

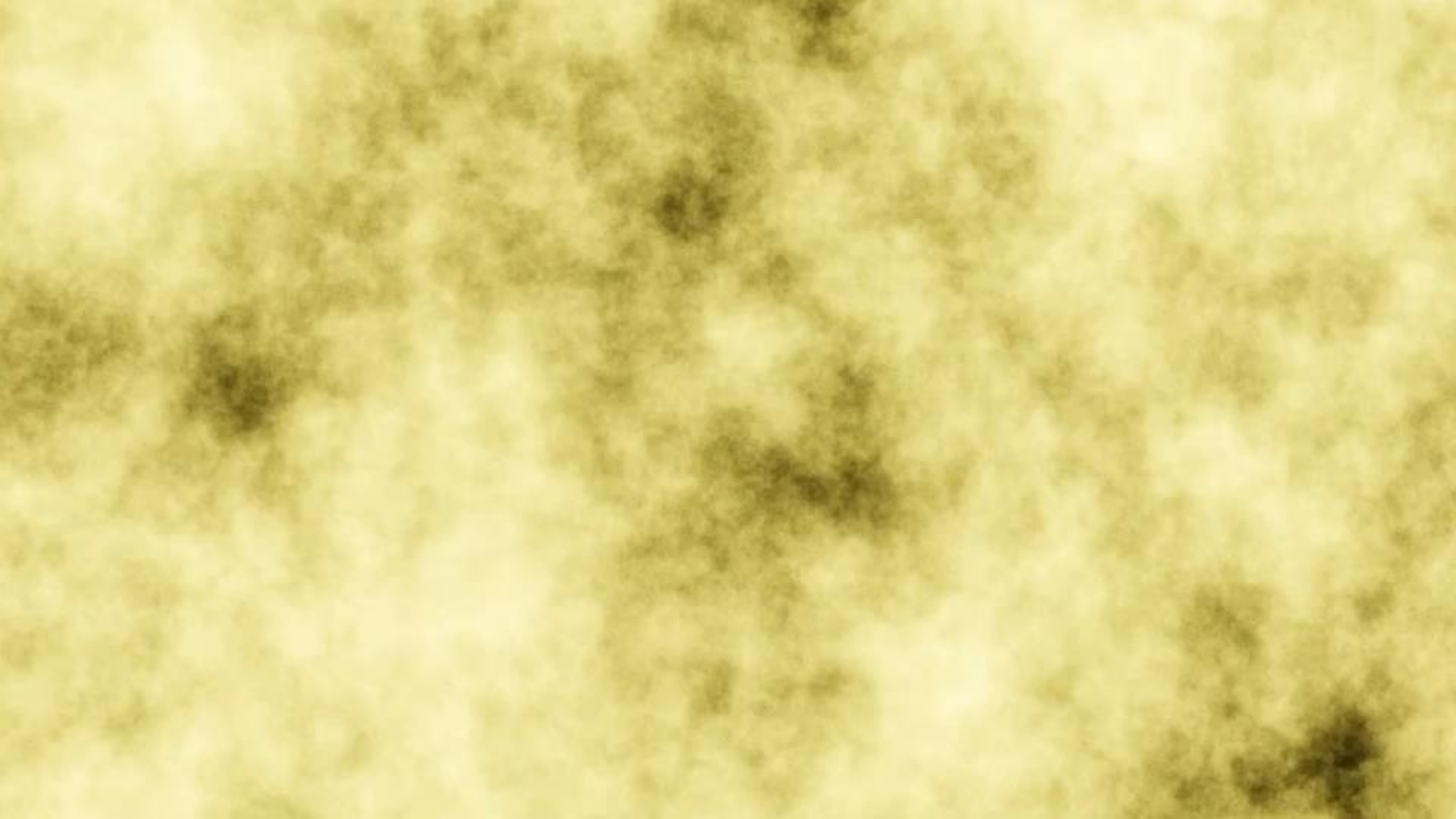
Chasing Cosmic DinoSTARS*: Using Next-Generation GW Detectors To Unravel the Fossil Origins of Compact Object Mergers

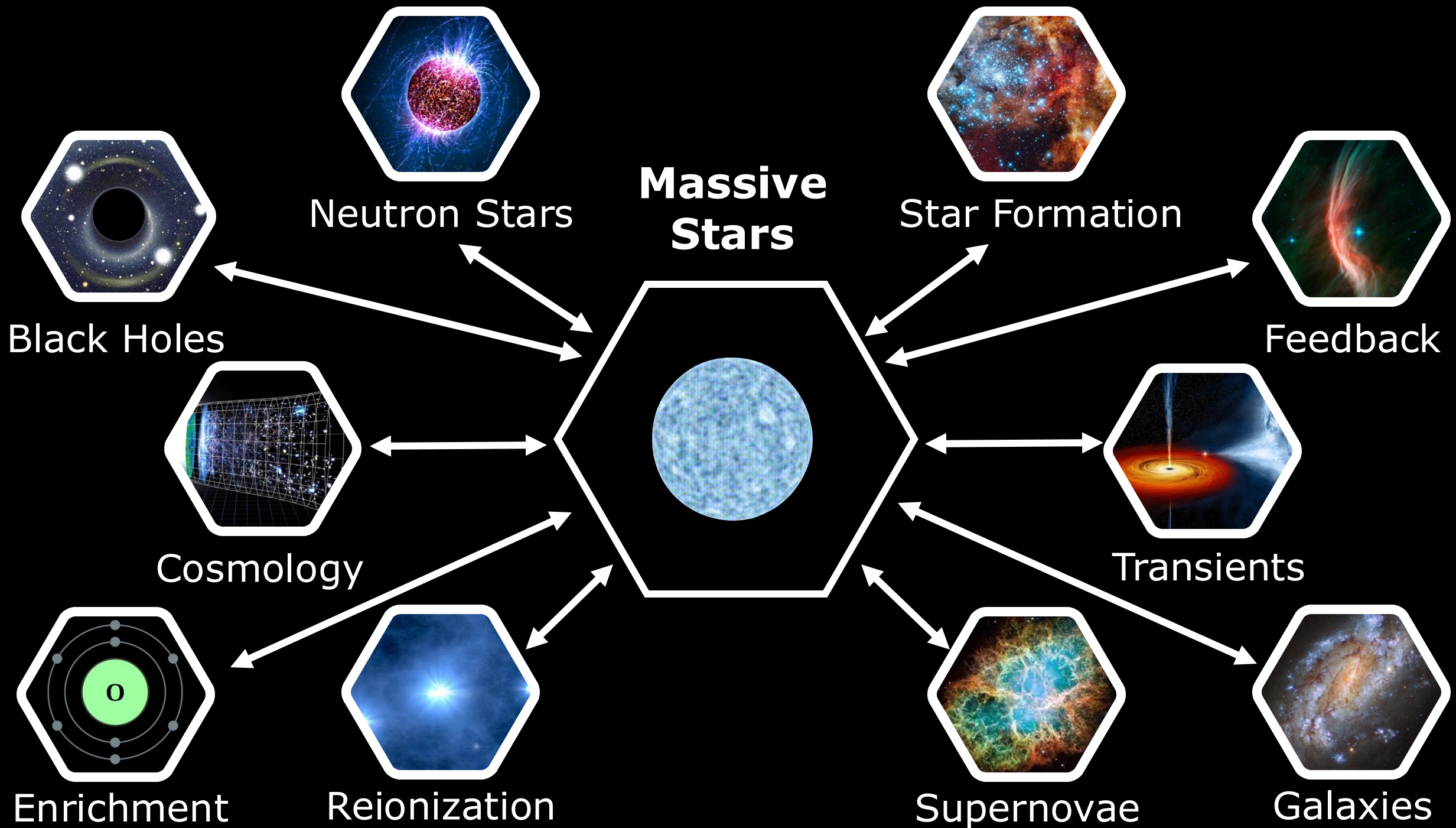
Floor Broekgaarden

Assistant Professor, Astronomy & Astrophysics, UC San Diego



* The concept “DinoSTARS” is from Lucas Ellerbroek’s son





It is challenging to observe a statistically significant population of Massive Stars...

- Massive stars live fast and die young

Kippenhahn & Weigert (1990):

$$\tau_* \propto M^{-2.8}$$

- Massive stars are rare

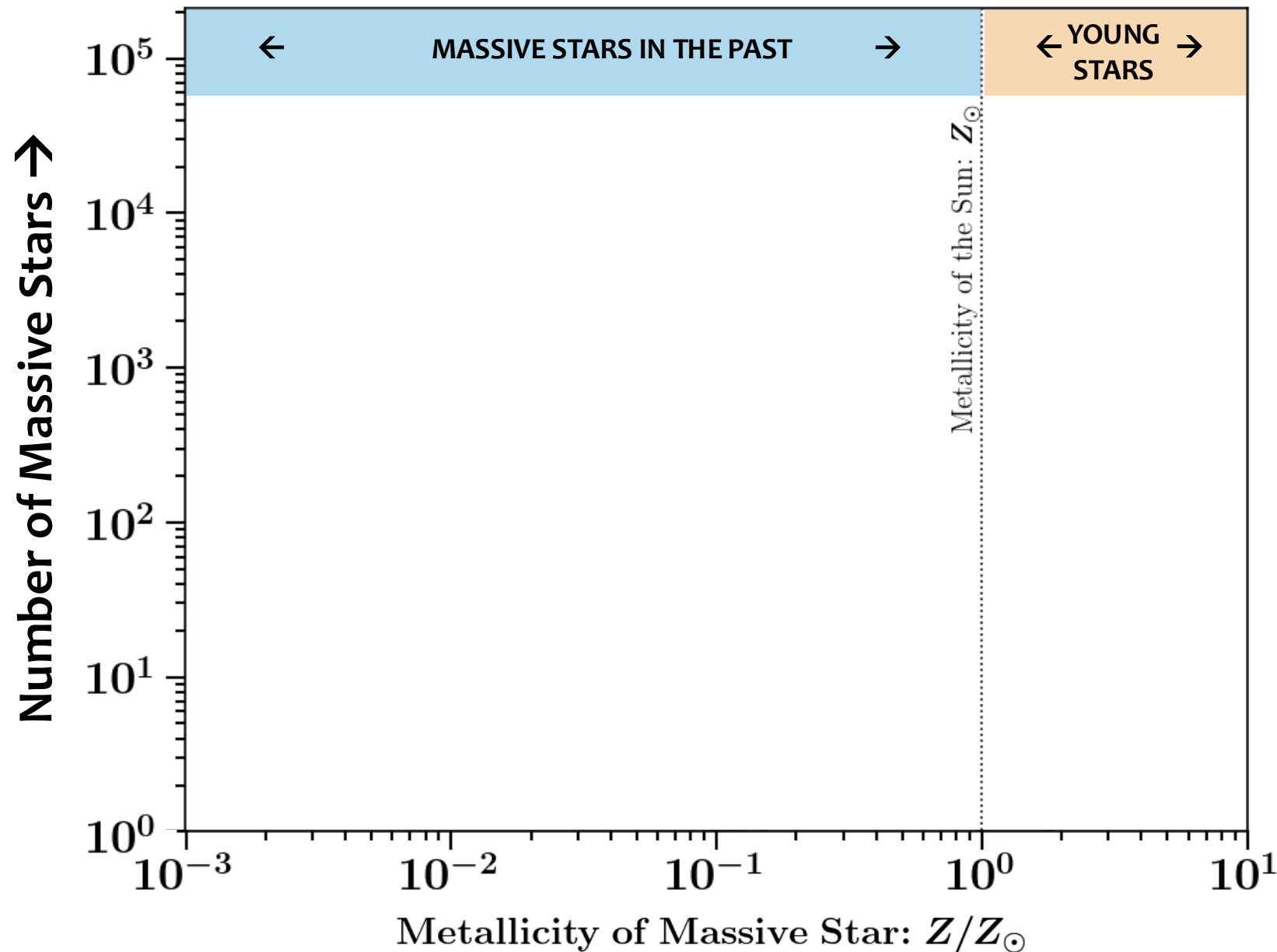
Salpeter IMF (1955):

$$p(M) \propto M^{-2.3}$$





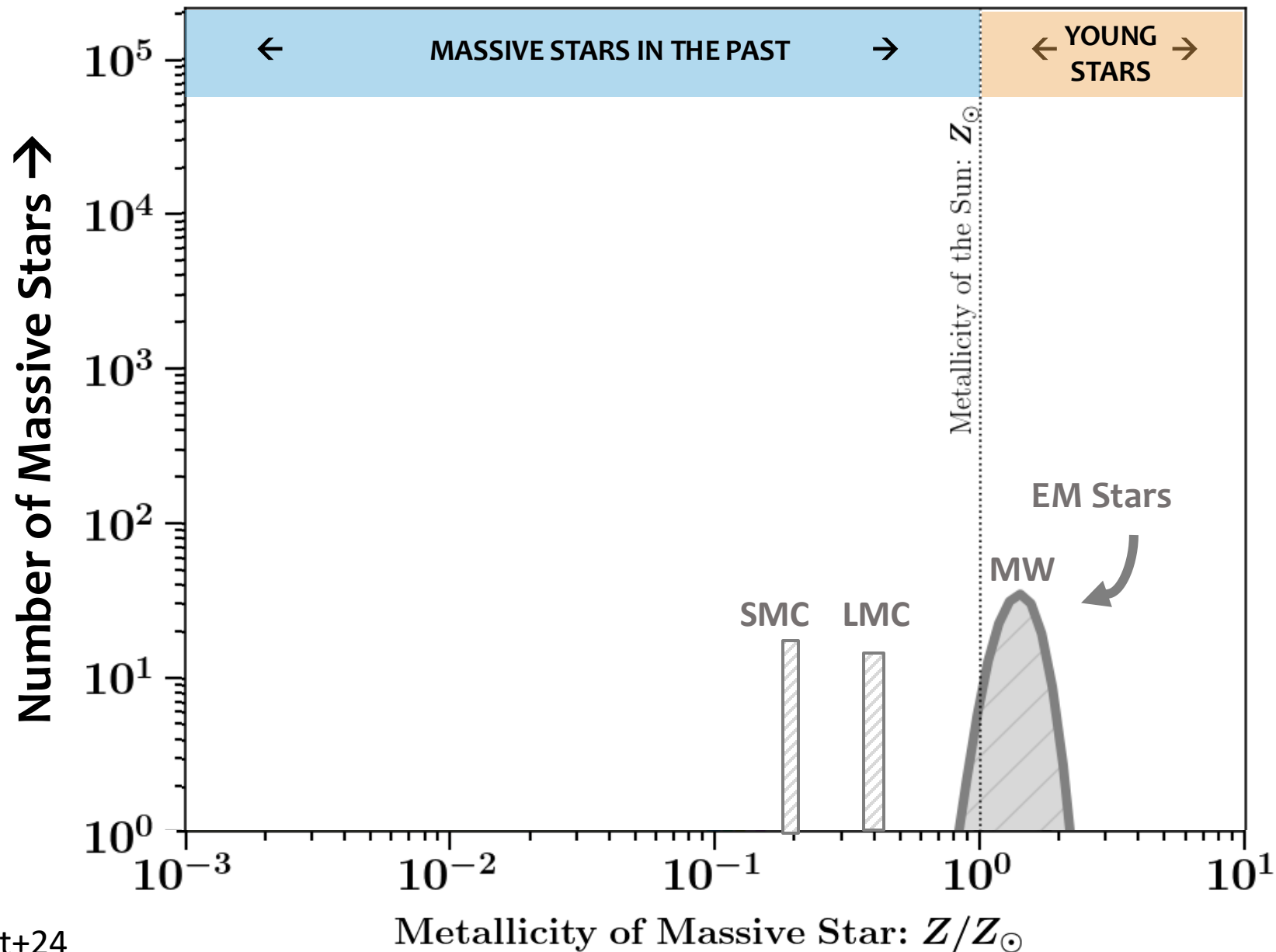
Massive stars we can study in detail*



* e.g. measure masses

Floor Broekgaarden, UCSD

Massive stars we can study in detail*



SMC: BloEm Survey (Shenar+25)

