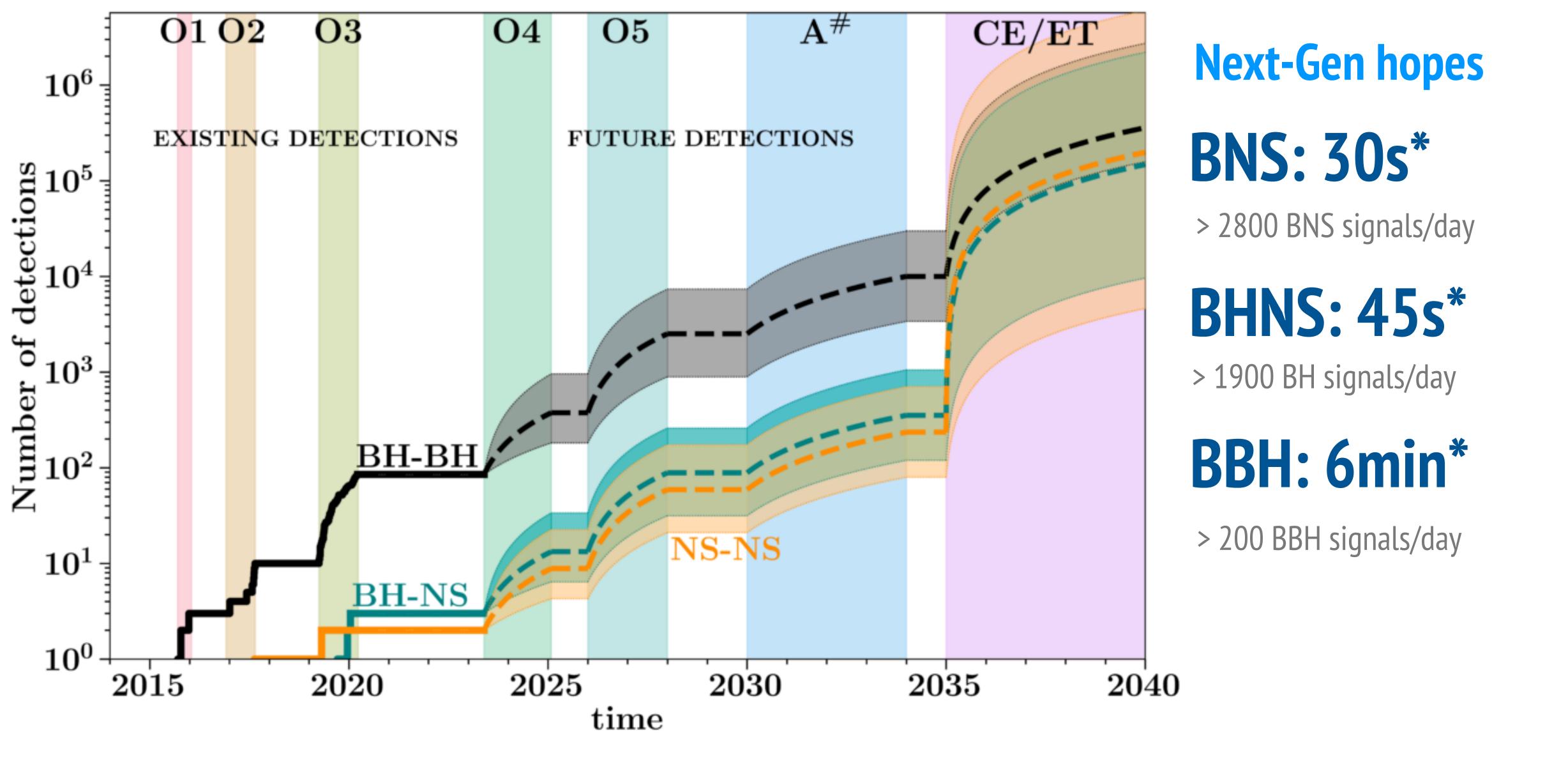
Too Many, Too Long, Too Loud

CBC Data Analysis Challenges in XG era

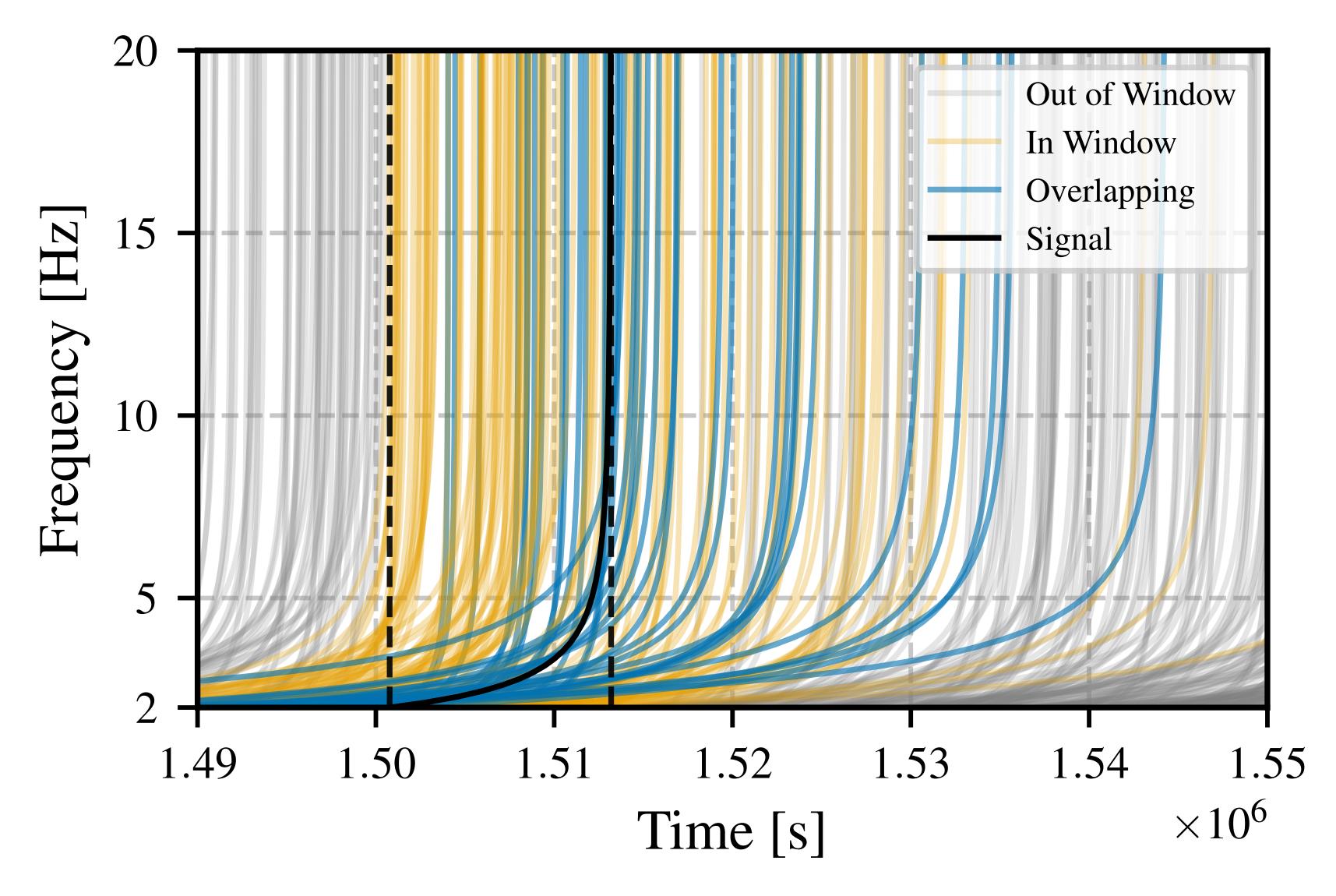
Koustav Chandra

