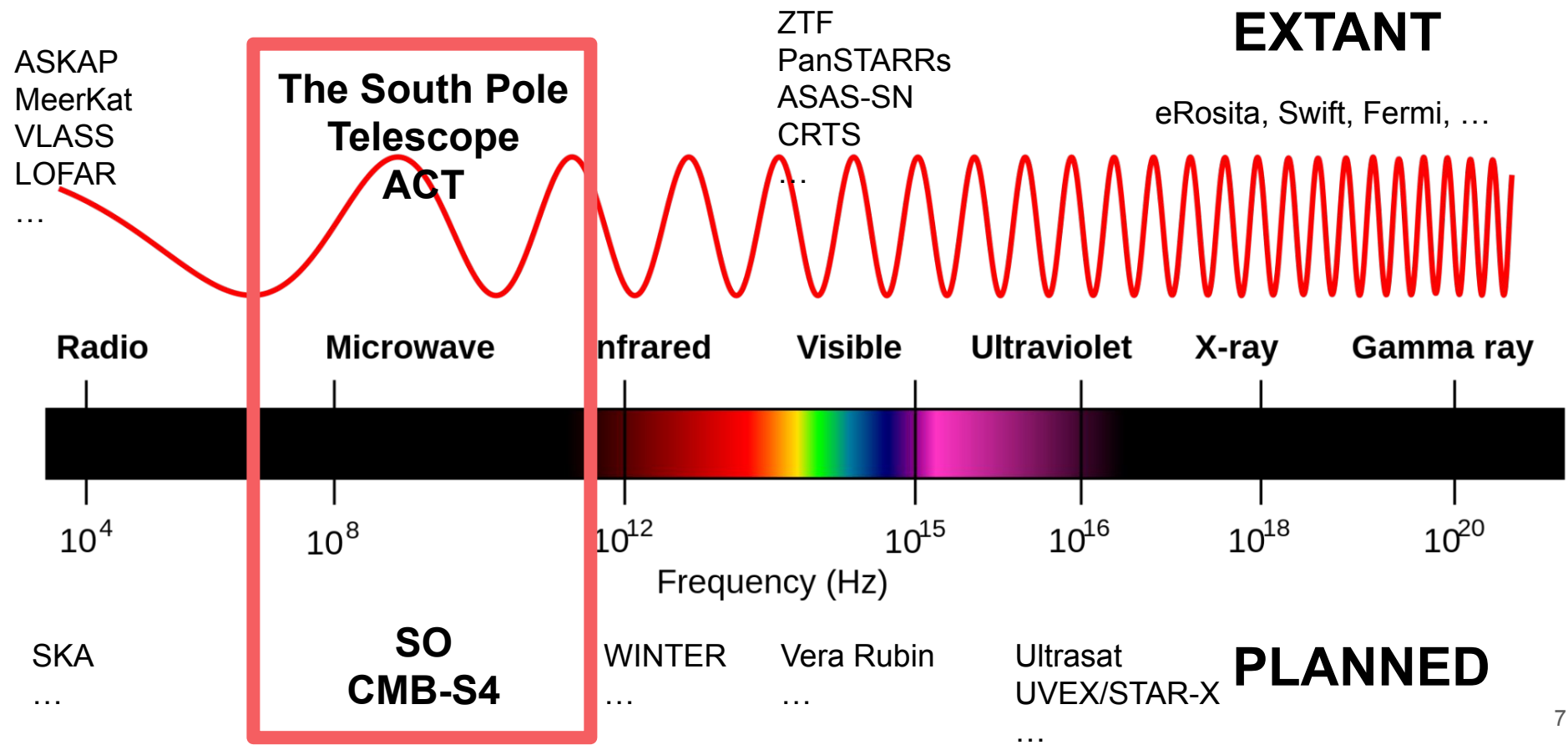
- observe where CMB peaks (95 / 150 / 220 GHz)
- low noise (μK)
- obsessively reobserve the same field every day for years
- wide survey area (1500 deg^2)
- high resolution (1' beam)
- good localization ($O(10'')$)

for every point in our field: 12 hour cadence, 3-band lightcurves with 6 / 7 / 25 mJy noise levels from 2019 – 2024+