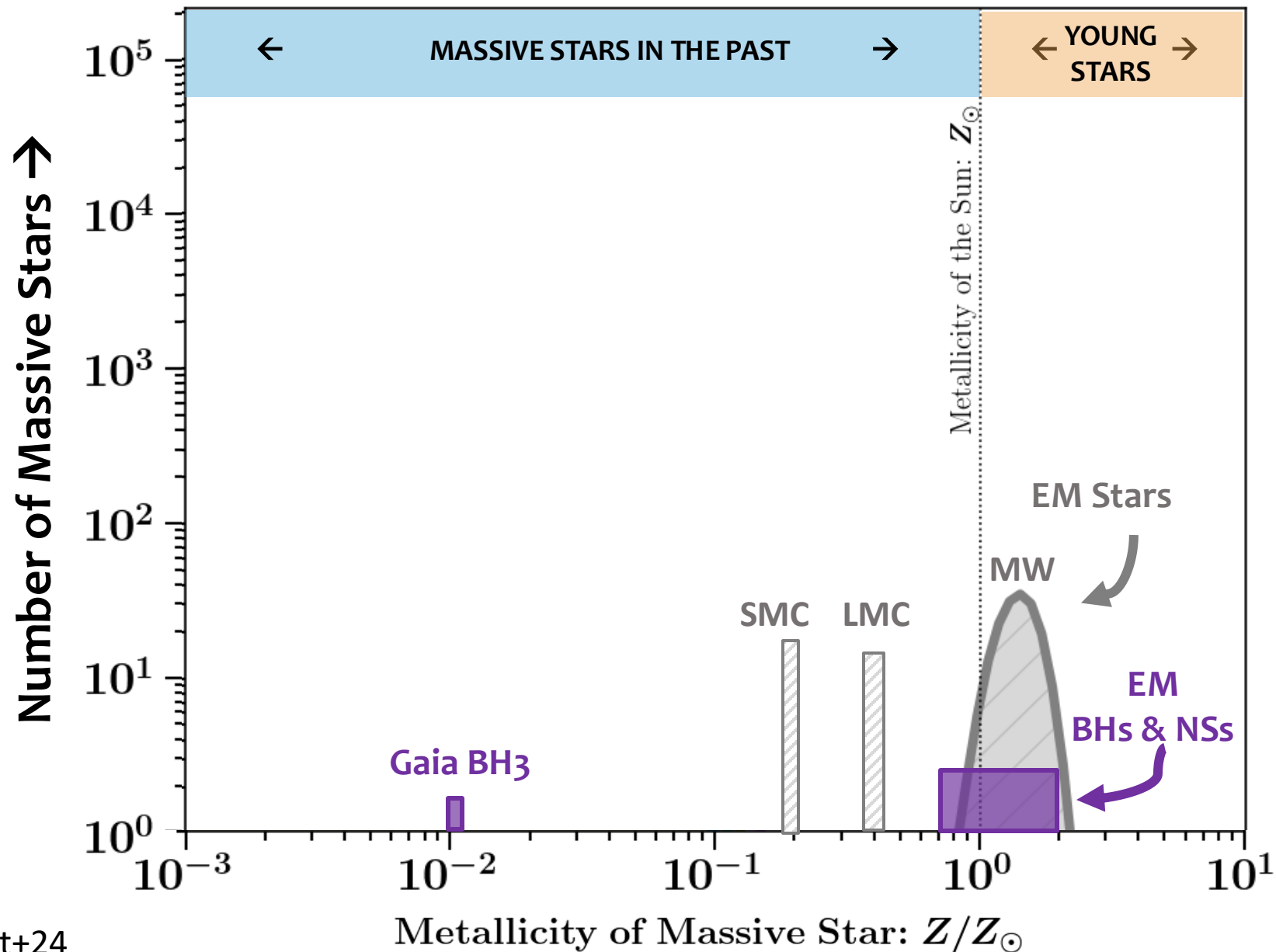
Gaia Collaboration, El-Badry+24, Balbinot+24

* e.g. measure masses

(billion years) \leftarrow Time in the Past

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Massive stars we can study in detail*



SMC: BloEm Survey (Shenar+25)

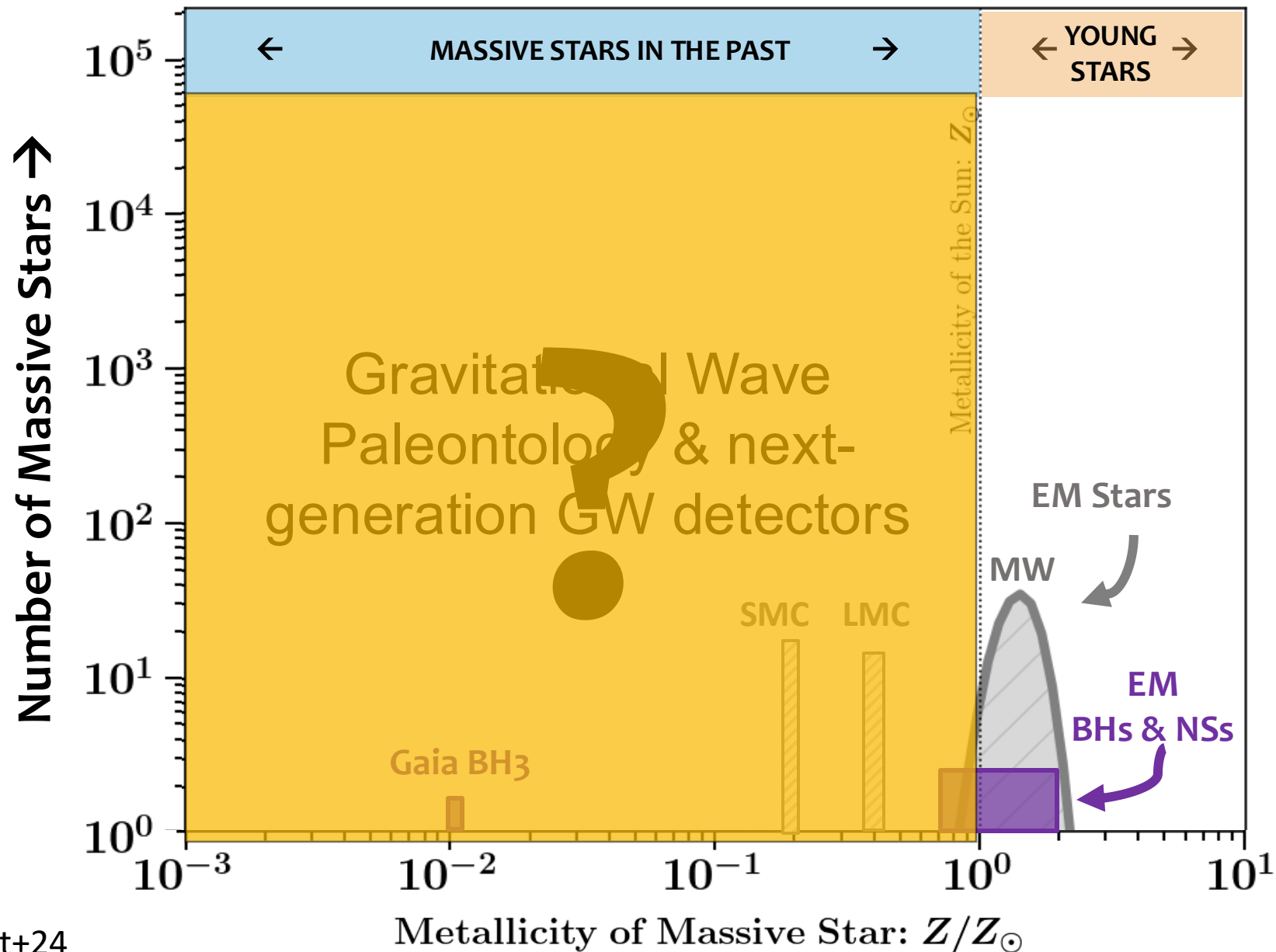
Gaia Collaboration, El-Badry+24, Balbinot+24

* e.g. measure masses

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Massive stars we can study in detail*



SMC: BloEm Survey (Shenar+25)

Gaia Collaboration, El-Badry+24, Balbinot+24

* e.g. measure masses

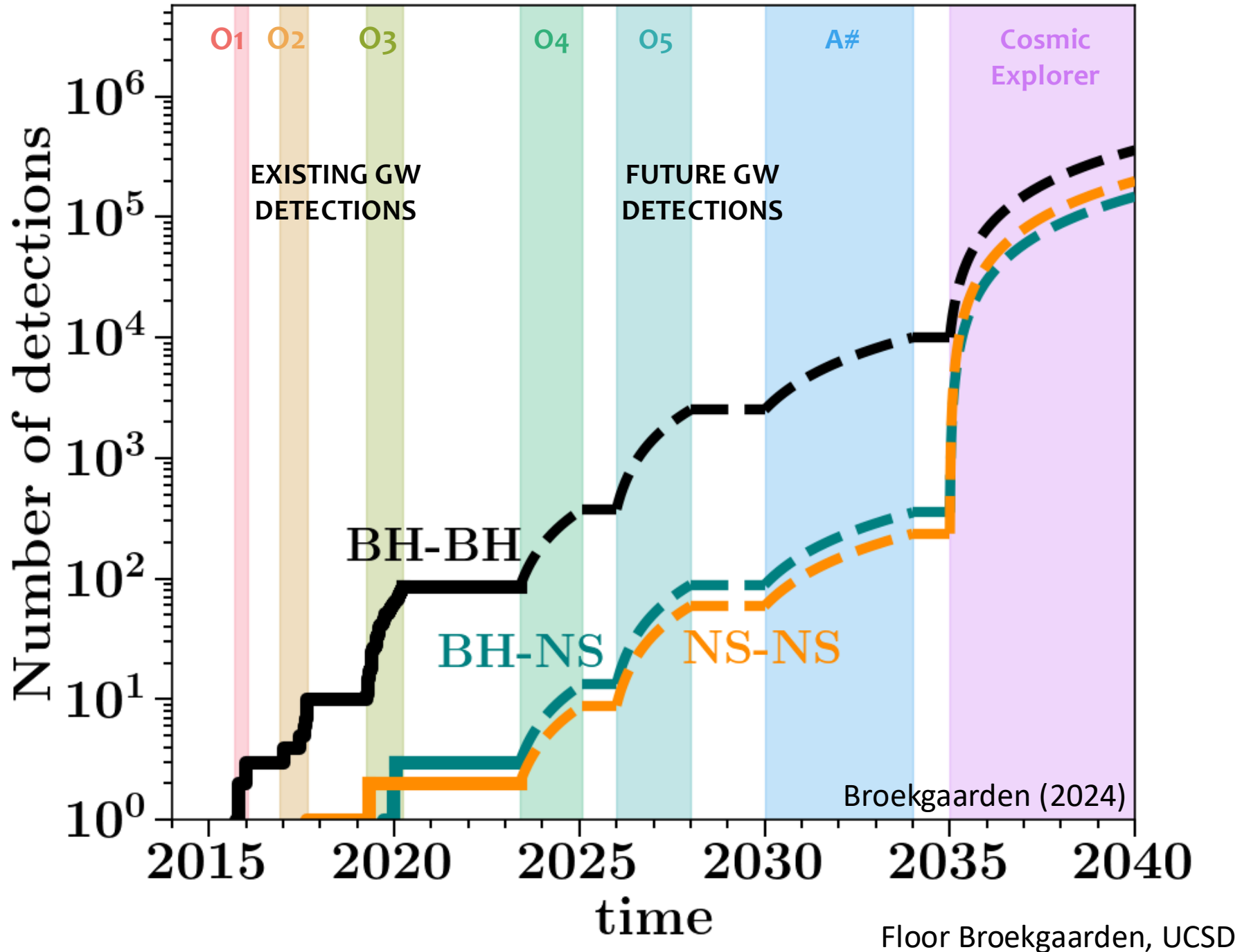
(billion years) \leftarrow Time in the Past

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Gravitational Wave
Paleontology & next-
generation GW detectors
Are poised to change our
understanding of massive
stars across cosmic time!

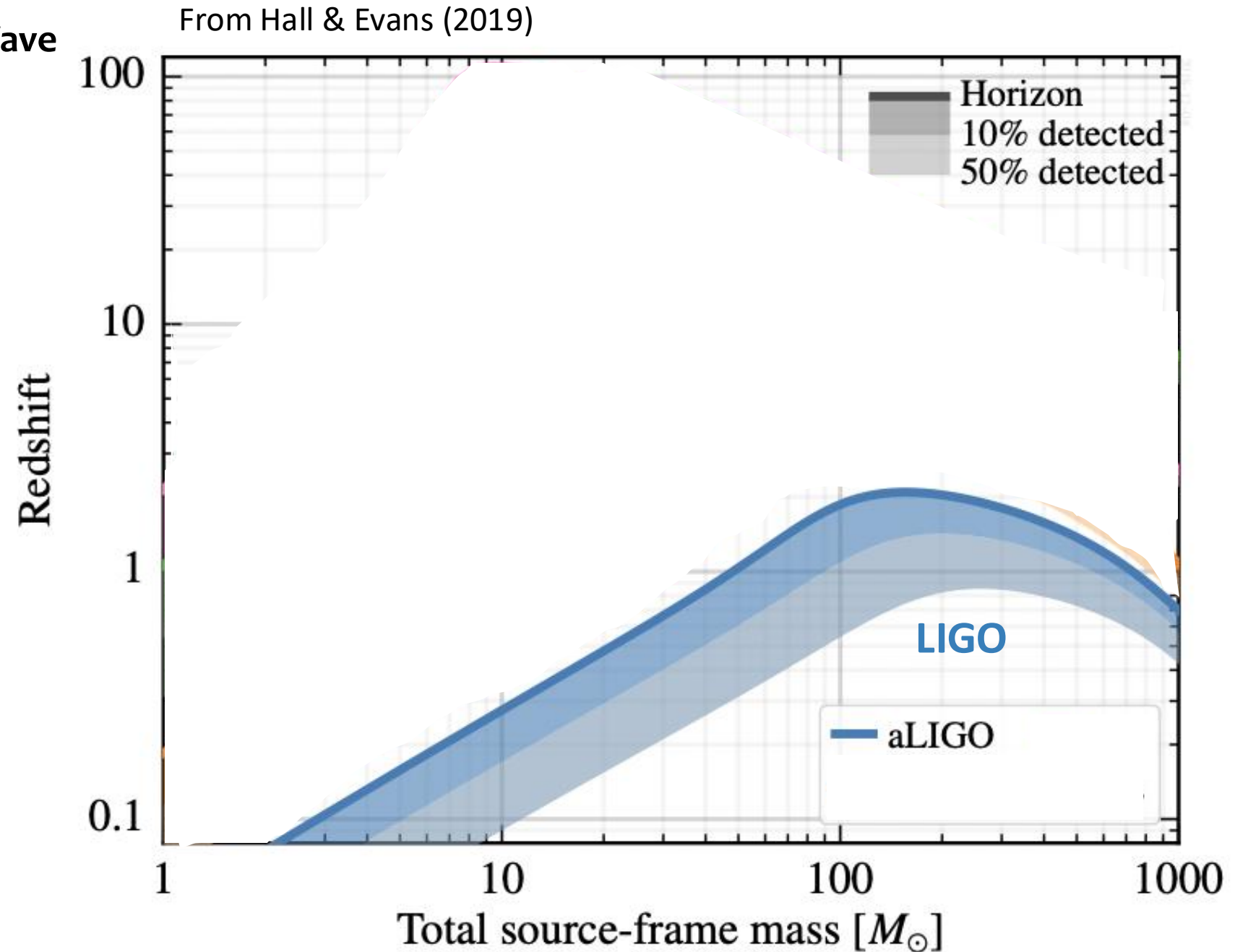
1) The Number of detected Gravitational-Wave Sources is rapidly increasing

Cosmic Explorer (& Einstein Telescope)
numbers based on: Gupta+24, Gupta+23,
Ronchini+22, Iacovelli+22, Borhanian &
Sathyaprakash+22, Baibhav+19



2) We will detect Gravitational-Wave Sources across vast Cosmic Time!

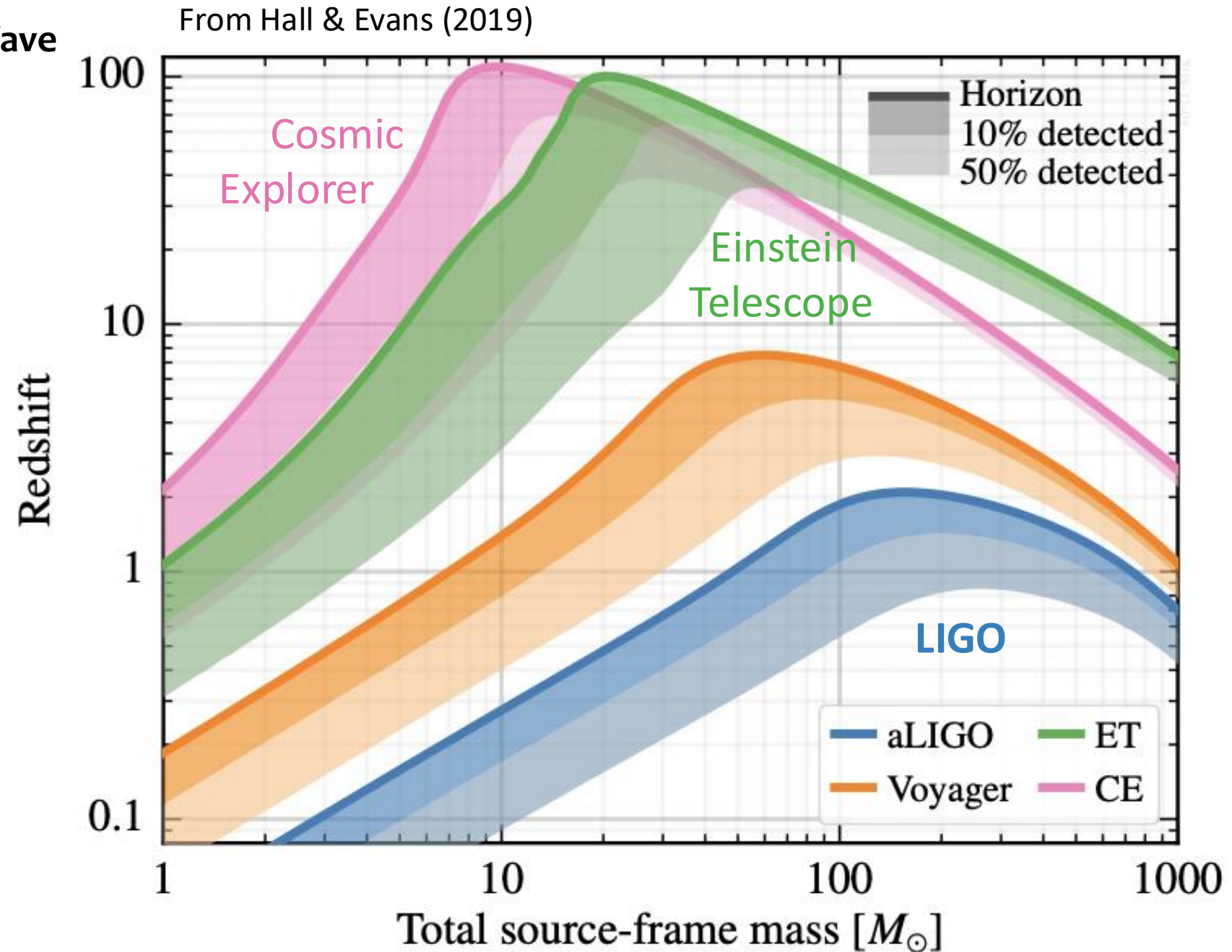
e.g. Hall & Evans +19, Evans +21, but see also Gupta+24, Gupta+23, Ronchini+22, Iacovelli+22, Borhanian & Sathyaprakash+22, Baibhav+19