2nd July 2025 | PAX X



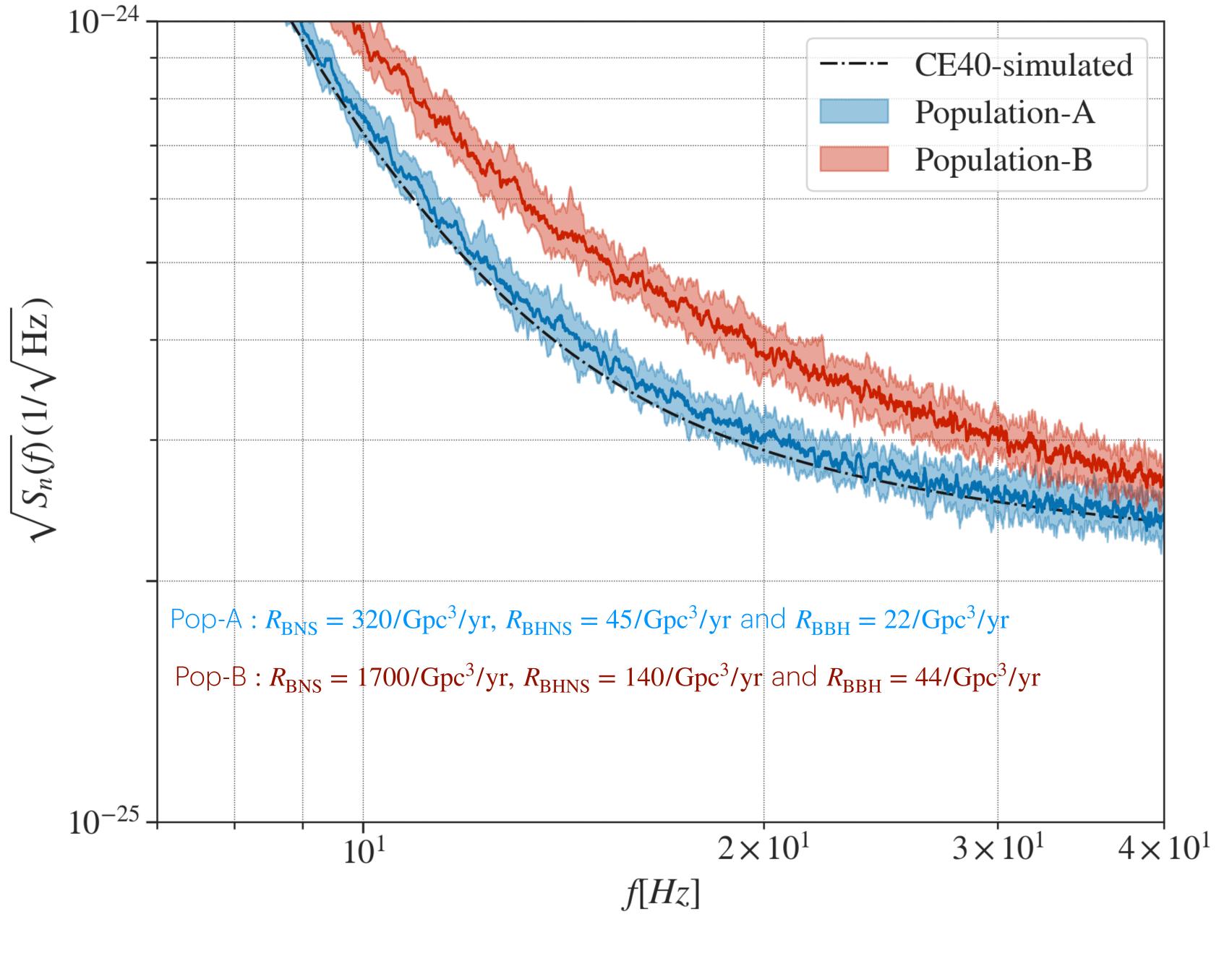


^{*}exact number depending on uncertainties in the binary merger rate and distribution function