The transient survey landscape



The transient survey landscape



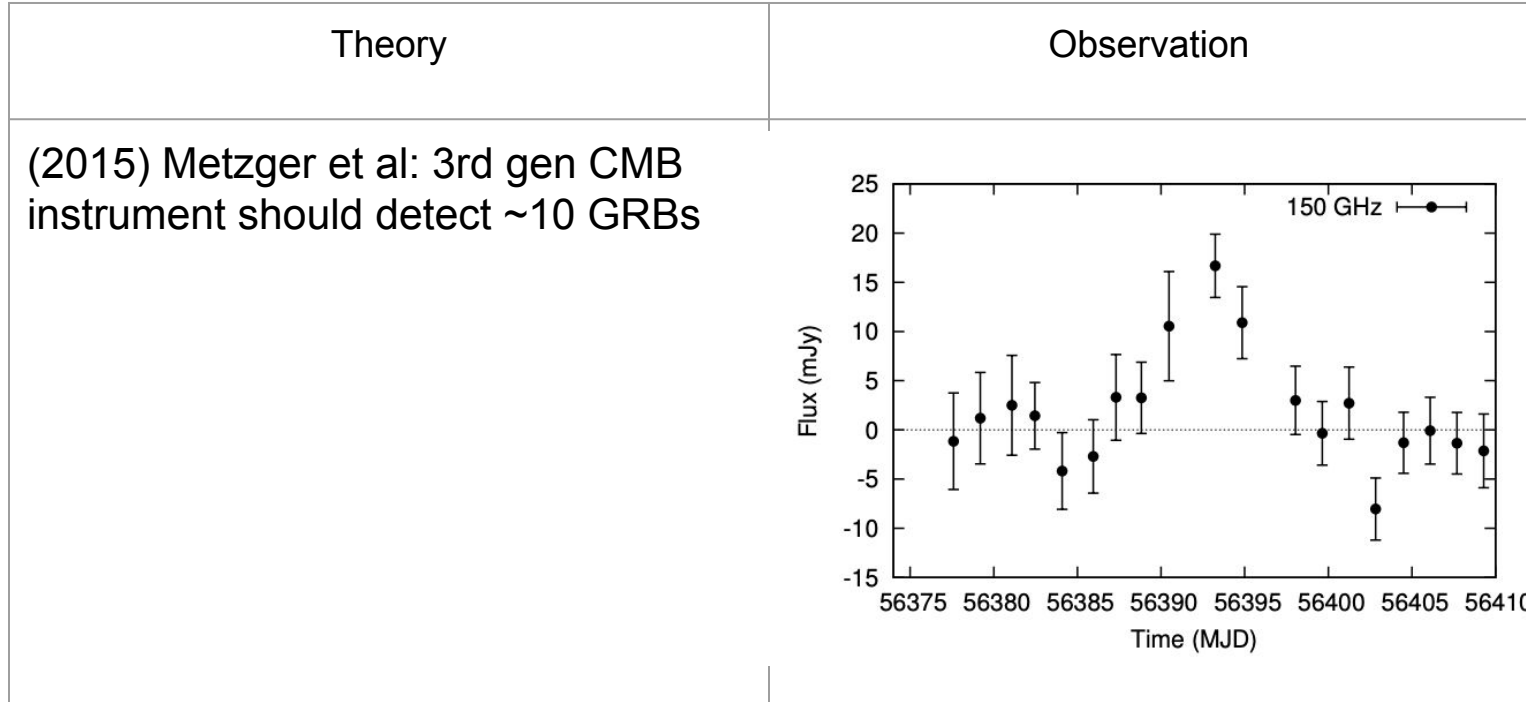
The science case for untriggered millimeter transients

Theory	Observation
(2015) Metzger et al: 3rd gen CMB instrument should detect ~10 GRBs	

The science case for untriggered millimeter transients

Theory	Observation
(2015) Metzger et al: 3rd gen CMB instrument should detect ~10 GRBs	(2016) SPT, Whitehorn et al: one 20 mJy event, 7 day duration, no counterpart