2) We will detect Gravitational-Wave Sources across vast Cosmic Time!

e.g. Hall & Evans +19, Evans +21, but see also Gupta+24, Gupta+23, Ronchini+22, Iacovelli+22, Borhanian & Sathyaprakash+22, Baibhav+19

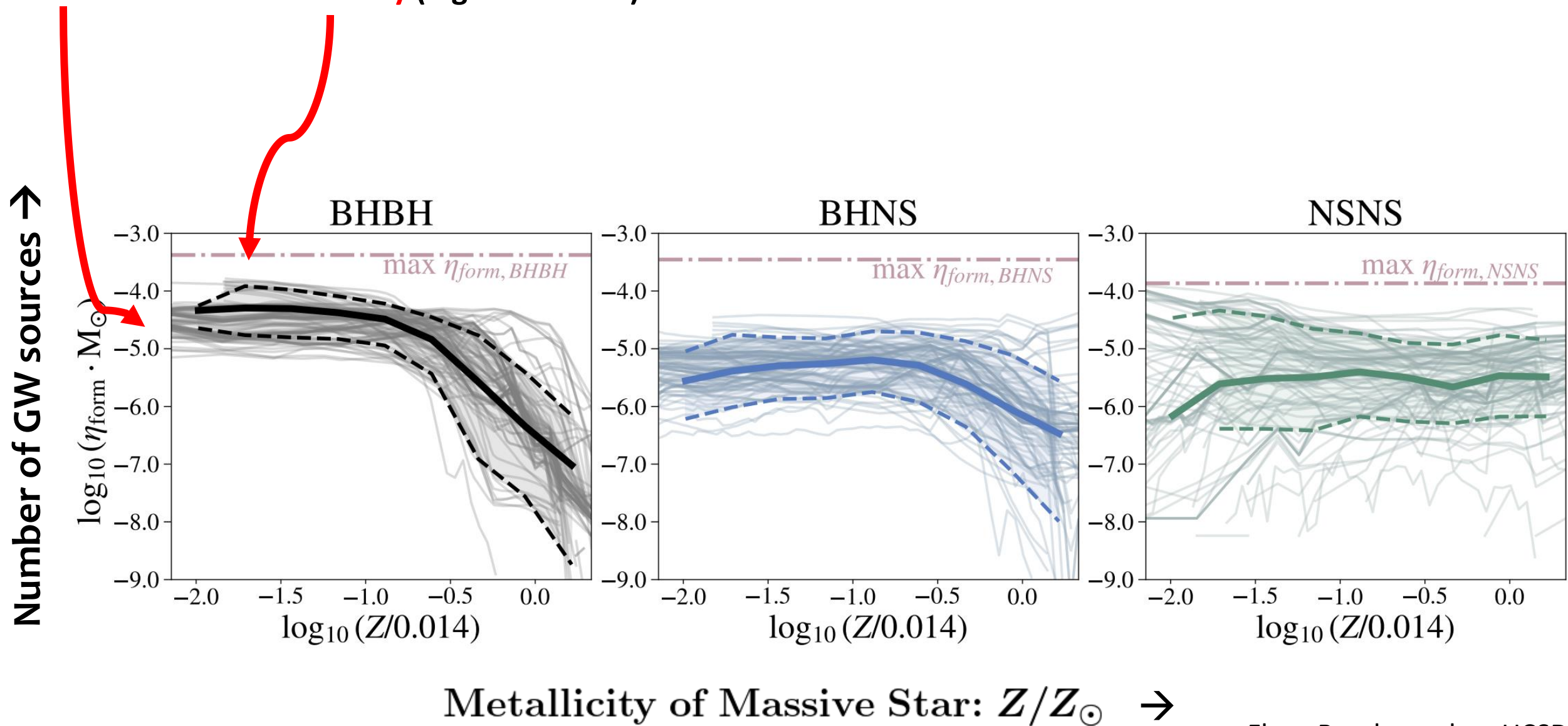
We might see **every** merging stellar mass BH-BH in the Universe!



3) Formation of stellar-mass GW sources - especially BBHs - is often significantly **boosted at lower metallicity** (higher redshift)

Figure from Van Son +25 (based on earlier work from Iorio+24)

See Giacobbo & Mapelli 2018; Klencki+2018; Neijssel+2019; Spera+19, Santoliquido+2021; Broekgaarden+2022,2023; Iorio et al. 2023, Chruslinska+23, Smith+25, van Son+25,



4) Many formation channels produce (a subset of systems with) **long delay times** of several Gyr

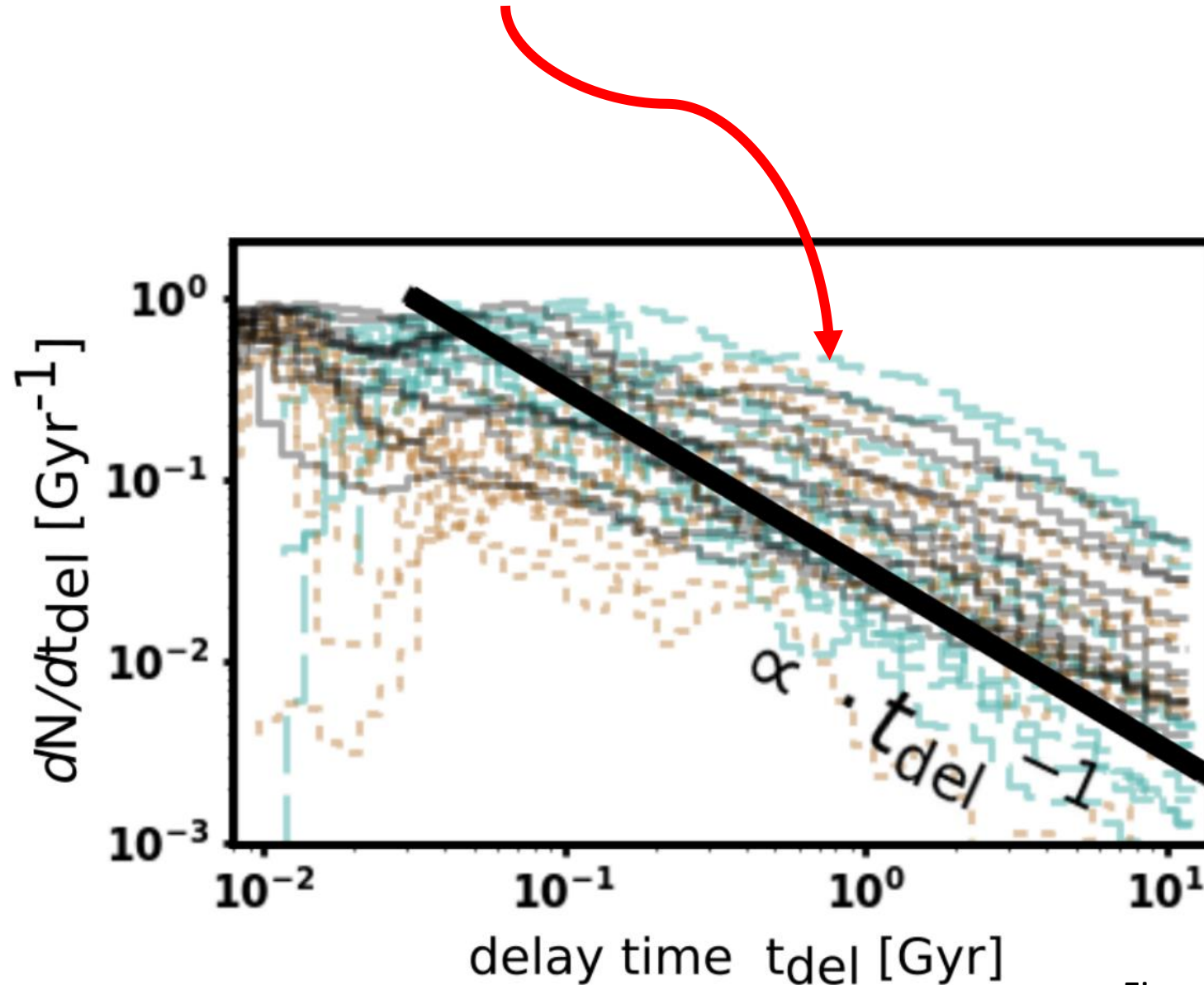
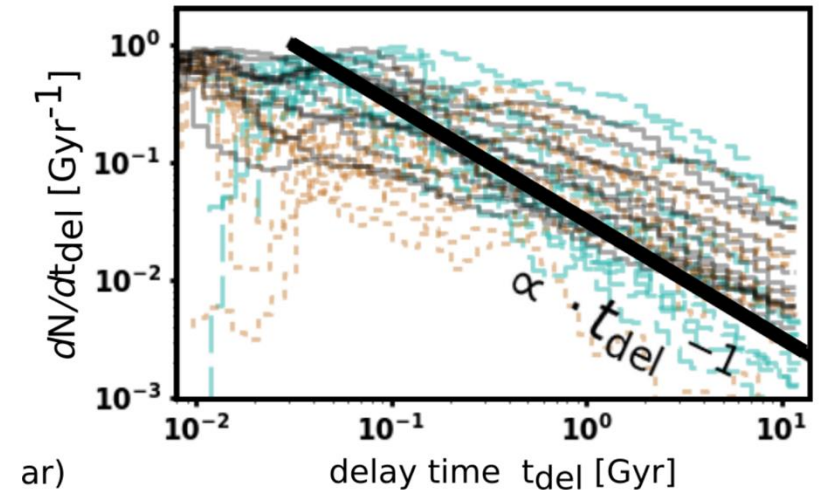
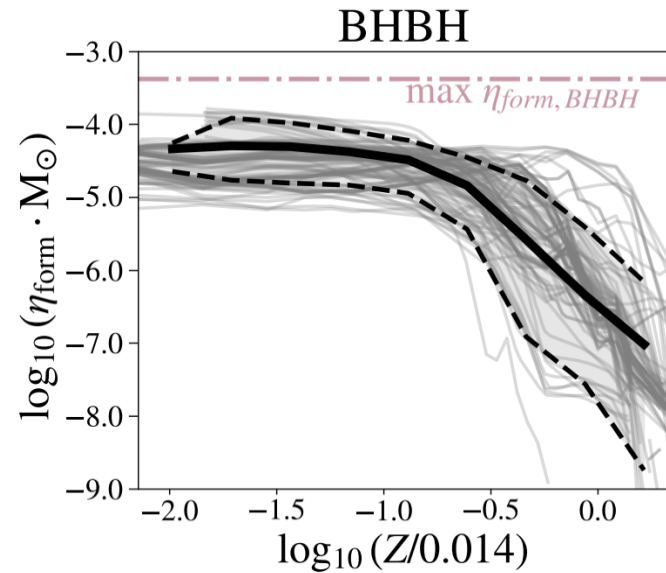
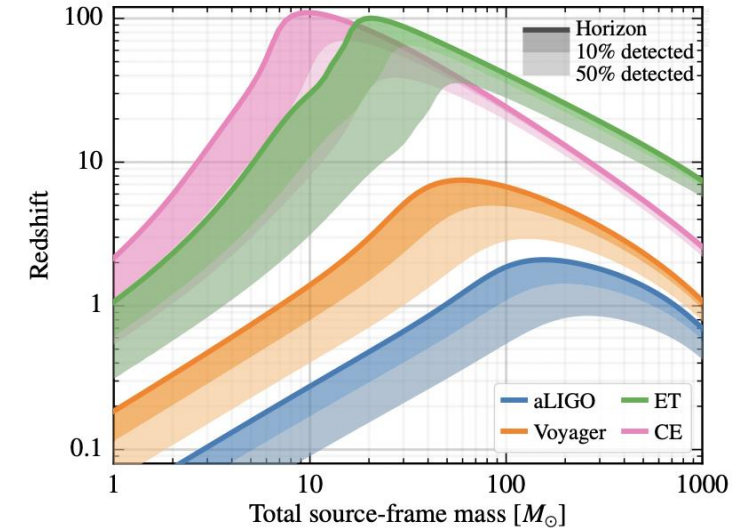
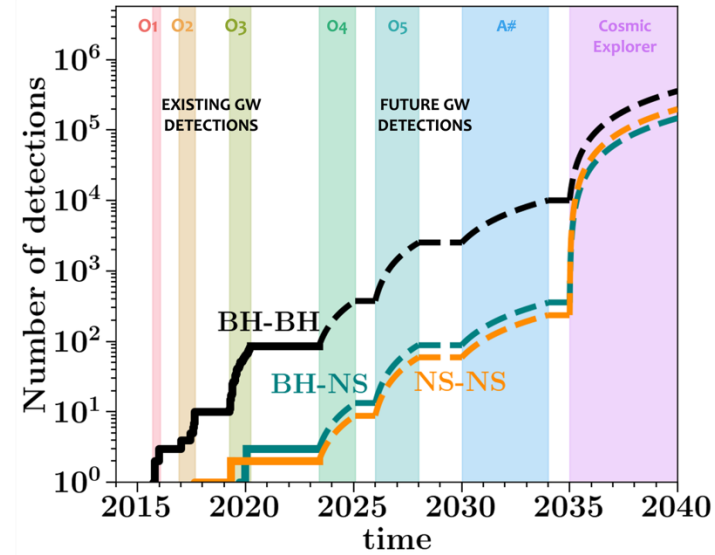


Figure from “[The Science of the Einstein Telescope](#)”+25

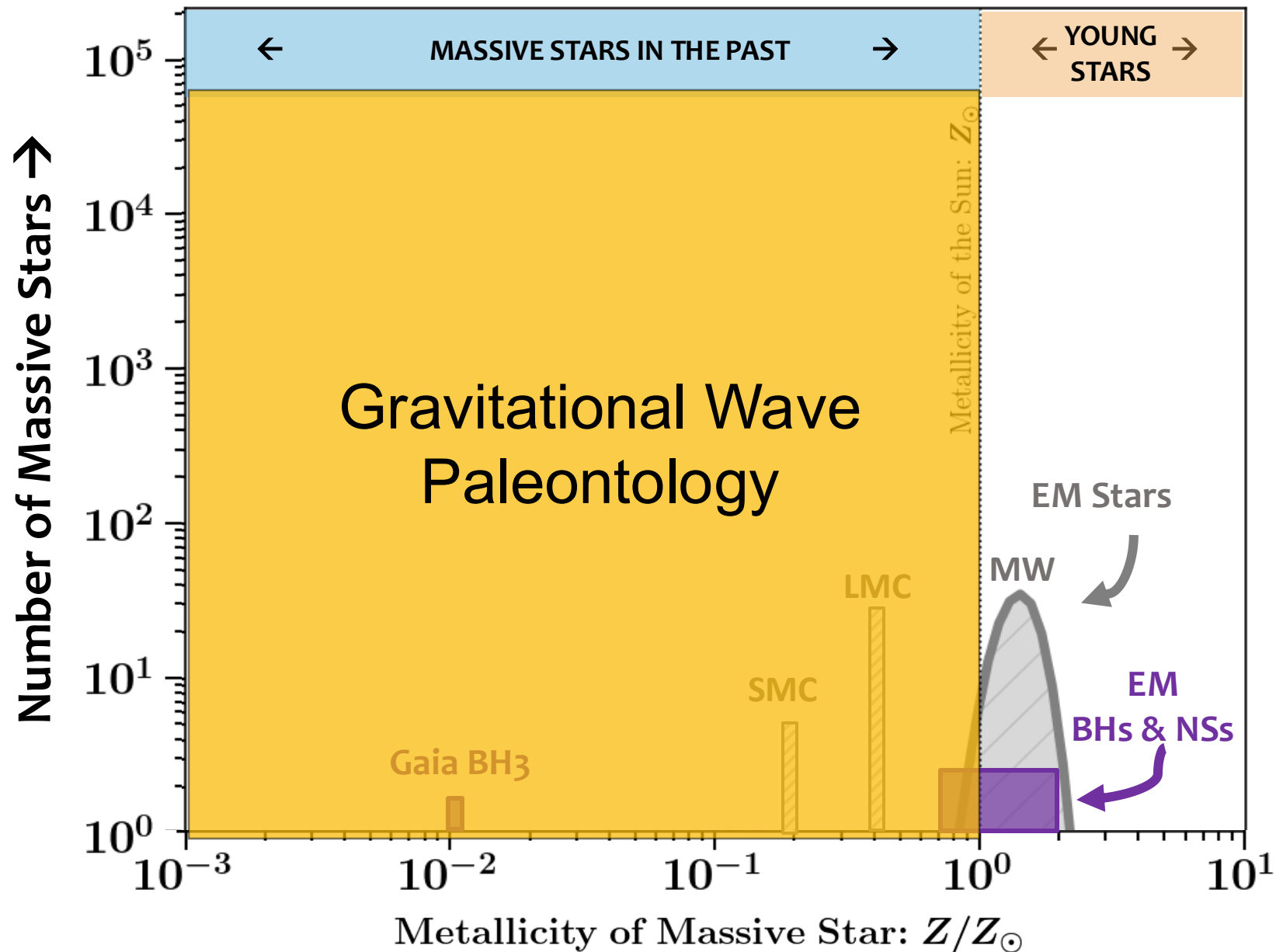
Combined, this means that GW detections - even “local” ones - already probe massive stars* across vast cosmic times & metallicities!