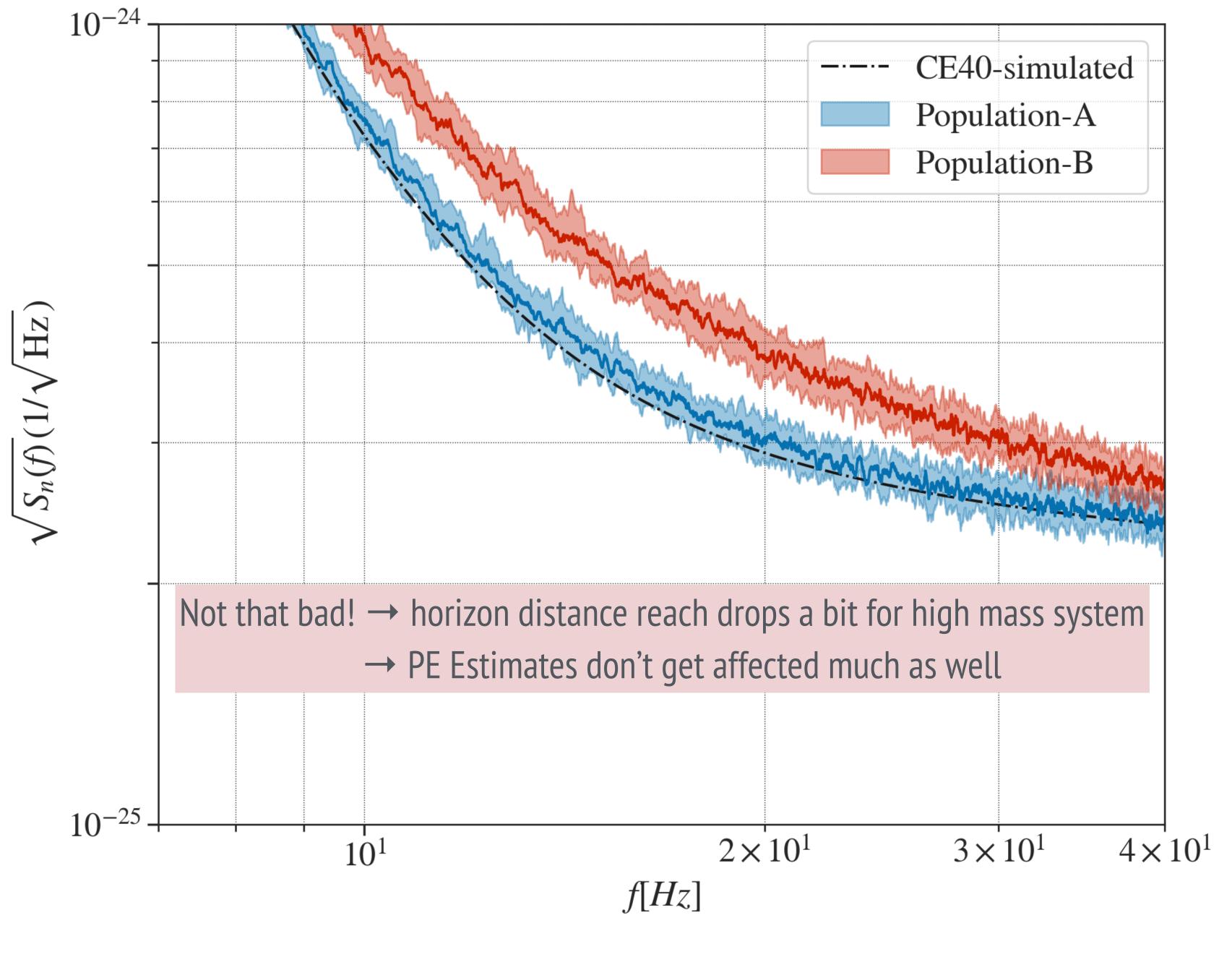
Foreground noise

- Long-lived and frequent signals →
 Overlapping in-band signals
 - At any given time, multiple signals may overlap in the detector band, introducing non-stationary features in the data.
- Low-frequency signal strain ≠ widesense stationary Gaussian noise
- This impacts accurate estimation of the data power spectrum
- But how bad it is?