The science case for untriggered millimeter transients



Whitehorn et al 2016

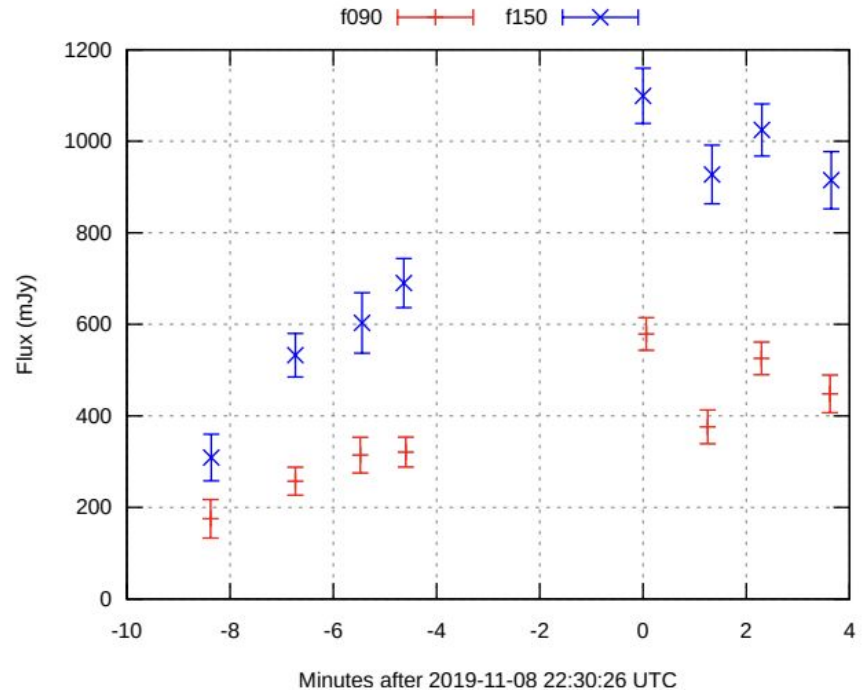
The science case for untriggered millimeter transients

Theory	Observation
(2015) Metzger et al: 3rd gen CMB instrument should detect ~10 GRBs	(2016) SPT, Whitehorn et al: one 30 mJy event, 3 day duration, no counterpart (2020) ACT, Naess et al: 3 flare stars

The science case for untriggered millimeter transients

Theory

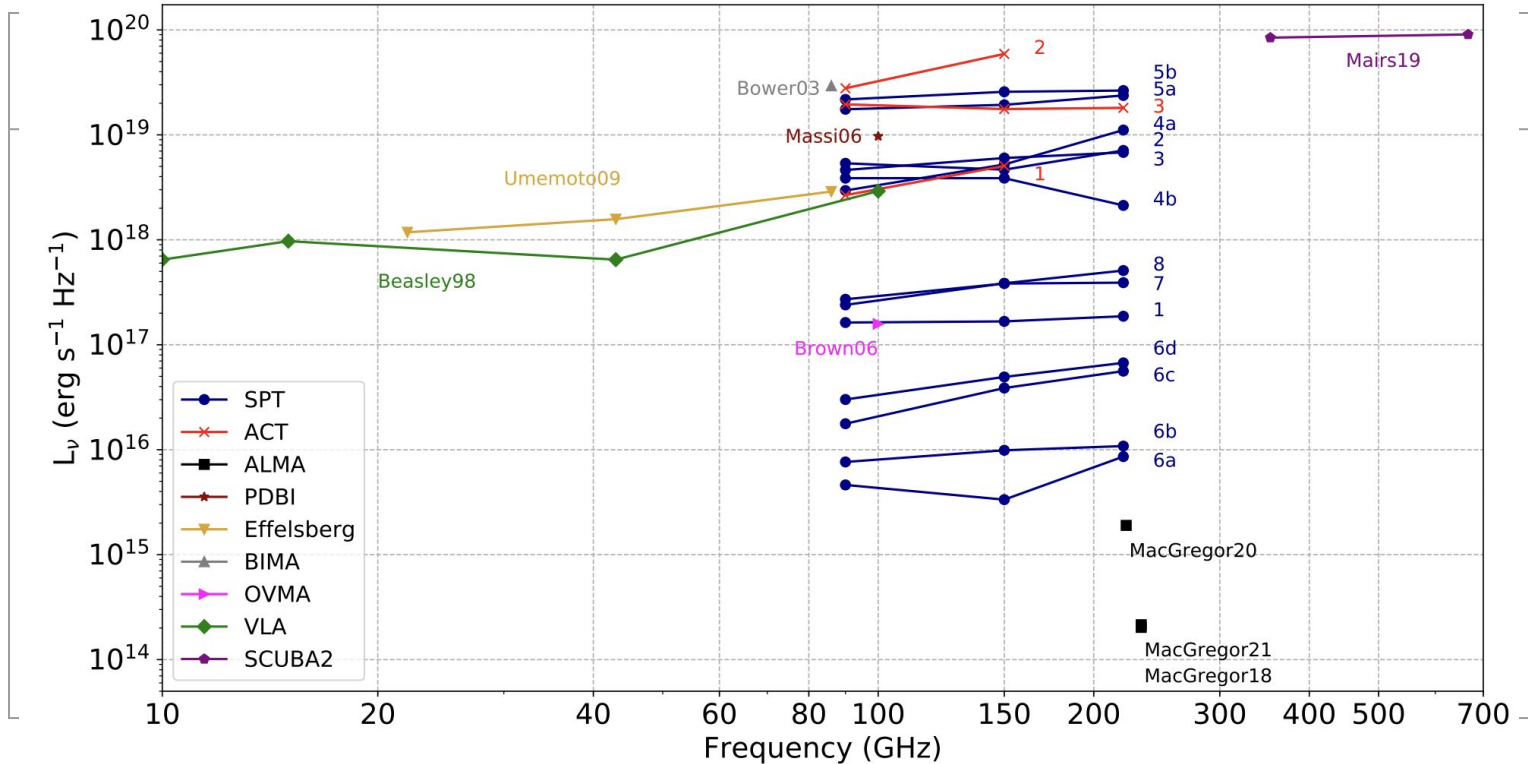
(2015) Metzger et al: 3rd gen CMB instrument should detect ~10 GRBs