See e.g. Chruslinska+22, Fishbach+23,+24



*or more generally: progenitors of GW sources

Massive stars we can study in detail*



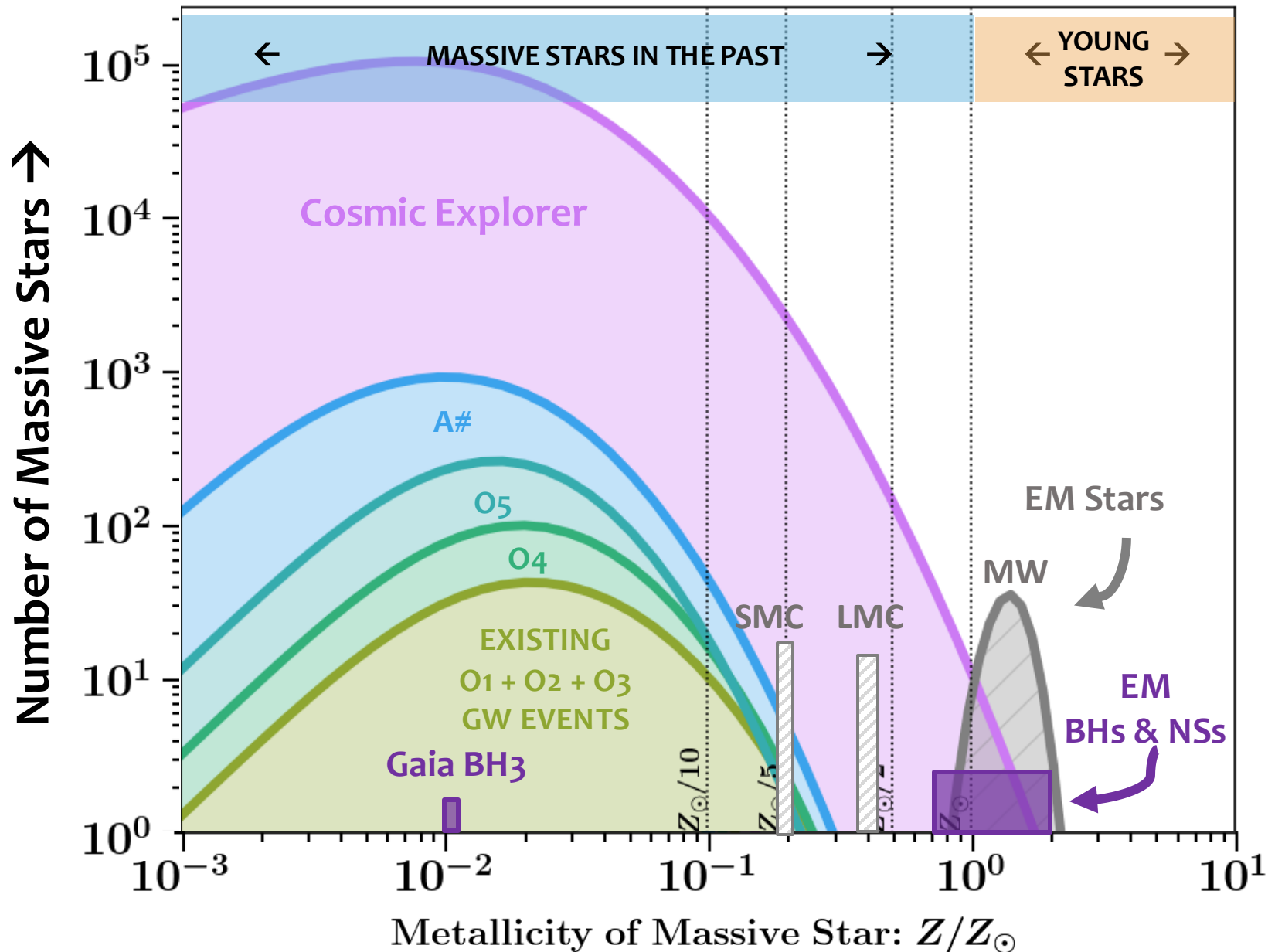
* e.g. measure masses

(billion years) ← Time in the Past

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Massive stars we can study in detail*

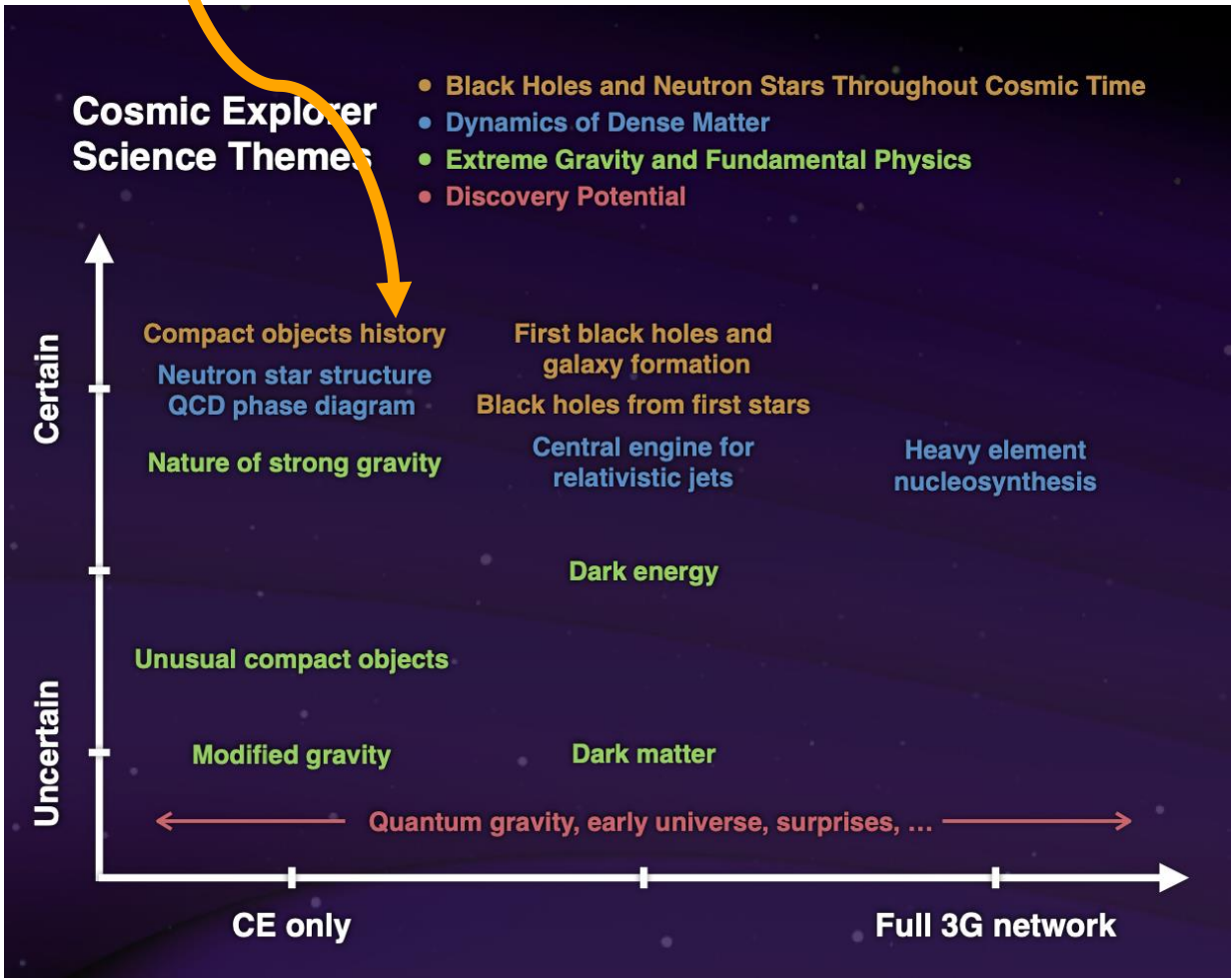
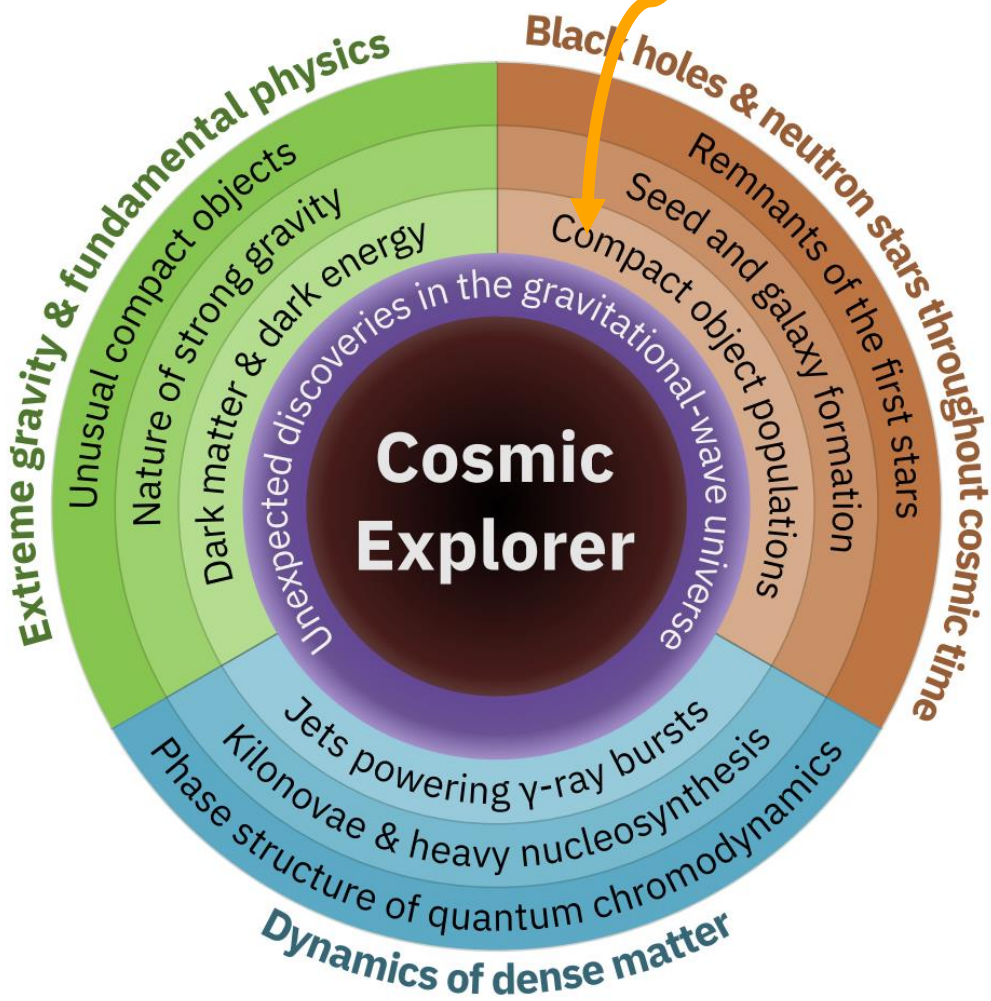
Cosmic Explorer:
“a unique instrument to study massive stars across cosmic time for \$100/star” if we can connect GW sources to formation pathway histories



* e.g. measure masses

Should we add “Exploring Massive Stars” / “Gravitational Wave Paleontology” more explicit to CE Science Case?

- low metallicity star formation - lives of massive stars - deaths of massive stars



Figures from Evans+21 (Cosmic Explorer Book)

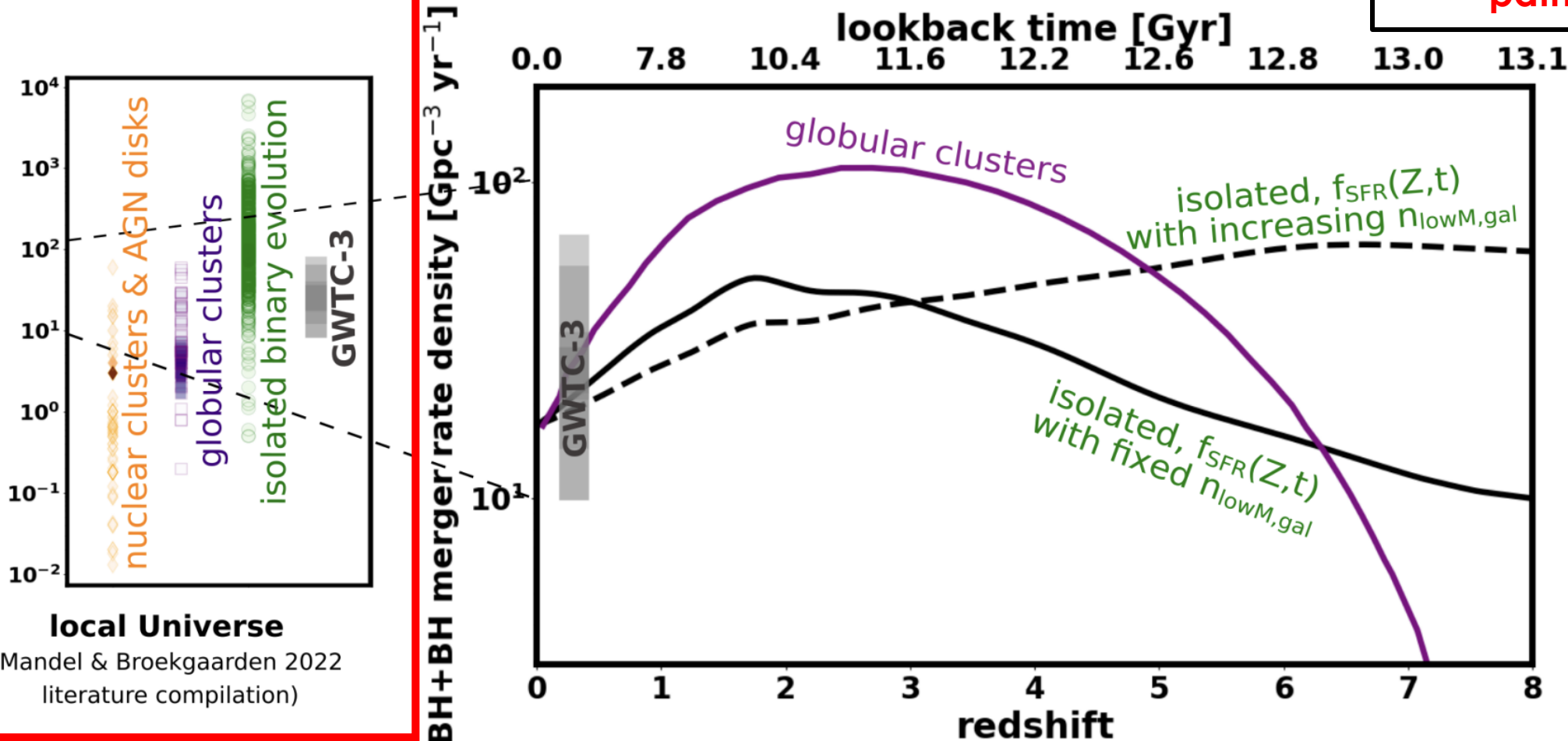
This endeavor & ambition comes with a big *if*:

Cosmic Explorer:
***“a unique instrument to
study massive stars across
cosmic time for \$100/star” if***
**we can connect GW
sources to formation
pathway histories**

This endeavor & ambition comes with a big *if*:

We learned in 2022: Formation Channels are uncertain!
(Rates alone cannot constrain channels)

Cosmic Explorer:
“a unique instrument to
study massive stars across
cosmic time for \$100/star” ***if***
we can connect GW
sources to formation
pathway histories



Gravitational Wave Paleontology today

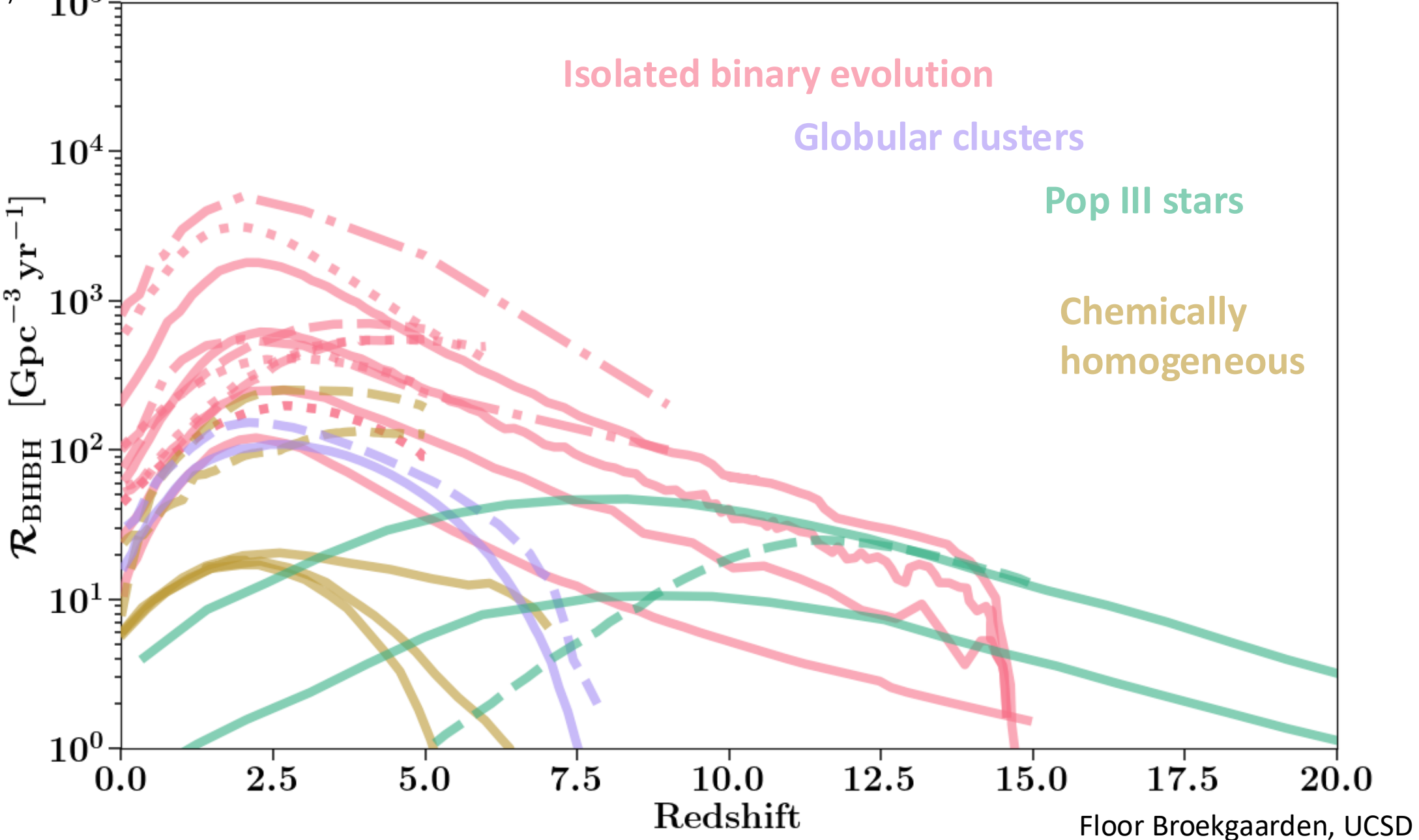


But there's a lot of hope & progress!