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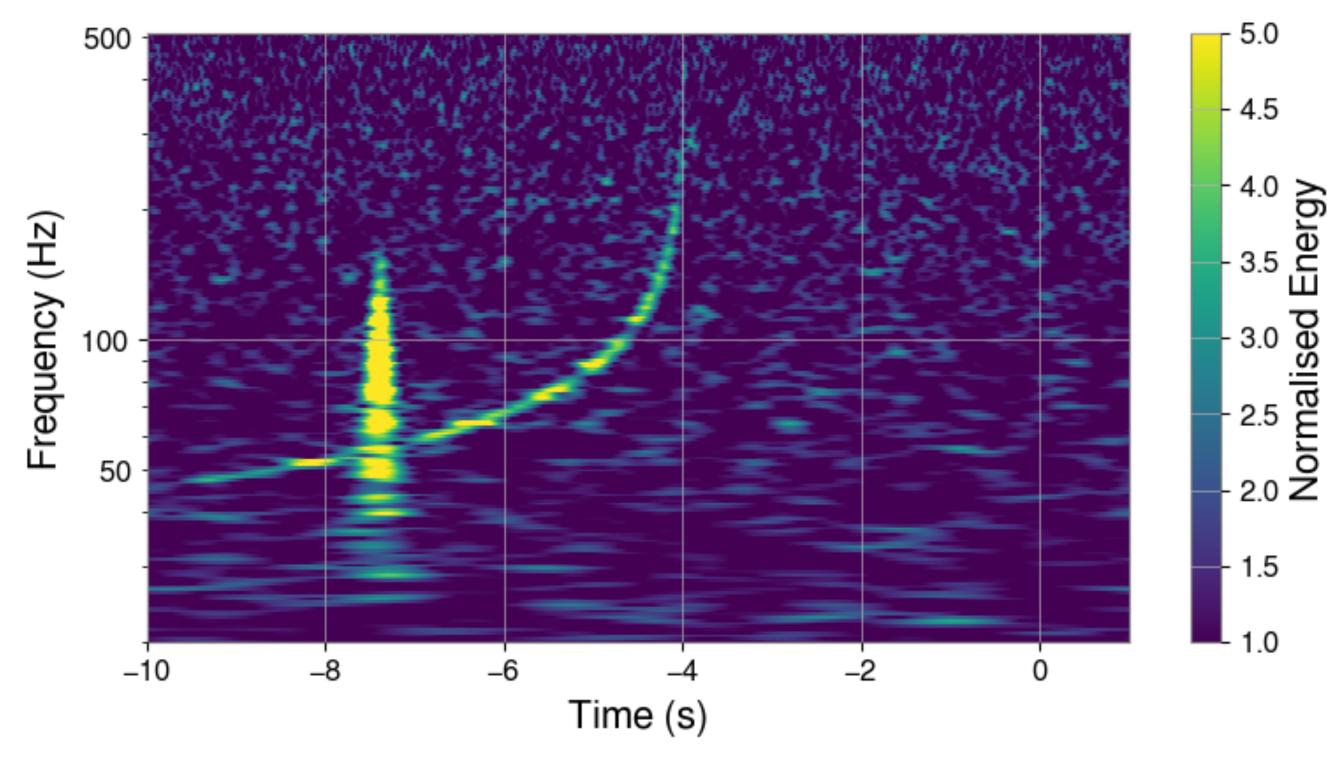


Foreground noise

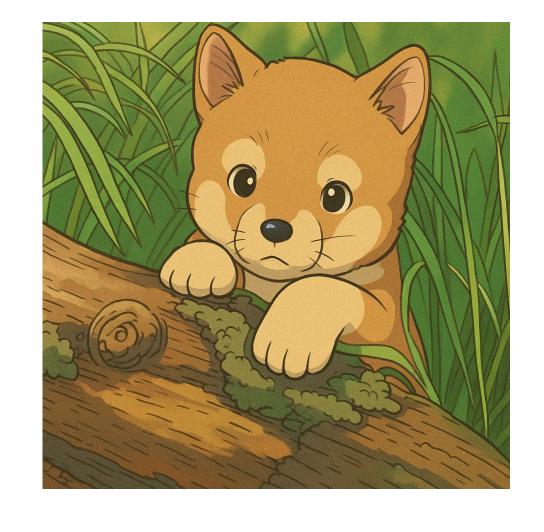
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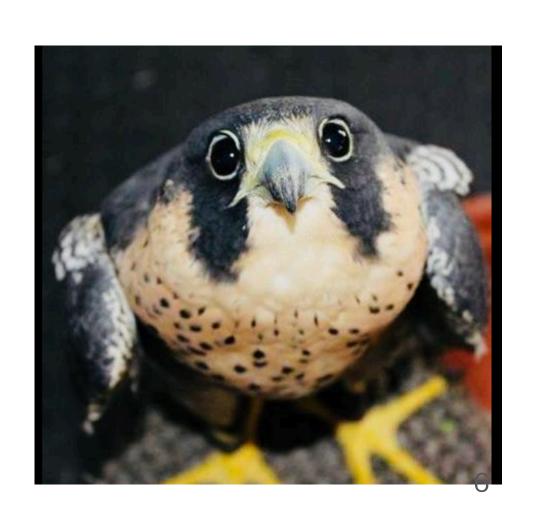