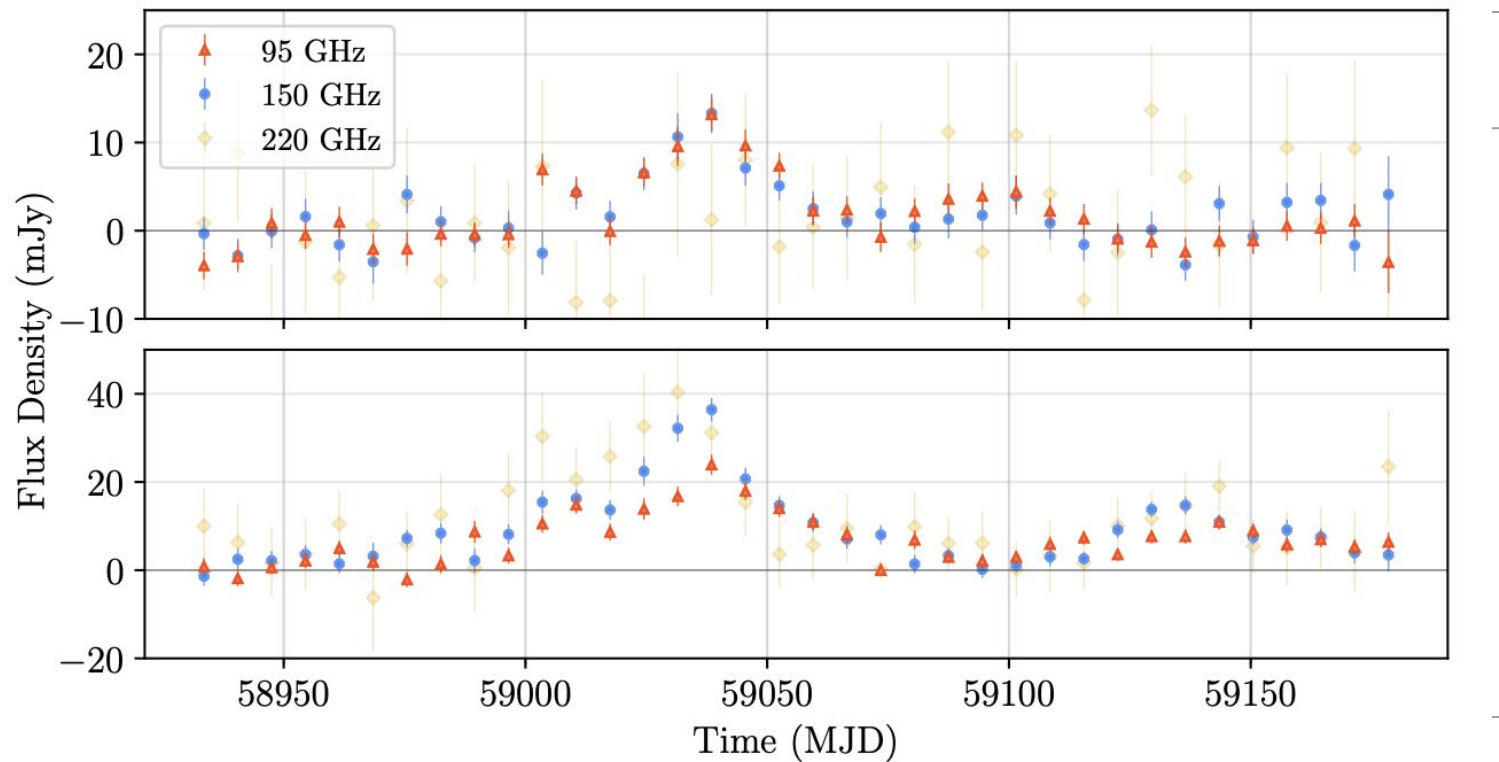
The science case for untriggered millimeter transients