Because it will be a game-changer!

Redshift dependent rates

Broekgaarden (in prep)



What causes features in the BBH distributions?

Star formation history artifact (Levina+in prep.)

Chemically Homogeneous Channel? (de Sa in prep.)

Stable mass transfer channel? (van Son+24)

Stable Mass Transfer? (Briel+24)

Globular Clusters?

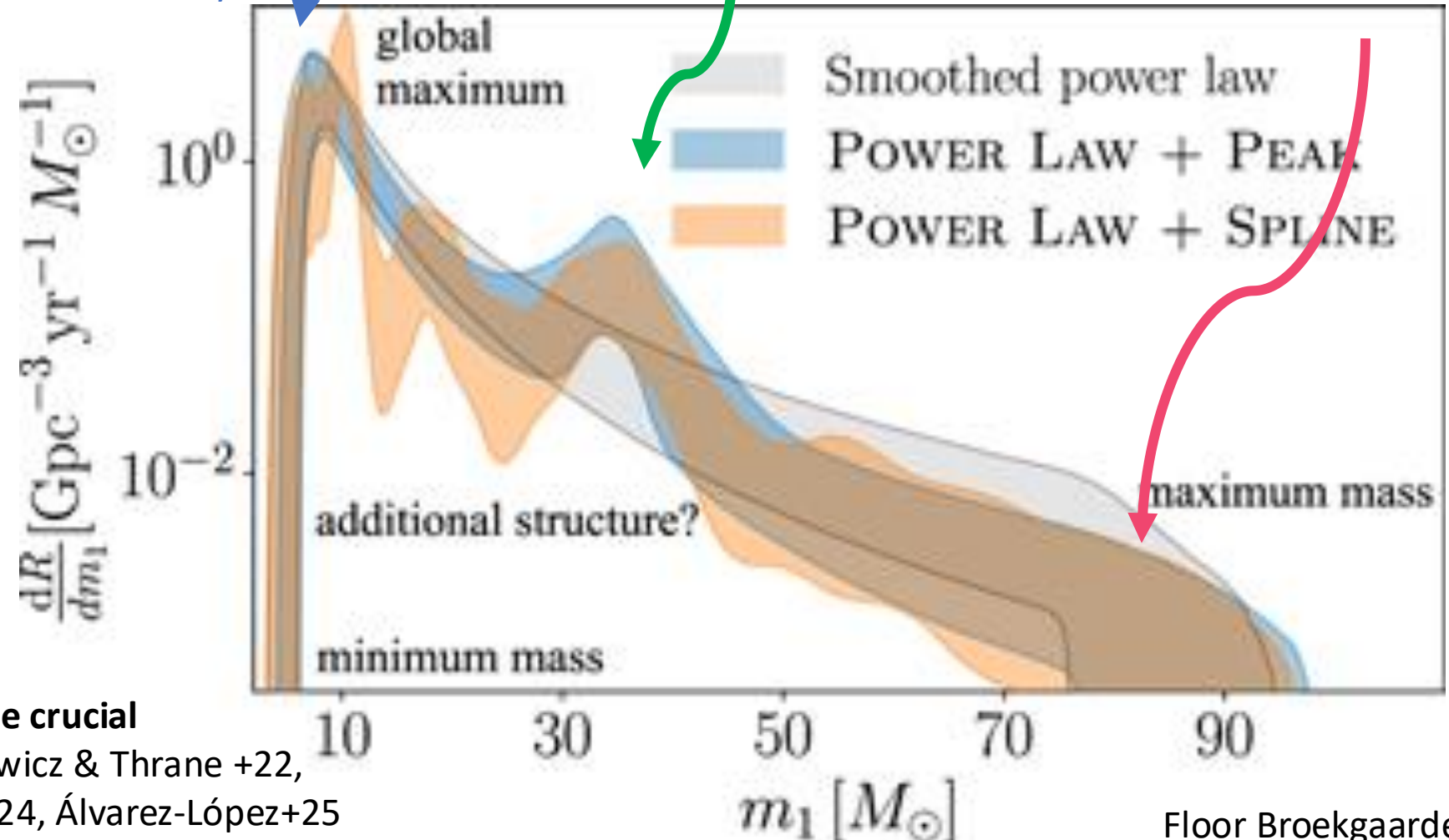
Pair instability Supernova?

Globular Clusters? AGN?

Model uncertainty?

Hierarchical mergers?

See Einstein Telescope & Cosmic Explorer Blue Book & References therein



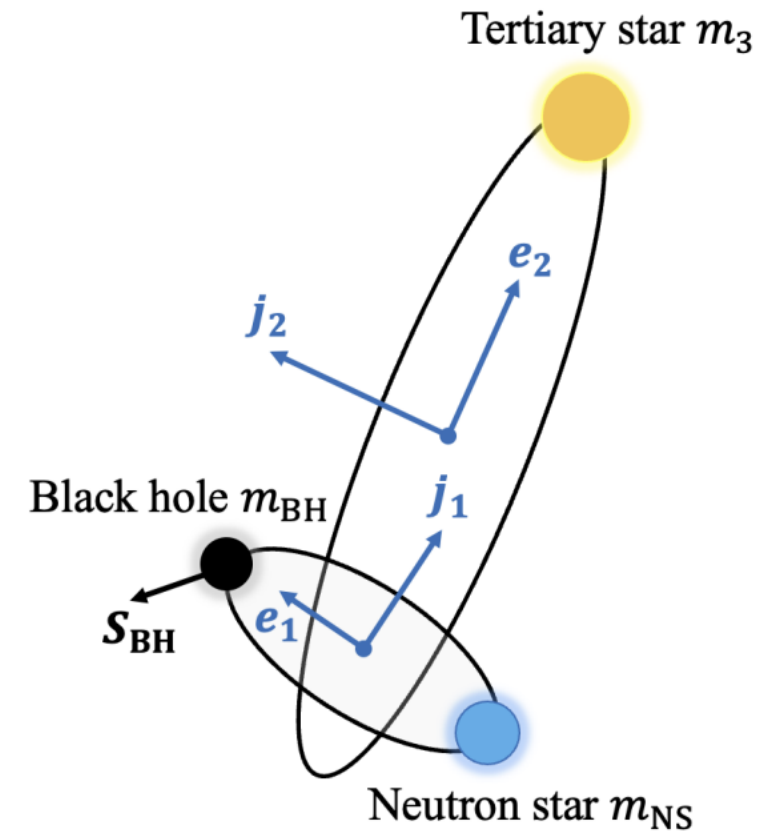
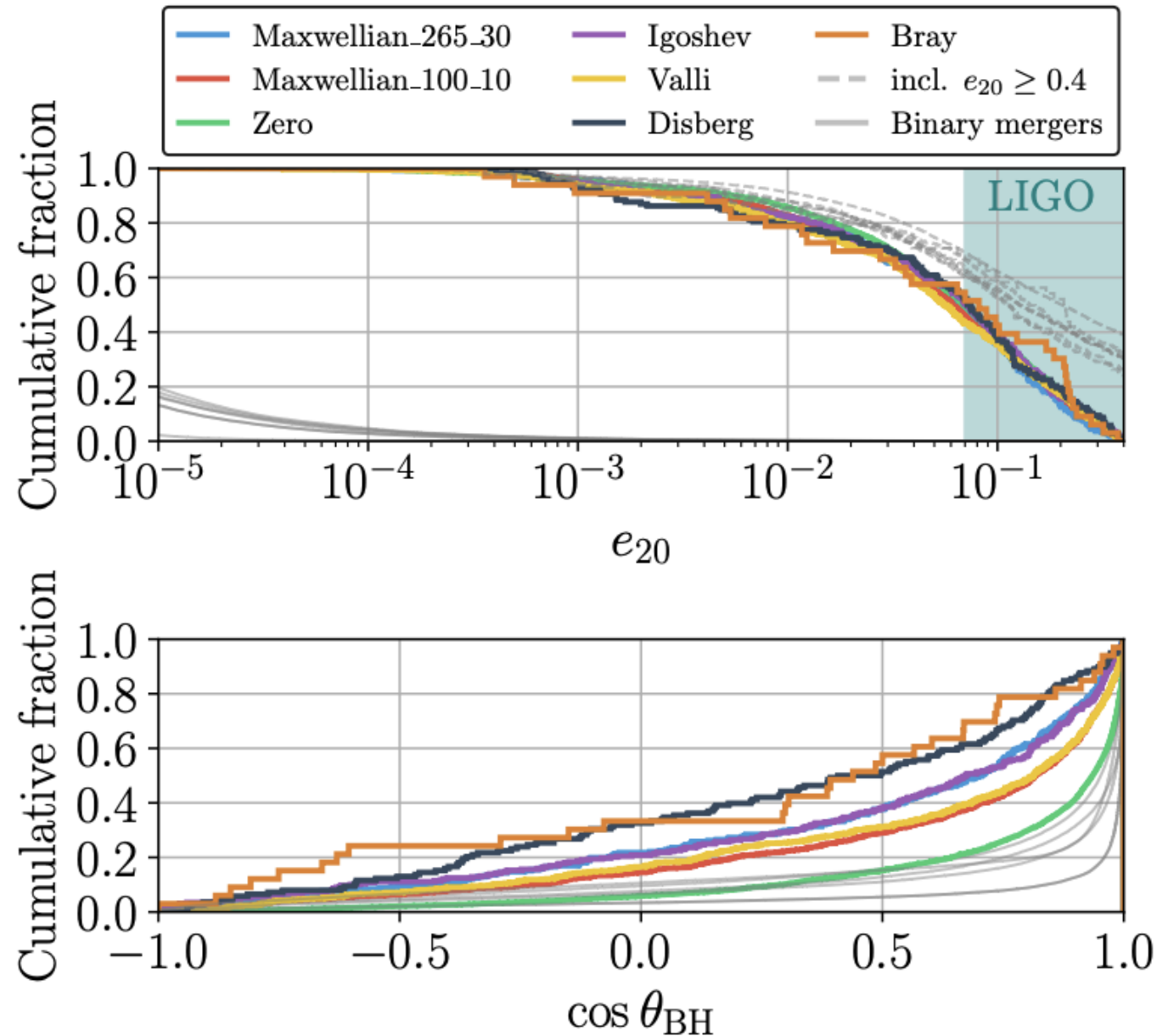
Parameter correlations will be crucial

e.g. Callister+21+24, Adamcewicz & Thrane +22,
Heinzel, Biscoveanu, Vitale +24, Álvarez-López+25

Floor Broekgaarden, UCSD

Triples might explain large BH spin-orbit misalignment observed in NSBH (and produce eccentric events)

Stegman & Klencki 2025



Floor Broekgaarden, UCSD

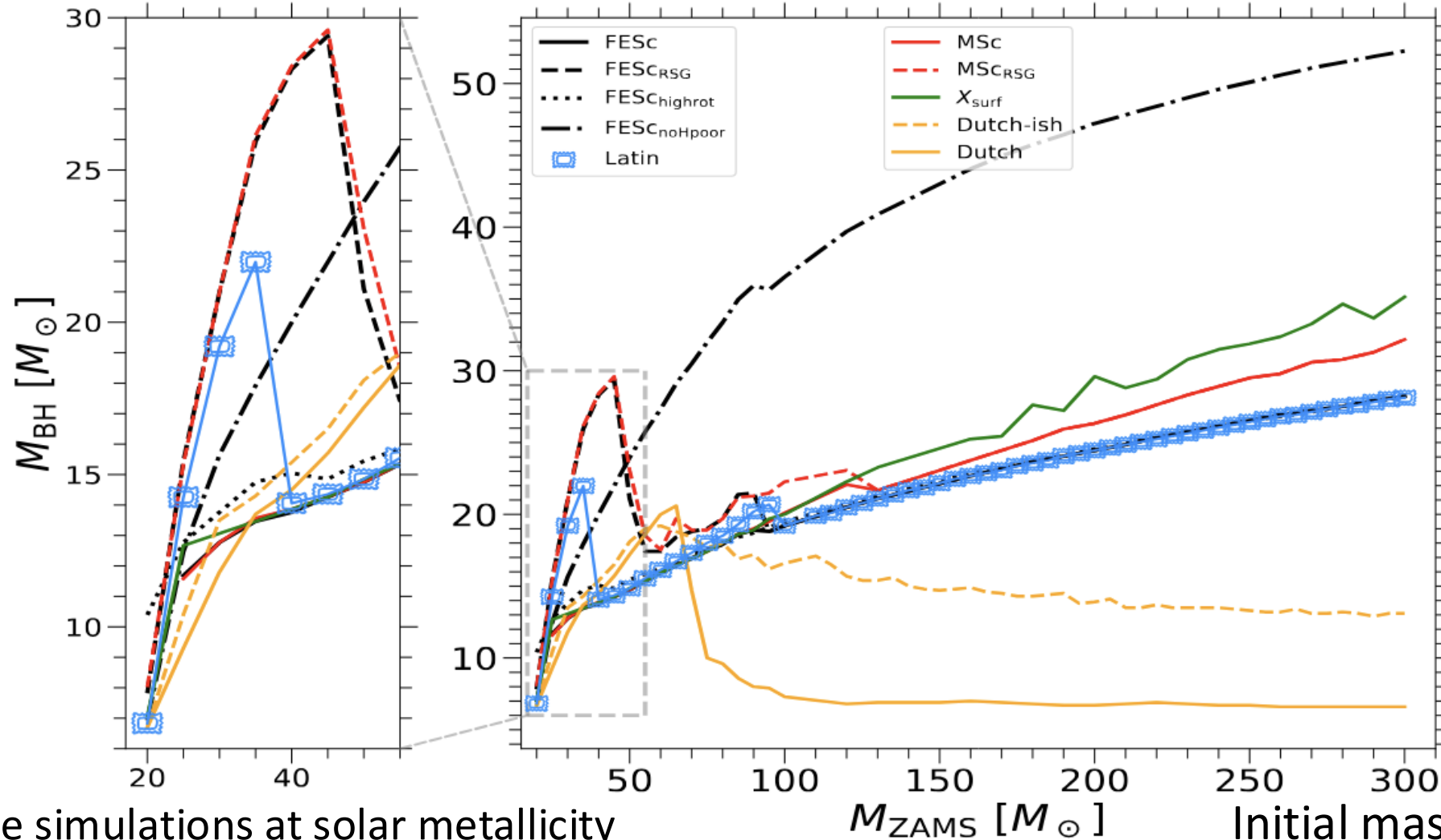
We are improving our models & Physics !