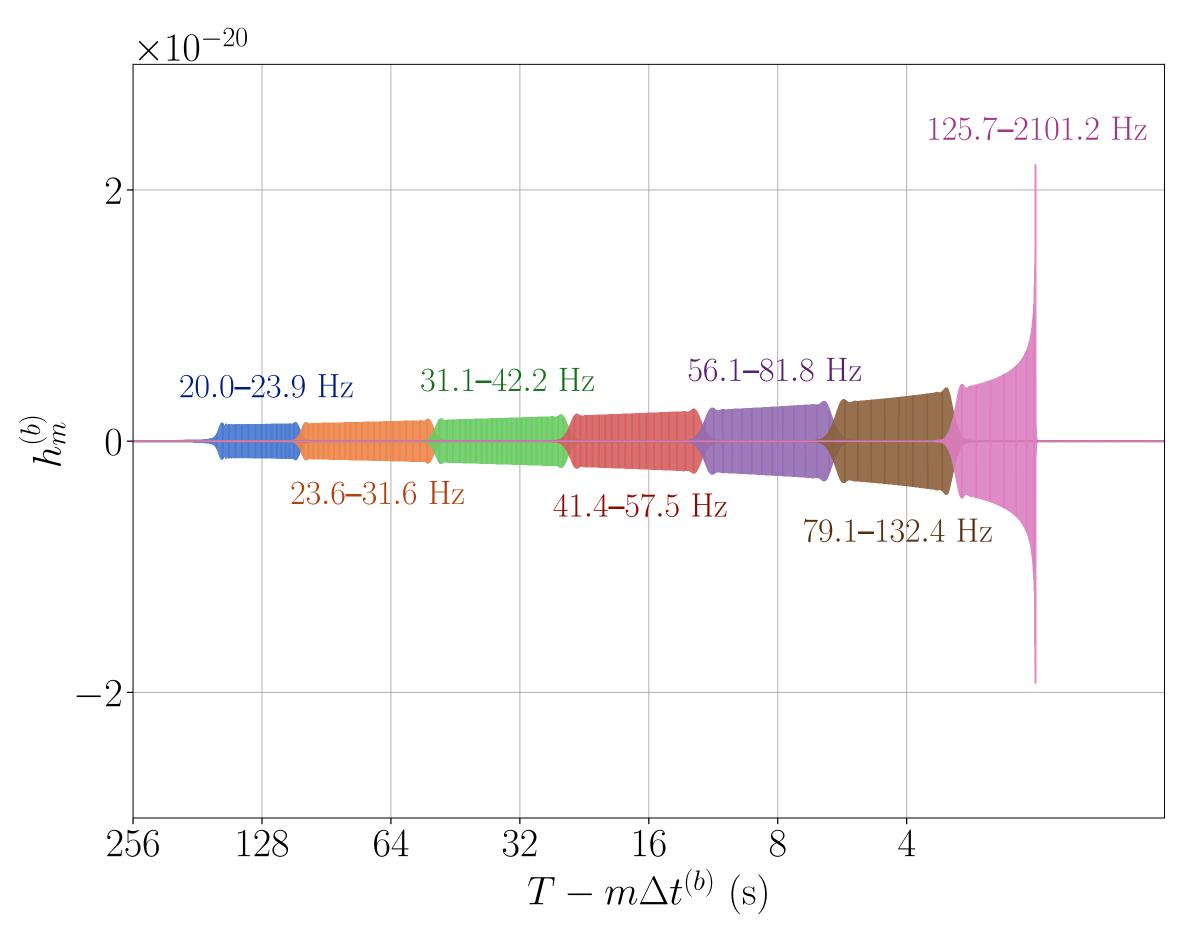
- Long-lived and frequent signals
- XG detectors will be more sensitive → more glitches
 - No idea how bad they will be!
 - Glitches already occurred >1 per minute in O4
 - Detected signals required glitch mitigation
- Glitch characterization & subtraction must be automated
 - Manual interventions (e.g., GW170817) not scalable
- Use non-Gaussian proof PE methods
 - Score-based likelihood (<u>Legin et al 2024</u>)
 - Simulation-based inference (<u>Dax et al 2024</u>, <u>Bhardwaj et al 2023</u>, Marx .)

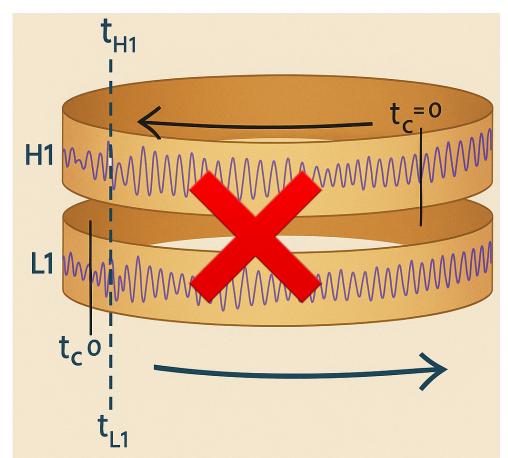


Nirban Bose, Archana Pai, Koustav Chandra et al (2020)