Theory	Observation
<p>(2015) Metzger et al: 3rd gen CMB instrument should detect ~10 GRBs</p>	<p>(2016) SPT, Whitehorn et al: one 30 mJy event, 3 day duration, no counterpart</p> <p>(2020) ACT, Naess et al: 3 flare stars</p> <p>(2021) SPT, Guns et al: 8 flare stars, 2 AGN</p>

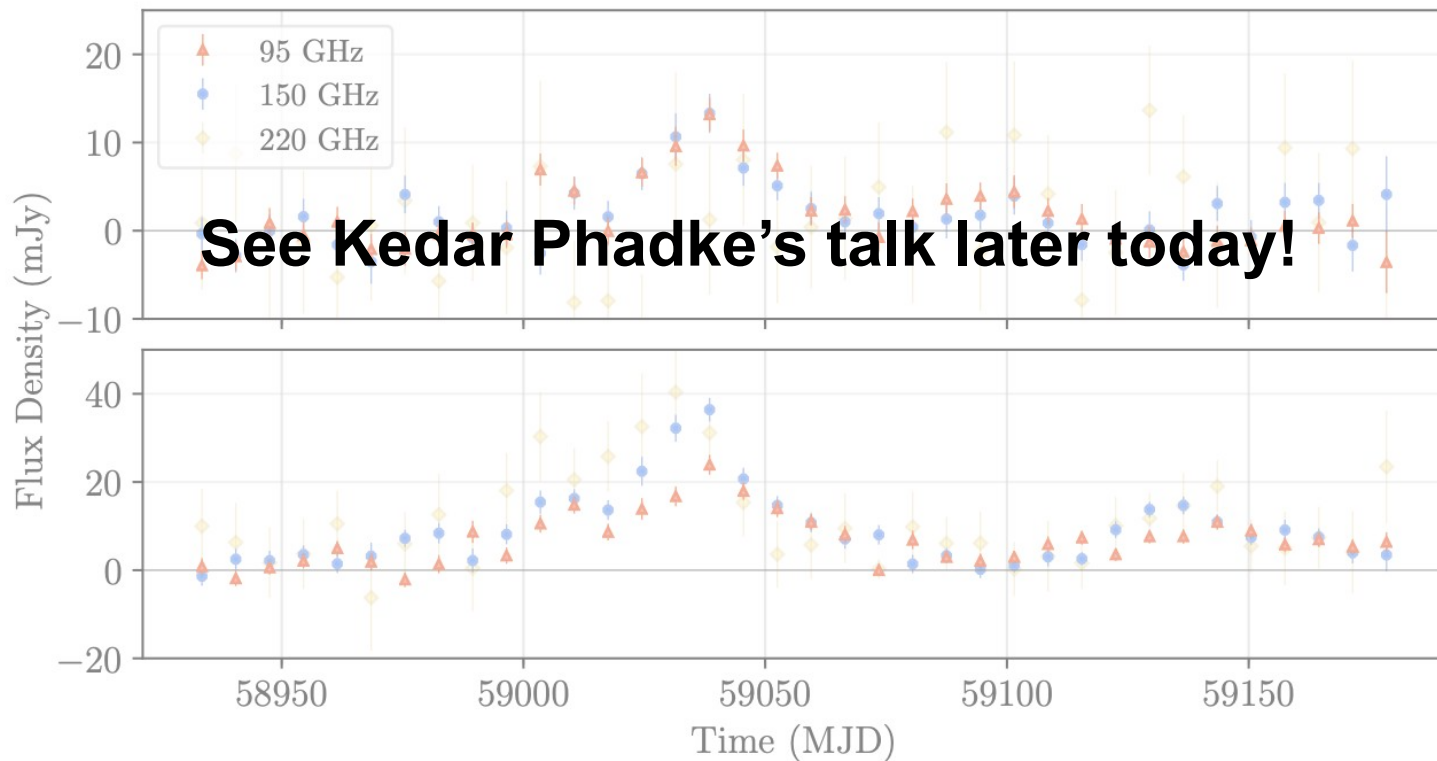
The science case for untriggered millimeter transients