New Supernovae & SN kick insights (e.g. Valli+25, Disberg+25), new/updated models: KRIOS (Tep+25), POSYDON (Fraggoss+24), SEVN (Iorio+24)



Amedeo
Romagnolo

How stellar winds impact BH masses: Romagnolo+25,



These are simulations at solar metallicity

Initial mass stars →

Gravitational Wave Paleontology

Today

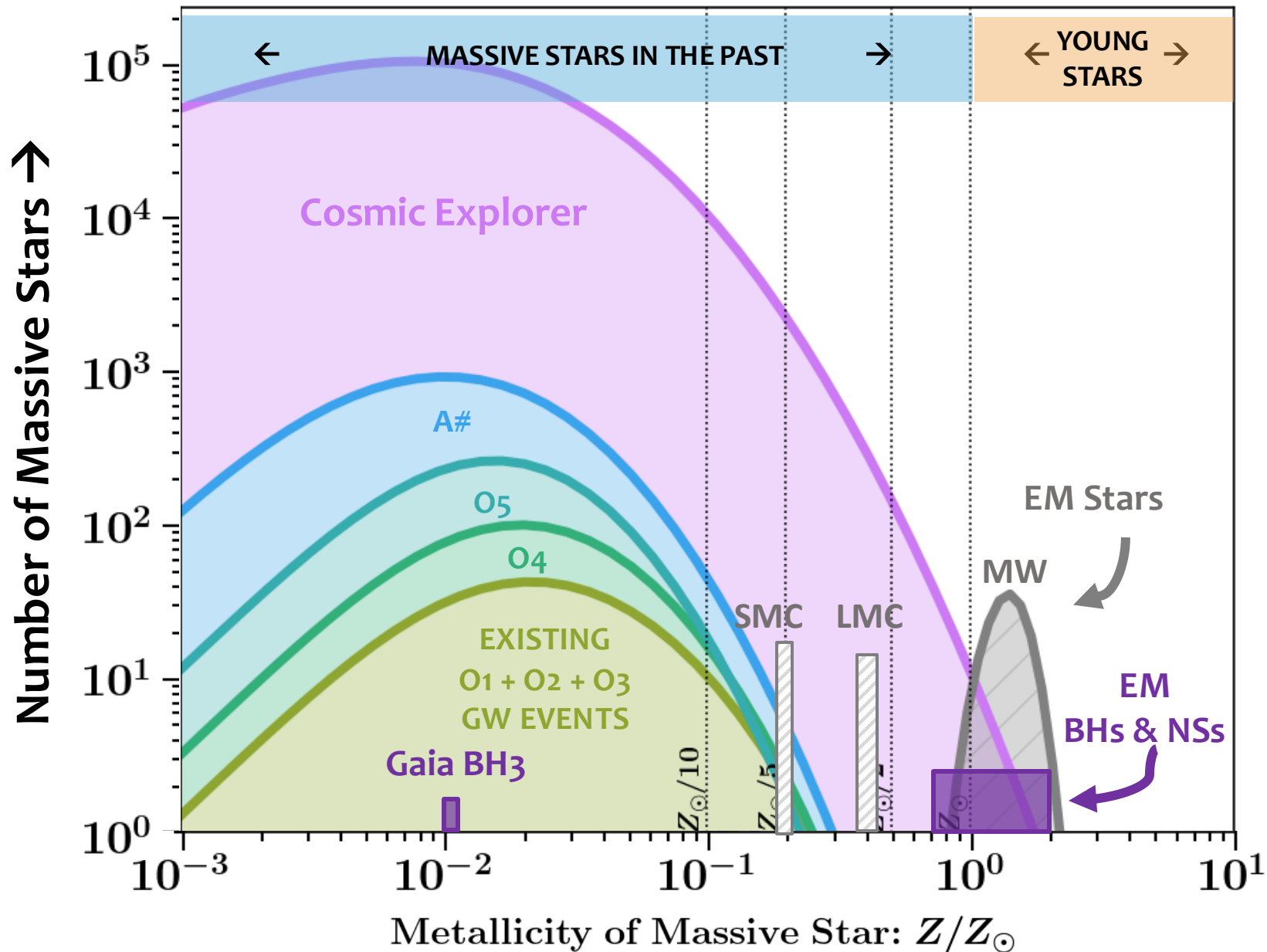


Future



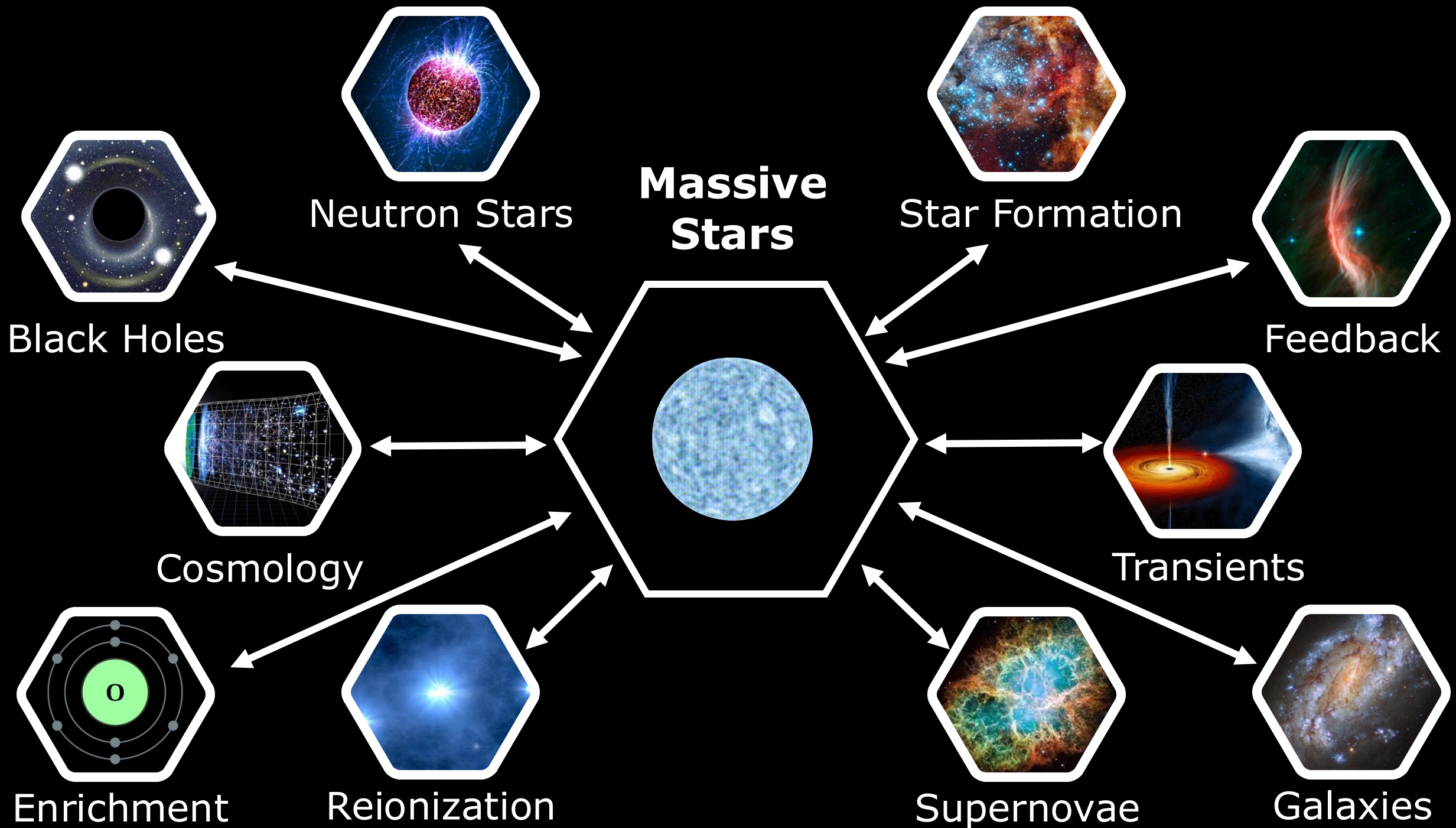
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(billion years) \leftarrow Time in the Past





UC San Diego Black Hole Paleontology Lab

visualizations/data science/software



Floor
Broekgaarden



Sasha Lavina



Amedeo
Romagnolo



Andrea Zonca



Tyler Smith



Tom Wagg



Ana Lam



Kyle Kremer



Kyle Rocha



Lucas de Sa



Marko Ristic



Steffani Grondin



Adam Boesky



Caua Rodrigues



Khushi
Karthikeyan



Jayanth
Bharadwaj



Shasha Arani



Esther Park



Meera Desawale



Laya Binu



Feranmi
Falodun



Aobo Li



Prajakta Saraf



Ann-Kathrin
Schuetz



Alex Migala



Eugene
Shang

YOU?

stellar evolution

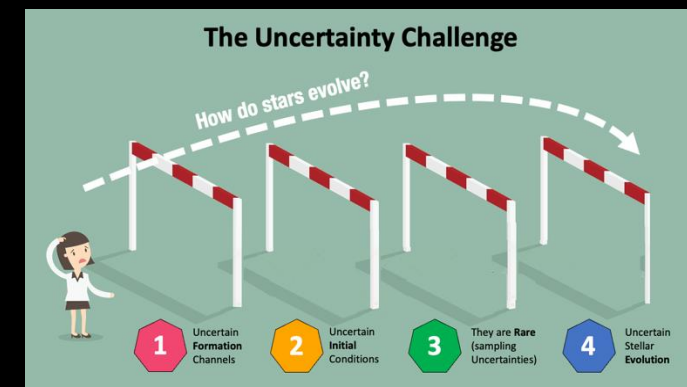
AI/ML Enrichment

Gravitational waves

Astro-statistics

transients

Cosmic star formation history



Other things to check out:

New Astro Resources websites

<https://www.broekgaarden.nl/floor/wordpress> With resources for:

- Applying for PhD/MSc/Postdoc/Fellowships/Faculty etc.
- Recorded workshops on applications, thriving in your PhD/PD etc.

Applying to PhD

Websites that provide overviews/resources.

- Melinda Soares-Furtado, has created a **Discord** for undergraduate students who are applying the coming year(s) to PhD programs and who would benefit from connecting with a digital network of grad school applicants (including current PhD students, postdocs and professor who can answer questions.). You can email her if you want to get an invite to this Discord: mmsoares@wisc.edu.

Additionally:

- A document on “applying for” students and gives tips and advices from former student. See: <https://docs.google.com/document/d/1uSp-sharing> – the department

USA focused:

- Really helpful Workshop/Vid Project: <https://thewocproject.org/>
- USA PhD application timeline on: <https://sites.astro.caltech.edu/>
- NSF GRFP: website with examples

Applying to Postdocs

General overviews demystifying the application process:

- “Guide to postdocs Applications part 1” <https://astrobites.org/2022/08/02/guide-to-postdoc-apps-part1/>, “Guide to postdoc Applications part 2” <https://astrobites.org/2022/08/03/guide-to-postdoc-apps-part2/>. These are great guides that walk you through the main things to think about when applying for postdocs, and provide links to resources for each stage of the postdoc application. *Astrobites: The series written by Mia de los Reyes and Luna Zagorac and edited by Pratik Gandhi, Lina Kimmig, Ishan Mittal, and Jenny Calahan.*

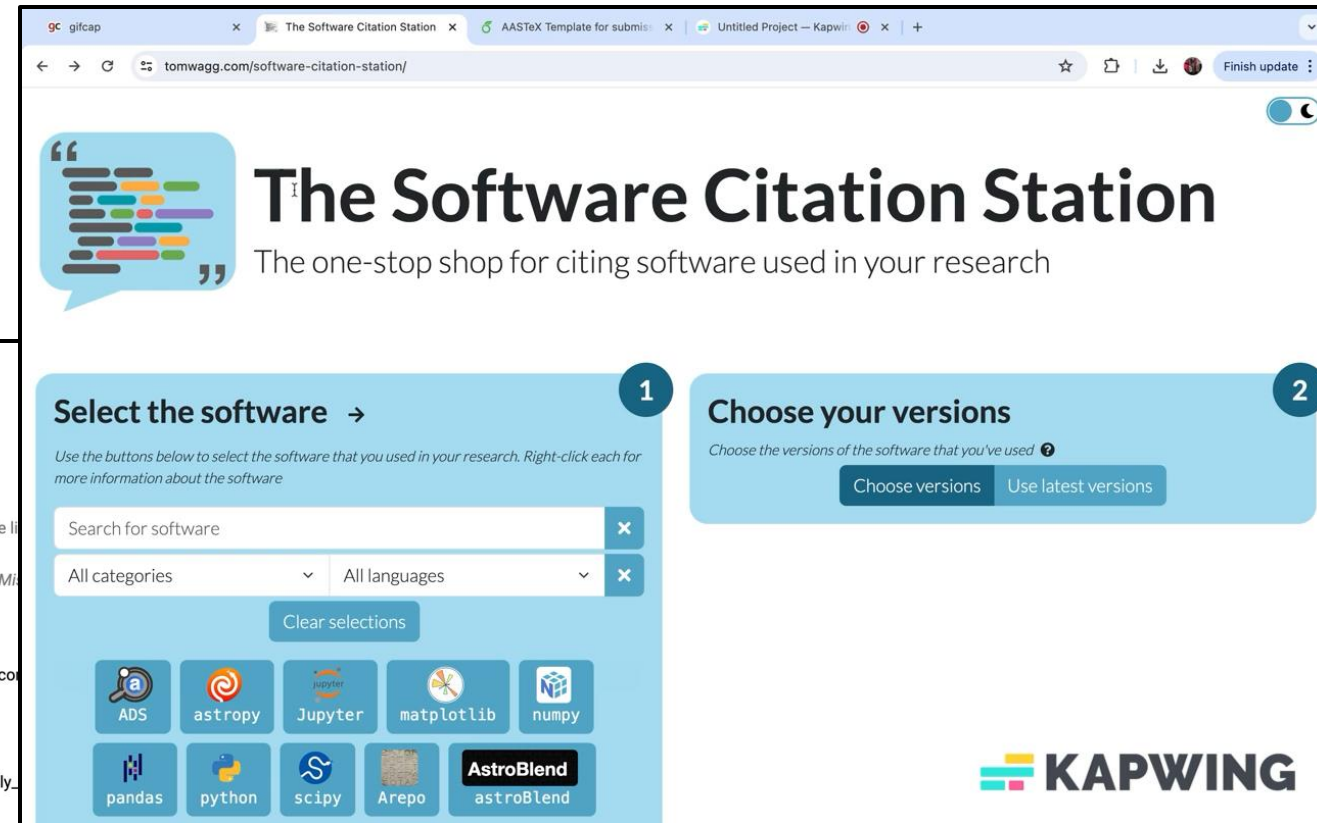
- Workshop presentation on “How to apply for PostDocs (in the USA)” by David Charbonneau: https://www.youtube.com/watch?v=hTc8_5B6zzU&ab_channel=HarvardAstronomyVideo, & slides: [Charbonneau_how_to_apply_to_a_postdoc](#)

More Postdoc application ppt slides and guides:

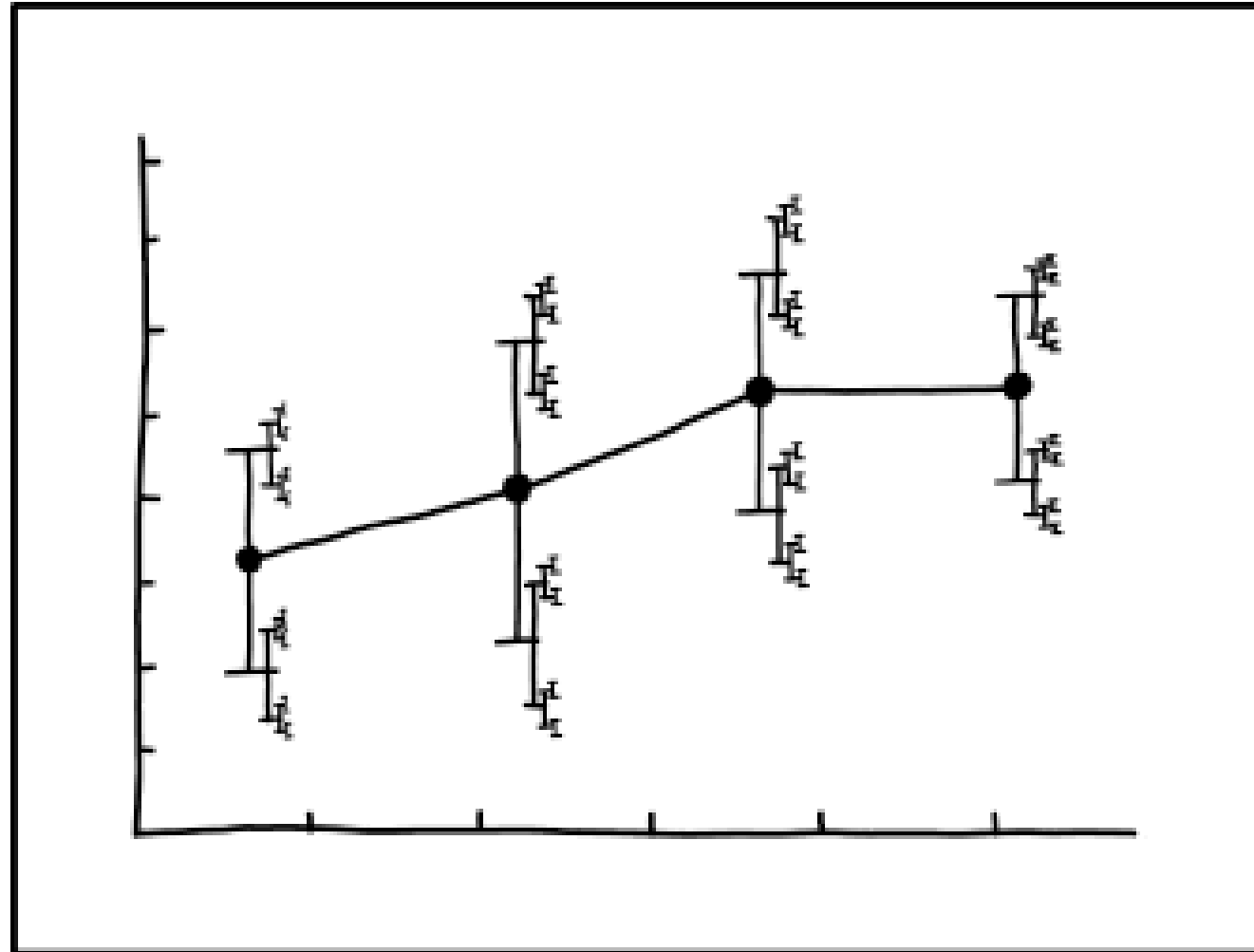
- Great PDF of presentation slides on “how to apply for PhD/PostDocs by Sera Markoff”, with many advices: https://www.dropbox.com/s/bxqtqisiktyaif9/how2apply_bhi_slides.pdf?dl=0
- Great PDF slides of “how to apply for postdocs” by Richard Alexander (U. of Leicester) https://www.astro.le.ac.uk/~rda5/rda5_jobs_talk_2019.pdf and more resources on Richard’s website: https://www.astro.le.ac.uk/~rda5/jobs_2019.html
- Great slides on “How to apply for postdoc positions in Astronomy” by Mary Loli Martinez Aldama (CFT/Poland) and Dhanya Nair (JIVE/NL) <http://astro.udec.cl/titans/postdoc.pdf>
- Guide by Bryan Gaensler & Sarah Maddison on “how to apply for postdocs” https://www.dunlap.utoronto.ca/~bgaensler/papers/gaensler_maddison_jobs_2012.pdf
- the many many helpful career resources by Astrobetter <https://astrobetter.com/wiki/Possible+Career+Paths...> Which includes many blogs and resources for careers outside of academia!! (which is so important!)