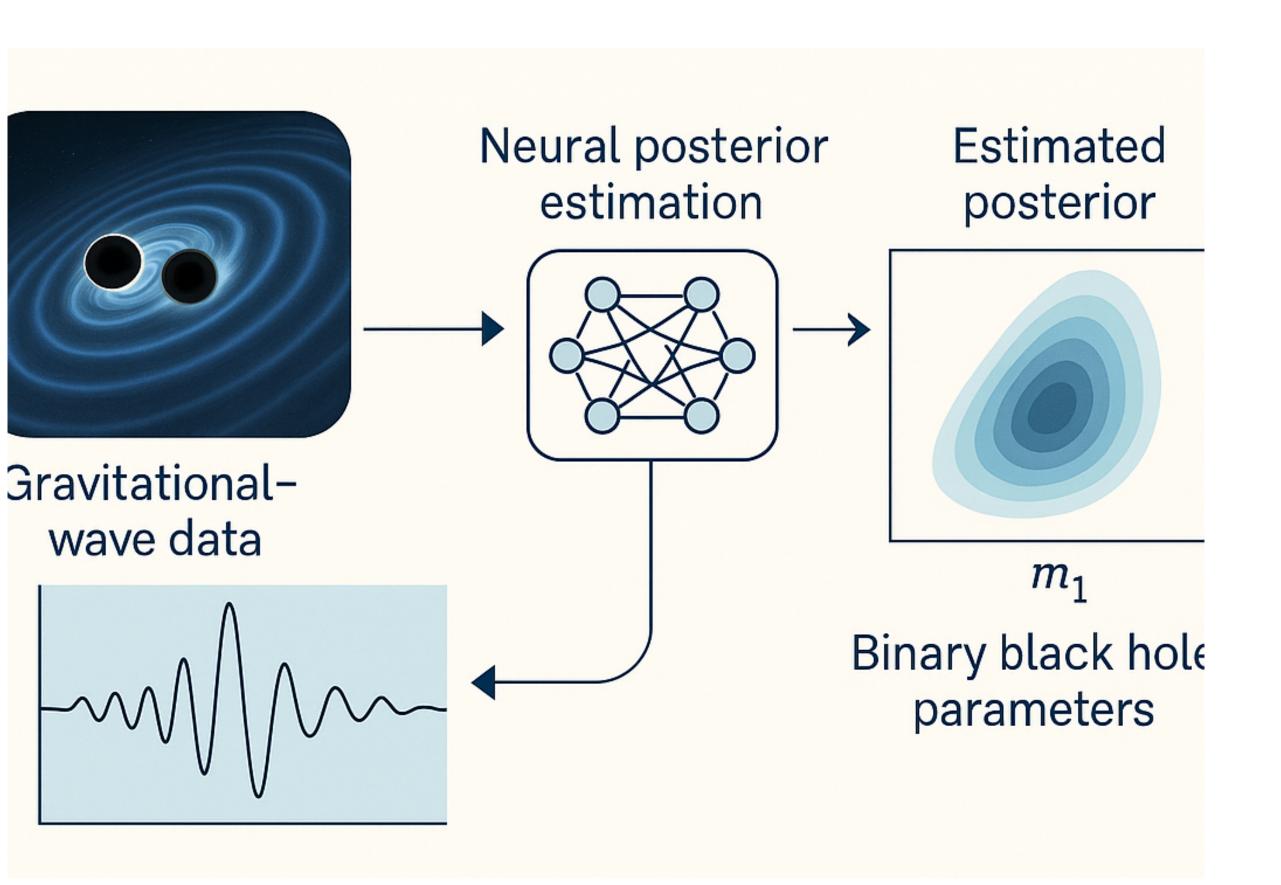
Sheer number and length of signals

Matched-Filter Searches:

- Longer Signals → Larger Banks → Size Reduction via SVD (already used by GstLAL)
- Overlapping Signals → not an issue as MF is a linear filter
- Accelerating the Search
 - Multi-banding: Slashes computation (already used by MBTA)
 - GPU Acceleration (Nitz 2015, Chu et al 2020, Huang et al 2024)
- Glitches: Use (semi-)coherent methods instead of coincident + Better vetoes (?)
- Background Estimation Limits: Time-slide techniques will break down in signal-rich data

Unmodelled Searches: can detect overlapping signals, but identify them as a single event rather than separating them (Relton et al 2022)