The science case for untriggered millimeter transients



The science case for untriggered millimeter transients



The science case for untriggered millimeter transients

Theory	Observation
<p>(2015) Metzger et al: 3rd gen CMB instrument should detect ~10 GRBs</p>	<p>(2016) SPT, Whitehorn et al: one 30 mJy event, 3 day duration, no counterpart</p> <p>(2020) ACT, Naess et al: 3 flare stars</p> <p>(2021) SPT, Guns et al: 8 flare stars, 2 AGN</p> <p>(2022) ACT, Li et al: 11 flare stars</p>

The science case for untriggered millimeter transients

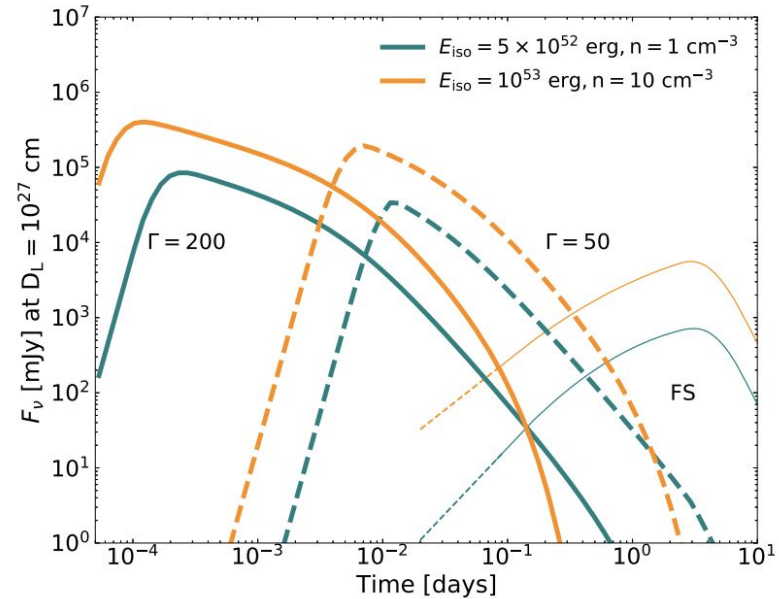
Theory	Observation
<p data-bbox="216 386 900 476">(2015) Metzger et al: 3rd gen CMB instrument should detect ~10 GRBs</p> <p data-bbox="216 543 927 787">(2022) Eftekhari et al: 3rd gen CMB instrument should detect ~10 GRB forward shocks, ~10-40 reverse shocks, maybe some TDE and Cows</p>	<p data-bbox="981 386 1665 530">(2016) SPT, Whitehorn et al: one 30 mJy event, 3 day duration, no counterpart</p> <p data-bbox="981 594 1684 635">(2020) ACT, Naess et al: 3 flare stars</p> <p data-bbox="981 698 1676 788">(2021) SPT, Guns et al: 8 flare stars, 2 AGN</p> <p data-bbox="981 851 1619 893">(2022) ACT, Li et al: 11 flare stars</p>

The science case for untriggered millimeter transients

GRB reverse shocks are exciting!

- Ultrabright and brief (2 hours – 1 day)
- Directly sample GRB ejecta, clean measurement of Lorentz factor
- Peak in mm, but only handful of detections so far (none untriggered)

Occupy same parameter space as stellar flares – SPT should see them!