The software citation station

Tom Wagg & FSB (2024): to make citing software incredibly easy



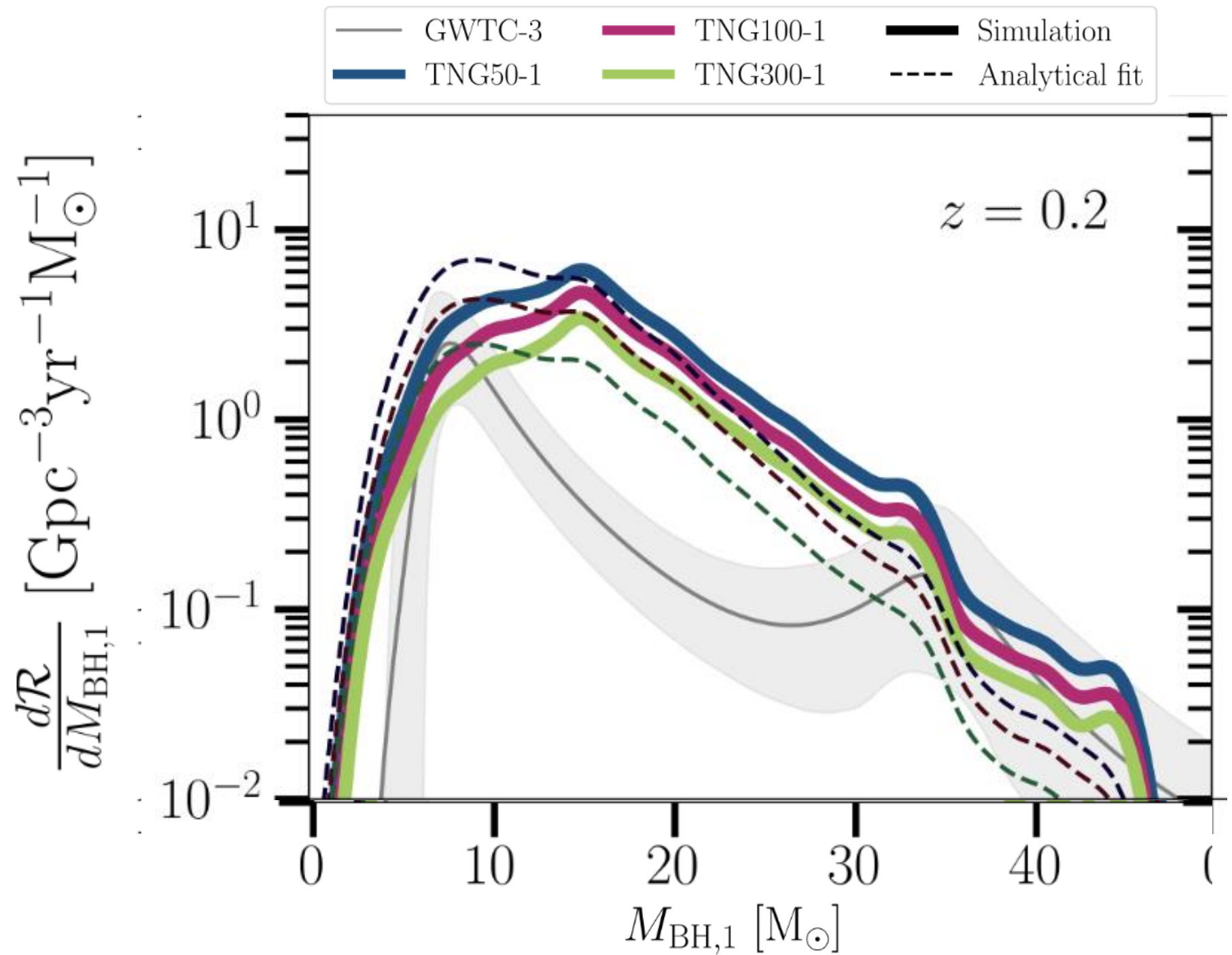
With so many uncertainties, can we learn anything from Gravitational Waves?



I DON'T KNOW HOW TO PROPAGATE
ERROR CORRECTLY, SO I JUST PUT
ERROR BARS ON ALL MY ERROR BARS.



Sasha Levina (2025a,b; in prep)



Do the majority of BHBH mergers experience a Common Envelope phase??

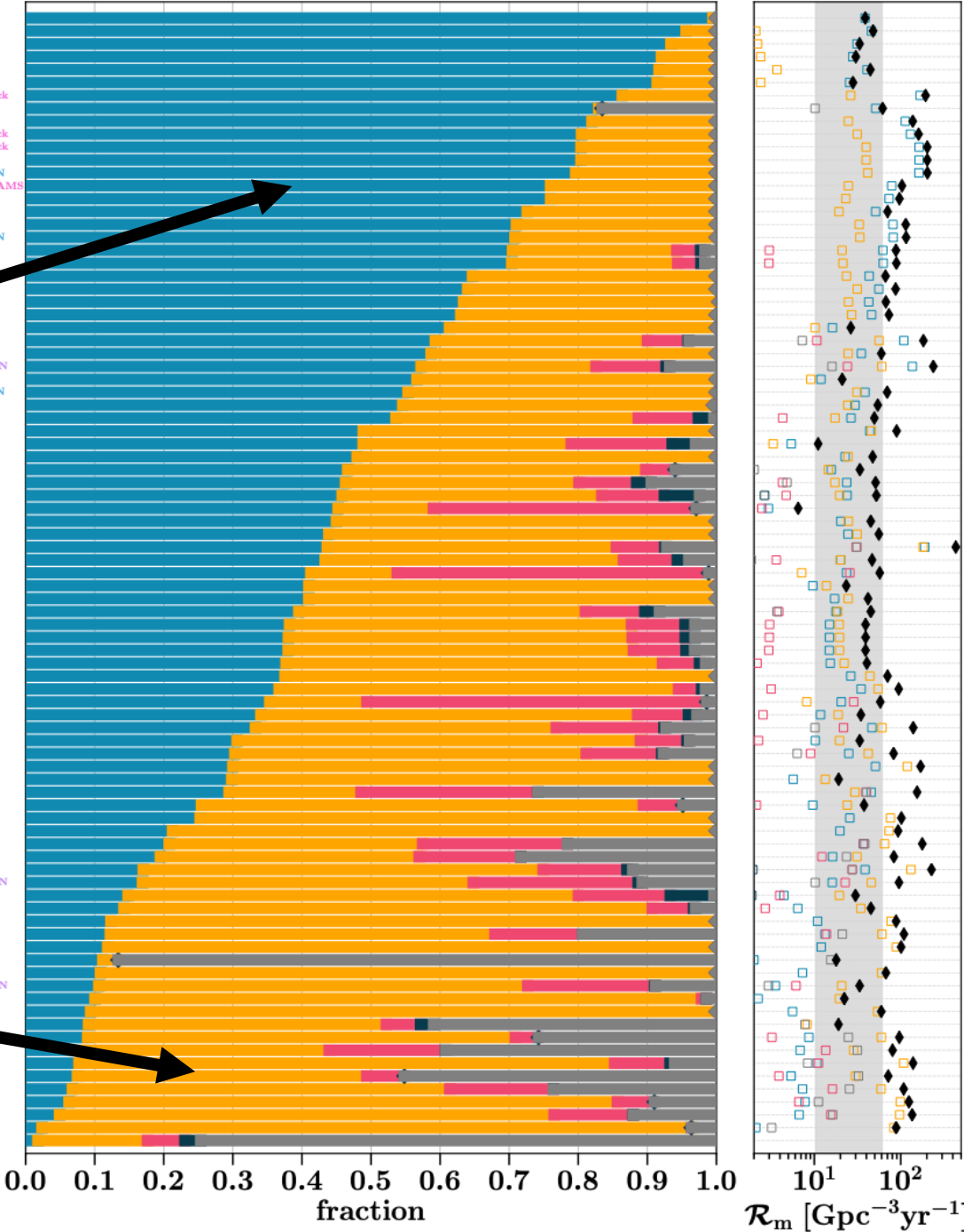
Yes!

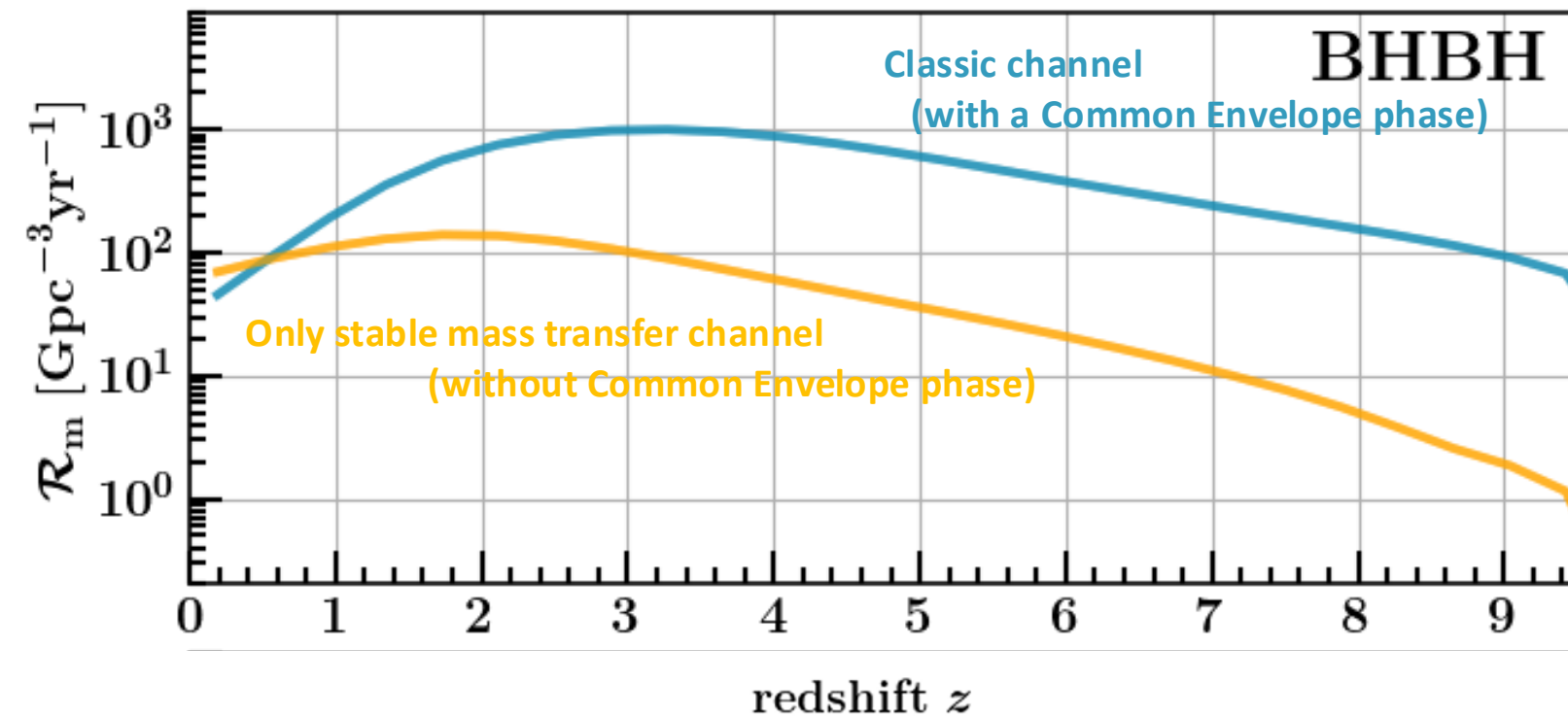
different codes and uncertainties

No!

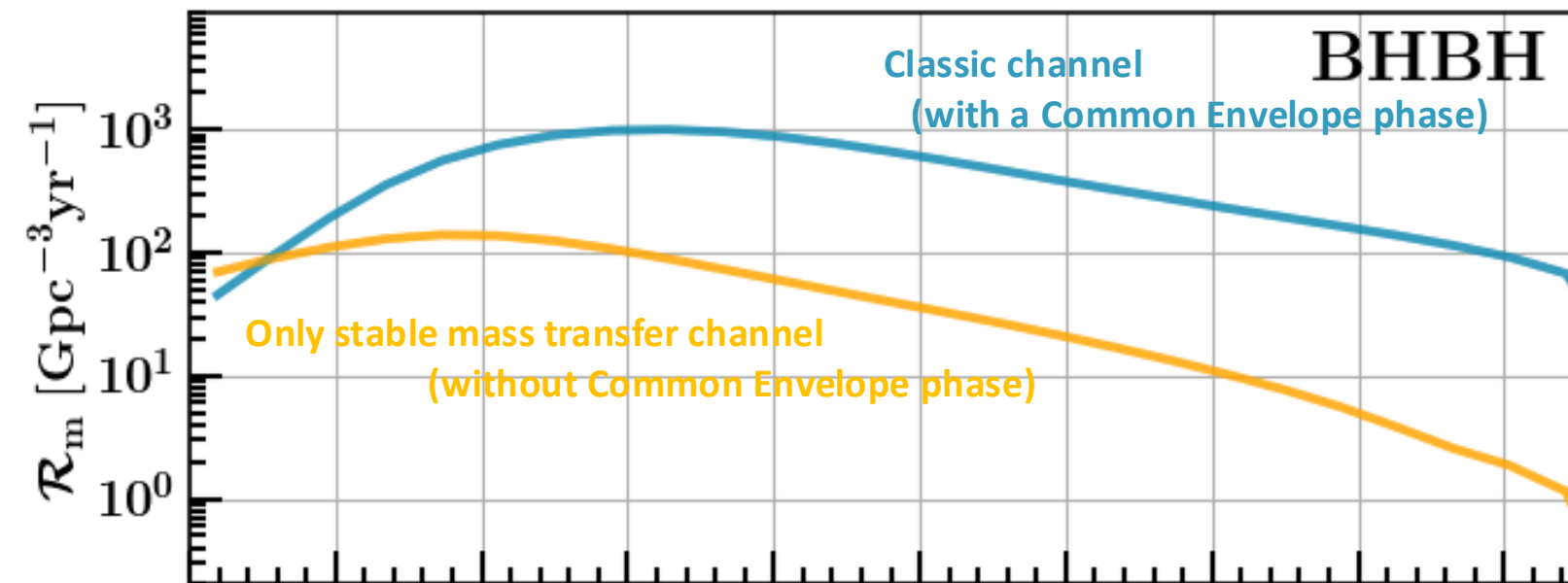
Broekgaarden+ (in prep.)

name	code	tracks	var1	var2
BA21	POSYDON	MESA	$10^4 \times \rho_{\text{He}}$	
DT22	SeBa	SSE	$\beta = 0.3$	
DT22	SeBa	SSE	$\beta = 0.3$	
DT22	SeBa	SSE	$\beta = 0.7$	
BA21	POSYDON	MESA	$10^6 \times \rho_{\text{He}}$	
DT22	SeBa	SSE	$\beta = 0.7$	
RG21	POSYDON	MESA	delayed SN	no BH kick
OL21	StarTrack	SSE	fiducial	
BA21	POSYDON	MESA	$\alpha = 0.2$	
RG21	POSYDON	MESA	rapid SN	no BH kick
RG21	POSYDON	MESA	N20 SN	no BH kick
RG21	POSYDON	MESA	N20 SN	
RG21	POSYDON	MESA	N20 SN	Full ECSN
BA21	POSYDON	MESA	logU prior	no MT ZAMS
BA21	POSYDON	MESA	logU prior	
RG21	POSYDON	MESA	delayed SN	
RG21	POSYDON	MESA	rapid SN	
RG21	POSYDON	MESA	rapid SN	Full ECSN
BRO22	COMPAS	SSE	optCE	
BRO22	COMPAS	SSE	UcaseBB	optCE
BA21	POSYDON	MESA	$10^3 \times \rho_{\text{He}}$	
BA21	POSYDON	MESA	no MT ZAMS	
BA21	POSYDON	MESA	$\alpha = 0.5$	
s822	COSMIC	SSE	$\sigma = 30$	
DT22	SeBa	SSE	$\beta = 0.7$	
BO24	COMPAS	SSE	$\sigma = 30$	rapid SN
BA21	POSYDON	MESA	$\alpha = 2$	
DT22	SeBa	SSE	$\sigma = 30$	delayed SN
DT22	SeBa	SSE	$\beta = 0.7$	IT
RG21	POSYDON	MESA	delayed SN	Full ECSN
BA21	POSYDON	MESA	$\alpha = 0.75$	
BRO22	COMPAS	SSE	no BH kick	
DT22	SeBa	SSE	$\beta = 0.3$	
BRO22	COMPAS	SSE	$f_{\text{WR}} = 5$	
BA21	POSYDON	MESA	$\alpha = 5$	
BRO22	COMPAS	SSE	$\beta = 0.25$	
BRO22	COMPAS	SSE	$\sigma = 30$	
BRO22	COMPAS	SSE	$\alpha = 2$	
BRI22	BPASS	STARS	fiducial	
BA21	POSYDON	MESA	$\alpha = 0.5$	
BA21	POSYDON	MESA	Belcz q_{crit}	
BO24	COMPAS	SSE	$\sigma = 30$	MM SN
BRO22	COMPAS	SSE	$m_{\text{NS}} \leq 2 M_{\odot}$	
BRI22	BPASS	STARS	rapid SN	
DT22	SeBa	SSE	$\beta = 0.3$	
BA21	POSYDON	MESA	$\alpha = 0.35$	
BRO22	COMPAS	SSE	$\sigma = 100$	
BRO22	COMPAS	SSE	fiducial	
BRO22	COMPAS	SSE	no PISN	
BRO22	COMPAS	SSE	UcaseBB	
BRO22	COMPAS	SSE	rapid SN	
DT22	SeBa	SSE	$\beta = 0.3$	
BRO22	COMPAS	SSE	$f_{\text{WR}} = 0.1$	
BRI22	BPASS	STARS	delayed SN	
BRO22	COMPAS	SSE	$m_{\text{NS}} \leq 3 M_{\odot}$	
BO24	COMPAS	SSE	$\sigma = 265$	rapid SN
BRO22	COMPAS	SSE	$\alpha = 0.5$	
BO24	COMPAS	SSE	$\sigma = 750$	rapid SN
BA21	POSYDON	MESA	Clayton q_{crit}	
DT22	SeBa	SSE	$\beta = 0.3$	IT
BO24	COMPAS	SSE	$\alpha = 2$	$\beta = 0.25$
BRO22	COMPAS	SSE	$\beta = 0.5$	
DT22	SeBa	SSE	$\beta = 0.7$	
DT22	SeBa	SSE	$\beta = 0.7$	
BO24	COMPAS	SSE	$\alpha = 2$	$\beta = 0.5$
BO24	COMPAS	SSE	$\alpha = 10$	$\beta = 0.25$
BO24	COMPAS	SSE	$\alpha = 2$	$\beta = 0.75$
BO24	COMPAS	SSE	$\sigma = 265$	delayed SN
BRO22	COMPAS	SSE	$\alpha = 10$	
BRO22	COMPAS	SSE	$\beta = 0.75$	
DT22	SeBa	SSE	$\beta = 0.7$	$\zeta = 7.5$
BO24	COMPAS	SSE	$\alpha = 10$	$\beta = 0.5$
DT22	SeBa	SSE	$\beta = 0.7$	$\zeta = 7.5$
OL21	StarTrack	SSE	changed CEswitch	
DT22	SeBa	SSE	$\beta = 0.3$	$\zeta = 7.5$
BO24	COMPAS	SSE	$\sigma = 750$	delayed SN
BRO22	COMPAS	SSE	$\alpha = 0.1$	
DT22	SeBa	SSE	$\beta = 0.3$	$\zeta = 7.5$
BO24	COMPAS	SSE	$\sigma = 750$	MM SN
BO24	COMPAS	SSE	$\alpha = 0.1$	$\beta = 0.25$
BO24	COMPAS	SSE	$\alpha = 0.5$	$\beta = 0.25$
BO24	COMPAS	SSE	$\alpha = 10$	$\beta = 0.75$
BO24	COMPAS	SSE	$\alpha = 0.1$	$\beta = 0.25$
BO24	COMPAS	SSE	$\alpha = 0.5$	$\beta = 0.5$
BO24	COMPAS	SSE	$\alpha = 0.1$	$\beta = 0.75$
BO24	COMPAS	SSE	$\alpha = 0.5$	$\beta = 0.75$
OL21	StarTrack	SSE	changed CE	
BO24	COMPAS	SSE	$\sigma = 750$	MM SN