Sheer number and length of signals

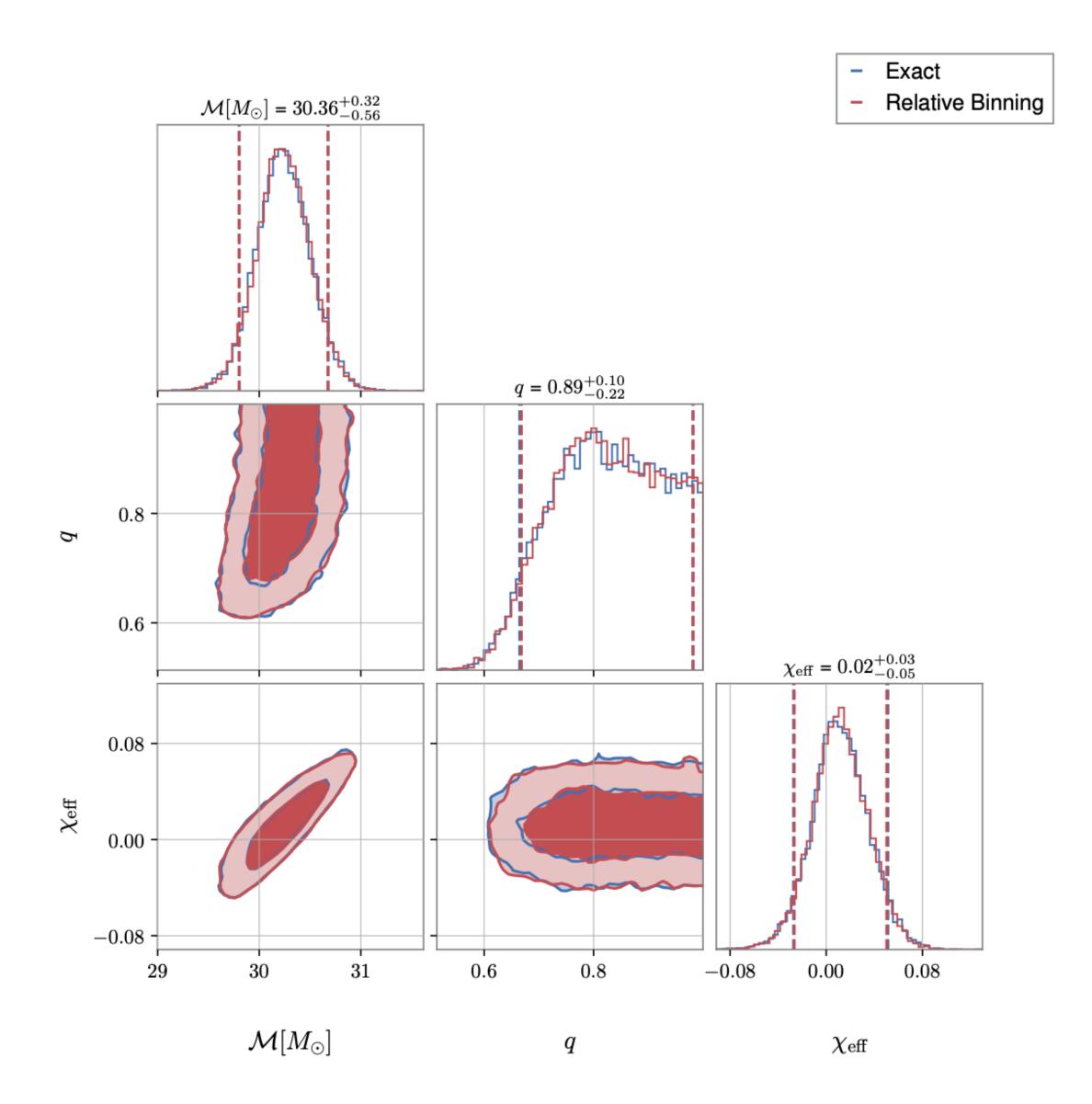


Parameter Estimation

- Simulation-based inference
 - Neural networks / amortization [<u>Dax et al 2024</u>, <u>Bhardwaj et al 2023</u>]
- Less of a need for faster waveforms



Sheer number and length of signals



Parameter Estimation

Simulation-based inference

 Neural networks / amortization [<u>Dax et al 2024</u>, <u>Bhardwaj et al 2023</u>]

Faster likelihoods

- faster waveforms → probably frequency domain surrogate models
- multi-banding, heterodyning, ROQ, etc.

Efficient sampling

- Hamiltonian MCMC (DeepMCMC [Perret & Porter 2025])
- GAMES [Nitz 2024]
- ML-enhanced samplers (nessai [Williams et al 2023])

- Fixed sky location assumption works for 2G signals
- BNS inspirals in XG detectors can last >1000s →
 Slow amplitude & phase modulation as IFO
 response time-dependent

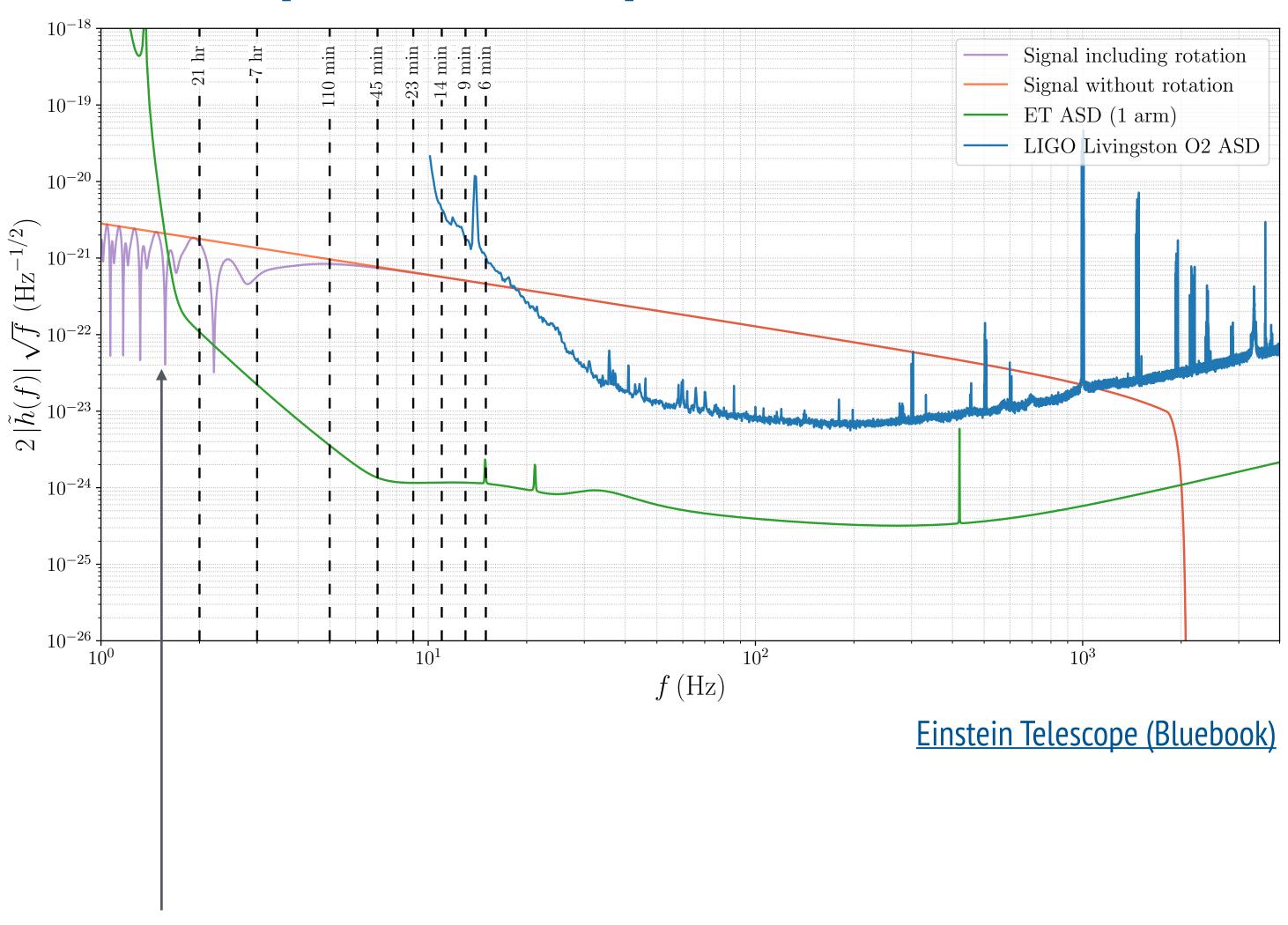
Searches

- Solution: Use template banks covering sky
 location + intrinsic parameters
- No need \rightarrow SNR loss negligible (χ_r^2 slightly large) for most detections [Meacher et al 2015]