Eftekhari et al 2022

The 2019 – 2022 SPT-3G short-duration transient search

Difference from 2021 search:

- 4x more data

- Better filtering and better understanding of noise and systematics

 - => increase data volume AND search down to 6.5 sigma (down from 10)

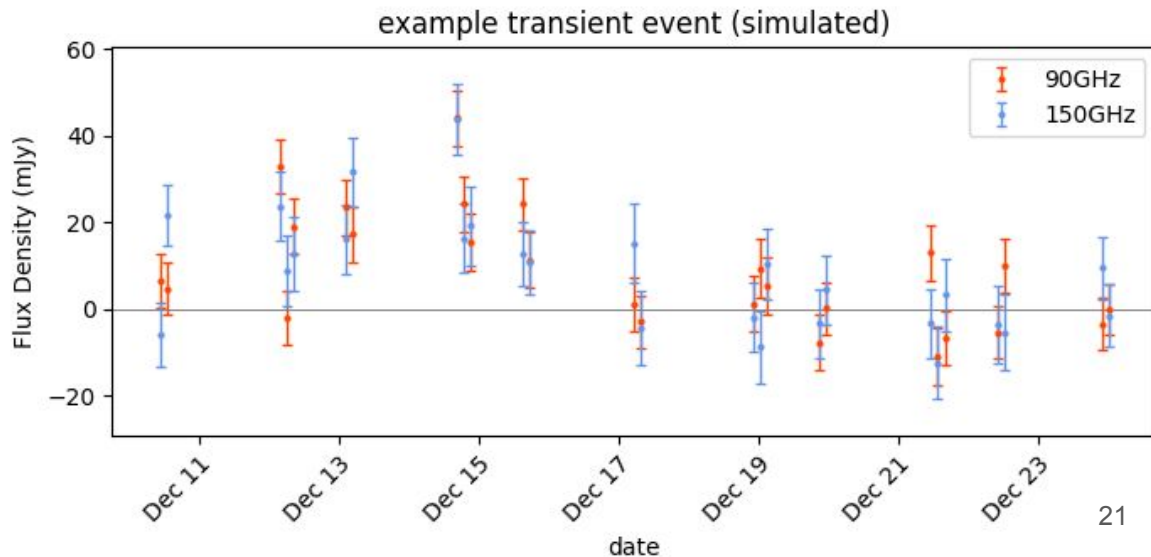
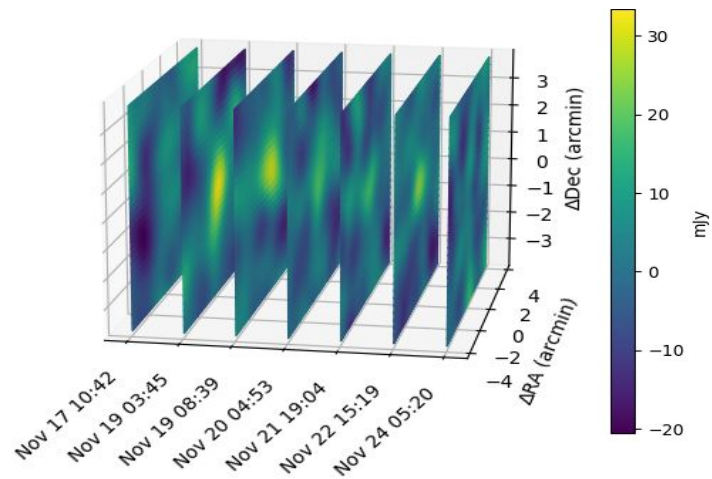
- Fit more lightcurve profiles

- Mask all AGN and other sources > 5 mJy (2.5% of survey area)