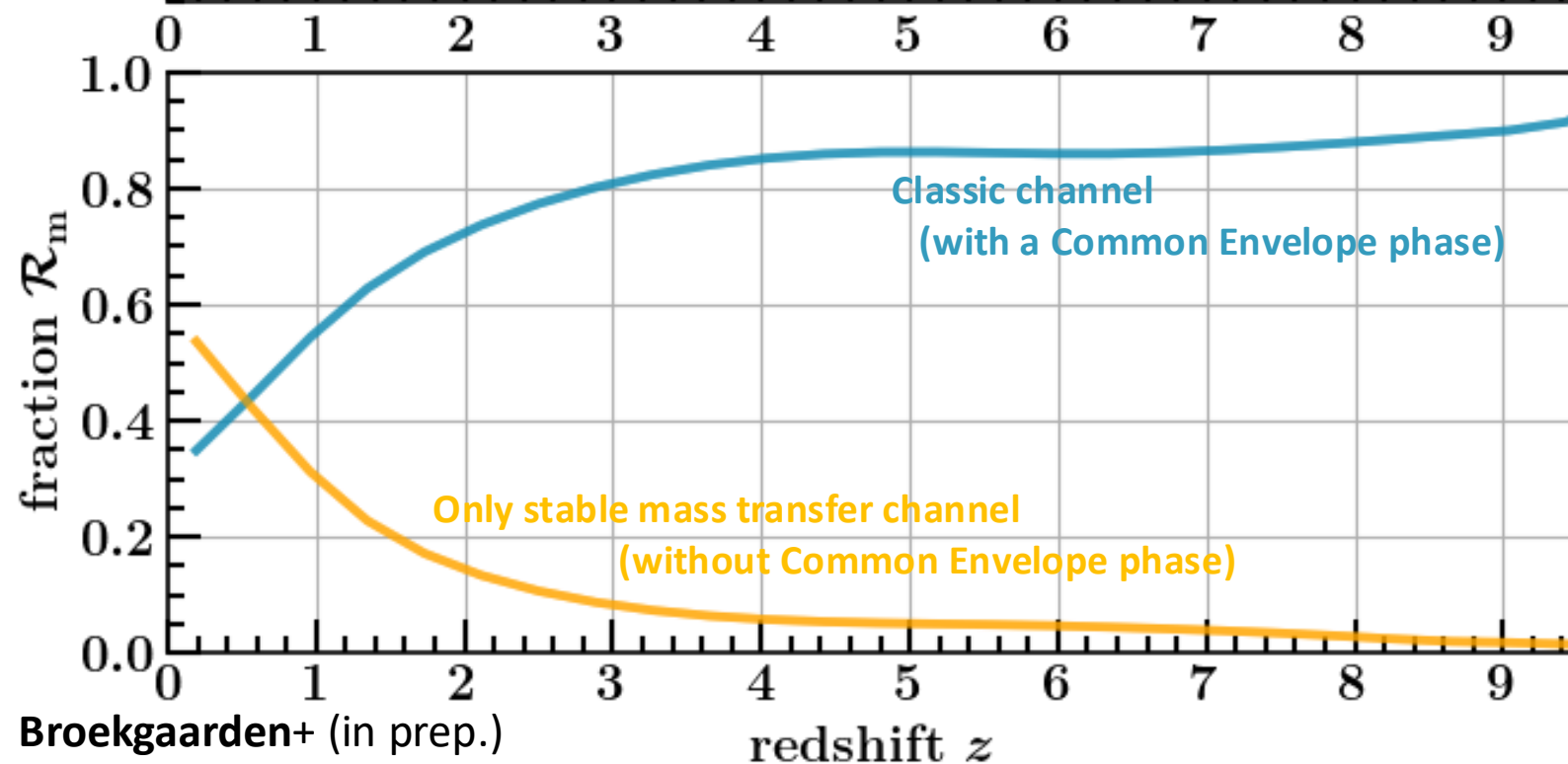




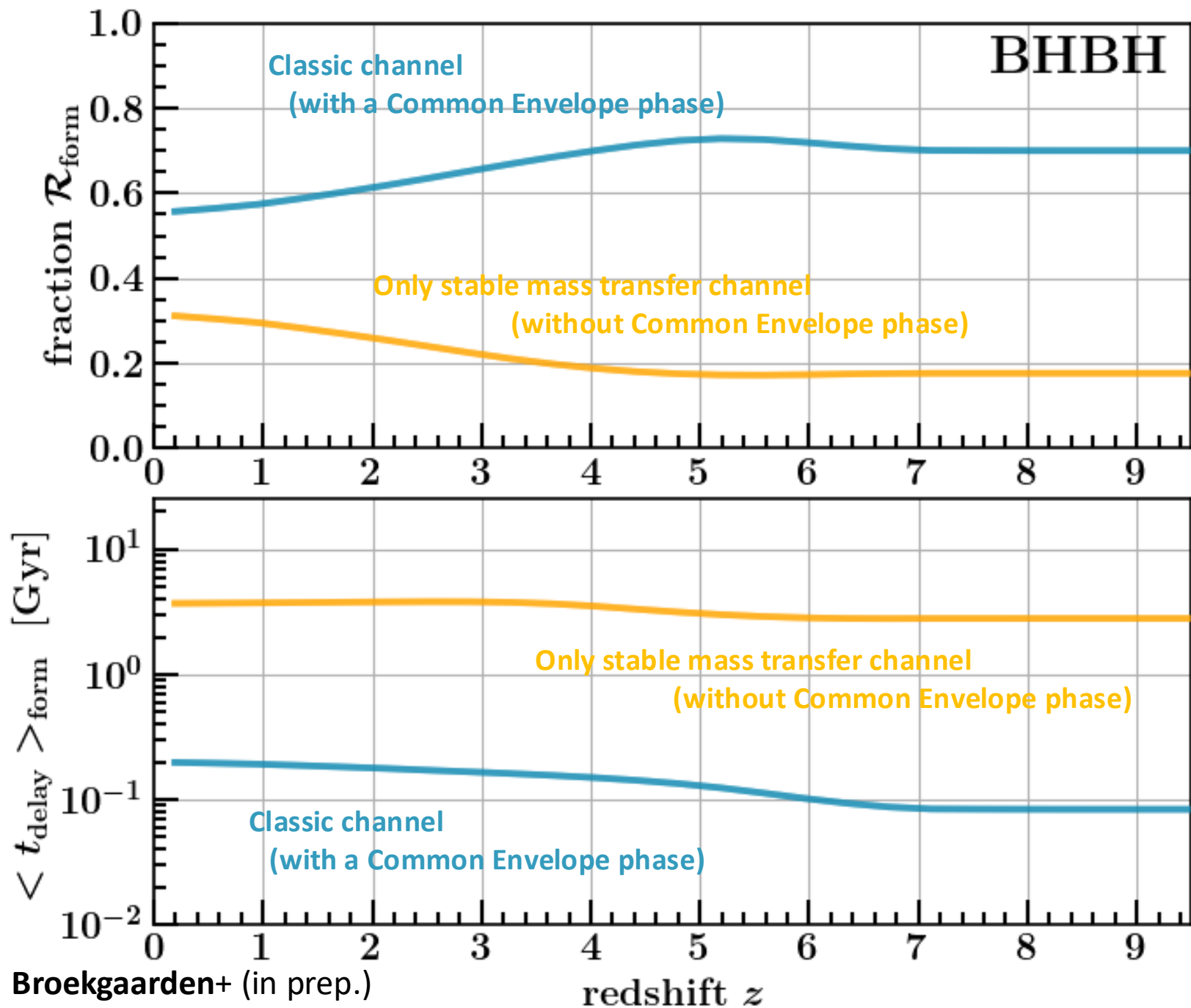
Merger Rate
without or **with** a
Common Envelope phase



Merger Rate
without or **with** a
Common Envelope phase

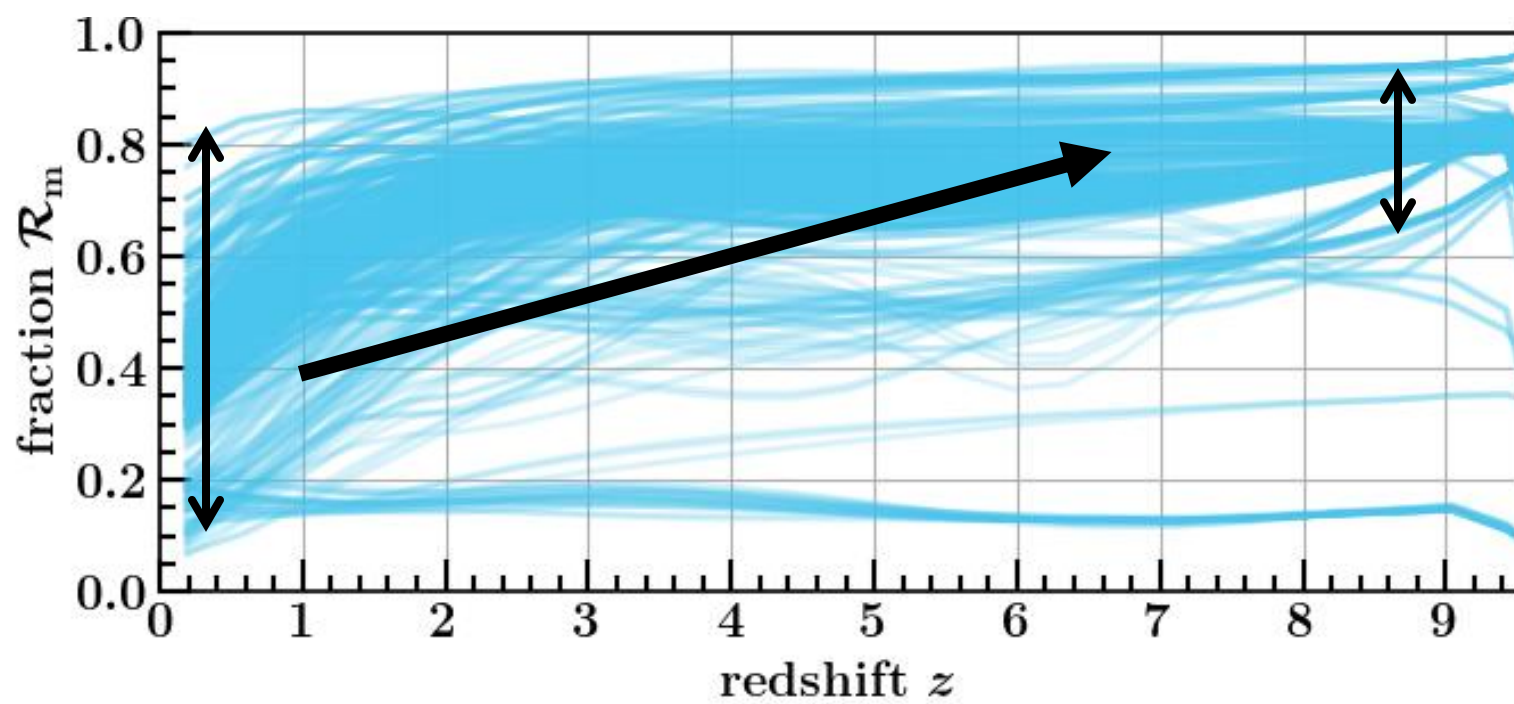


Merger Fraction
without or **with** a
Common Envelope phase

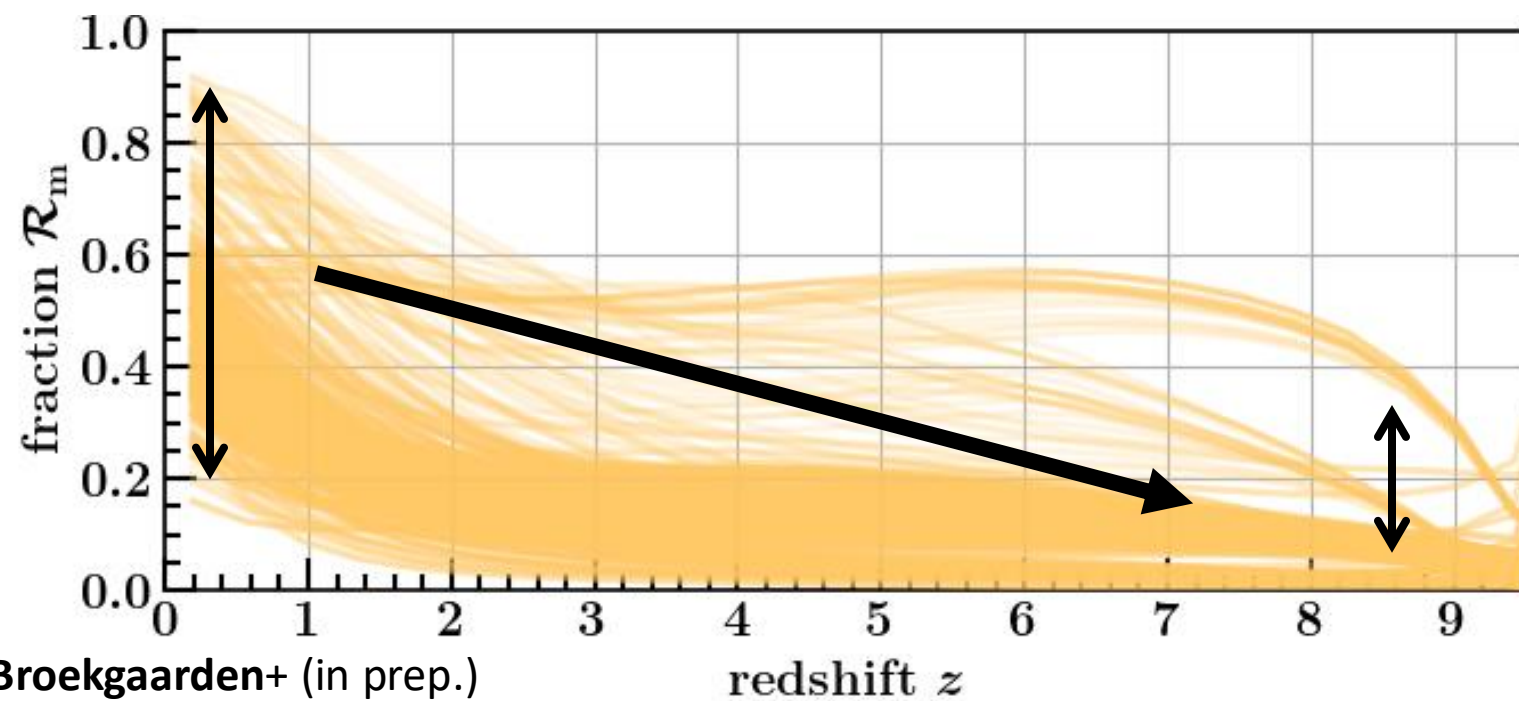


Formation Rate
without or **with** a
Common Envelope phase

Formation Inspiral time
without or **with** a
Common Envelope phase



Merger Rate
without or **with** a
Common Envelope phase



5

Constraints from intermediate stages



Ana Lam