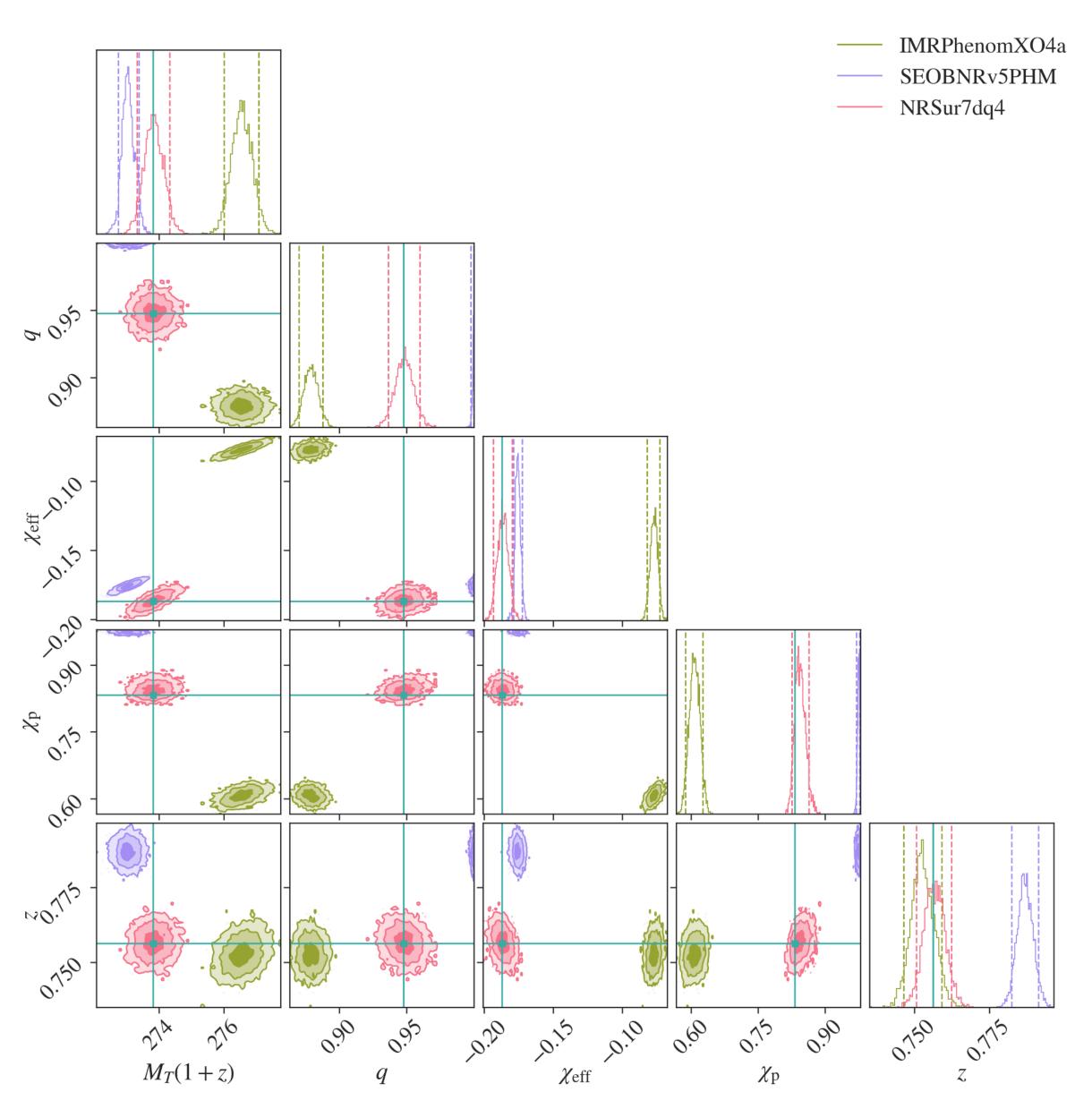
PE

• Frequency-dependent antenna-response codes exists [Biwer et al (2018), Baral et al (2025)]

Comparison of Amplitude modulation



GW370817 (with and without accounting for Earth's rotation)



Example 2 List of my headaches (Should be yours too)

- Noise Modelling & Glitches Now with bonus jump scares!
- Too Many Waveforms, PE Methods, Searches "Lost in parameter space 🍑 🌀 — Send help (and maybe a flowchart)."
- Waveform Systematics "Still haunting us like a ghost ex 😐 🖳 — Sorry Rossella, don't 🦠 me!"
- Calibration Uncertainty

"What even is the true strain? 🍑 📉 — If I don't know, it's definitely a problem!" & we want no human intervention (even if to do just astrophysics)

Computational Resources

"WE NEED MORE CLUSTER!!! — \gg — and by that, I mean 🐞 🐞





What's your headache?