- Cut on > 10 day emission

Search algorithm

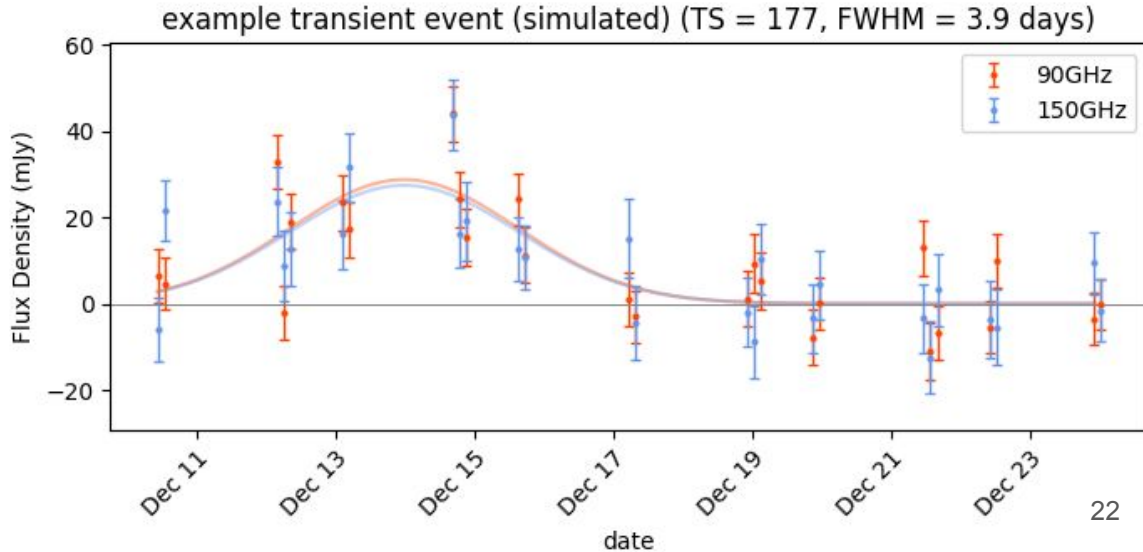
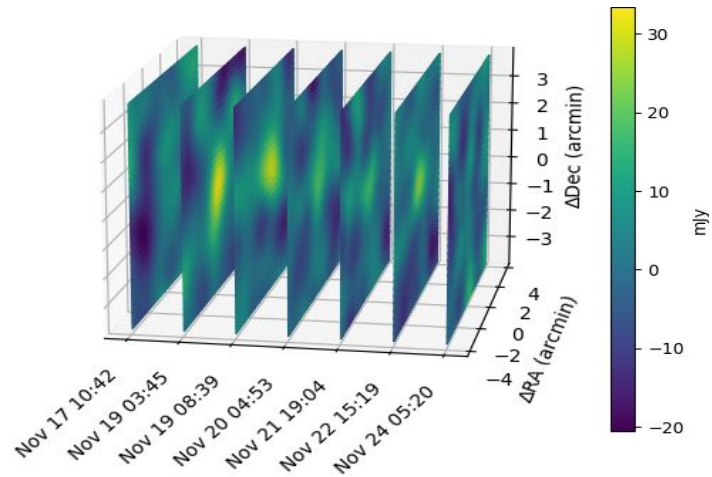
- 1) Filter maps, subtract a 3-year average, estimate noise
- 2) Select all > 3 sigma pixels and construct 12 day lightcurves
- 3) Fit a lightcurve profile and calculate the Test Statistic (likelihood ratio)



7-day image stack of SPT-3G filtered maps

Search algorithm

- 1) Filter maps, subtract a 3-year average, estimate noise
- 2) Select all > 3 sigma pixels and construct 12 day lightcurves
- 3) Fit a lightcurve profile and calculate the Test Statistic (likelihood ratio)



7-day image stack of SPT-3G filtered maps

Results!

Results!

- 100+ flares from 60+ stars!

See Chris Tandoi's talk later today!

Results!

- 100+ flares from 60+ stars!
- No detections of extragalactic flares :(

Results!

- 100+ flares from 60+ stars!
- No detections of extragalactic flares

What's next?

- Convert non-detection to rate limits – all-sky $N(>S)$

Results!

- 100+ flares from 60+ stars!
- No detections of extragalactic flares

What's next?

- Convert non-detection to rate limits – all-sky $N(>S)$
- Keep looking! Our transients pipeline is running through 2024 & beyond (& still getting better)

Results!

- 100+ flares from 60+ stars!
- No detections of extragalactic flares

What's next?

- Convert non-detection to rate limits – all-sky $N(>S)$
- Keep looking! Our transients pipeline is running through 2024 & beyond (& still getting better)
- Do a search for > 10 day signals!

Results!

- 100+ flares from 60+ stars!
- No detections of extragalactic flares

What's next?

- Convert non-detection to rate limits – all-sky $N(>S)$
- Keep looking! Our transients pipeline is running through 2024 & beyond (& still getting better)
- Do a search for > 10 day signals!
- Plan for the future – CMB-S4